

**CEOI 5<sup>th</sup> and 6<sup>th</sup> Open Calls  
Final Review**

***High-Level System Integration of  
UK Receiver Technology for  
STEAM-R & MWS***

***RAL Space, STAR-Dundee Ltd***

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BMA House, Tavistock Square, London

20 March 2013

# Outline

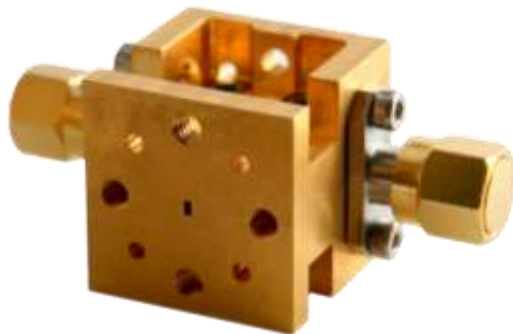
- Project Background
- Aims of Current Project
- Review of UK Support to STEAM-R
  - Introduction to STEAM-R
  - Previous development work related to current project
- Technical Report on Current Project
- Achievements v Goals, Issues & Problems Encountered, Positioning Achieved
- Roadmap / Future Activities

# Project Background

- UK invited to support Swedish-led development of STEAM-R (2003)
  - Follow-on to ODIN mission
  - Millimetre-wave payload of PREMIER mission (EE7)
- Core UK support funded through CEOI (from Phase 1)
  - Science support (Remote Sensing Group, RAL Space)
  - Critical hardware development (MMT Group, RAL Space; Astrium Ltd; STAR-Dundee Ltd)
- Related activities
  - Deployment of MARSCHALS in PREMIER-Ex & ESSENCE science campaigns (ESA/NCEO/STFC)
  - Development of wideband spectrometer (ESA/STAR-Dundee Ltd/Astrium Ltd)
  - MARSCHALS Upgrade (STFC/UKSA)
- EE7 User Consultation Meeting (this month)
  - Awaiting official decision (expected in May)
  - In event of PREMIER de-selection by PBEO, STEAM-R likely to become Swedish national mission

# Aims of Current Project

- System-level integration & characterisation of UK receiver technology
  - Development of Wideband Spectrometer v2 (WBS II) [STAR-Dundee Ltd]
  - Breadboarding & characterisation of total-power radiometer comprising 340 GHz sideband-separating receiver with high-resolution back-end (SHIRM + WBS II) [RAL Space]
  - Preliminary assessment of integration of SHIRM + WBS II receiver into MARSCHALS [RAL Space]
- Risk mitigation in preparation for full upgrade of MARSCHALS



*Sub-Harmonic  
Image-Rejection Mixer (SHIRM)*

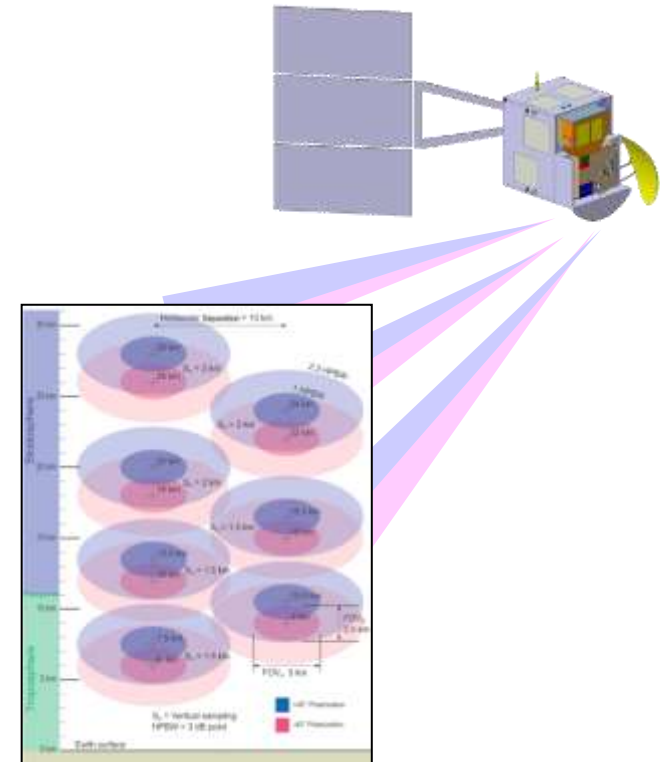


*Wideband Spectrometer II*



*MARSCHALS on Geophysica*

- Purpose of PREMIER
  - Provide better understanding of link between composition of UTLS and climate change
  - STEAM-R provides observations in presence of cloud (H<sub>2</sub>O, O<sub>3</sub>, N<sub>2</sub>O, CO, HCN, ...)
  
- Spectral coverage: 315 – 355 GHz
  
- Heterodyne receiver array providing 14 simultaneous views through the UTLS
  - DSB (baseline)
  - 2SB (high priority option for lowest views)
  
- High resolution spectrometer back-end
  - 168 GHz instantaneous BW (min.)
  - Incumbent technology is Digital Autocorrelation Spectrometer (DAS) from Omnisys Instruments

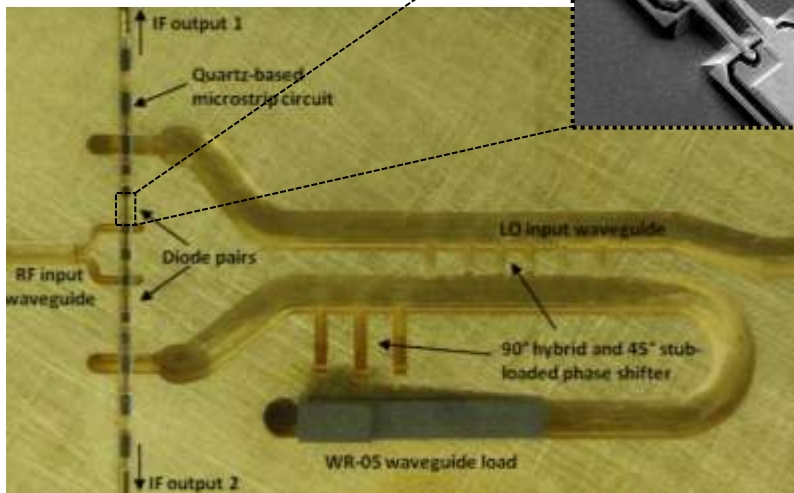


*STEAM-R viewing geometry: 14 simultaneous views in UTLS region. Receivers split into two 7-element arrays observing in orthogonal polarisations*

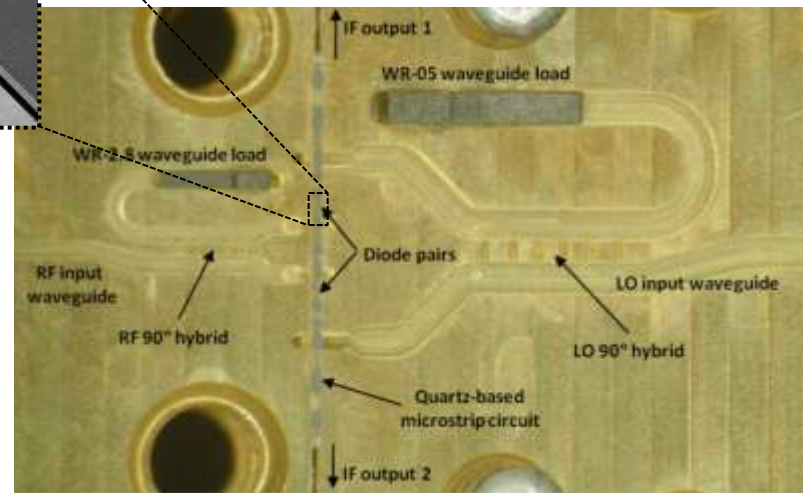
## *Development Review: Sub-Harmonic Image-Rejection Mixer*

- SHIRM Optimised performance
  - Sideband rejection: 15 dB min. (>20 dB nom.), IF BW = 2-14 GHz
  - SSB receiver noise temperature: ~3000 K
- Devices employ planar Schottky diode technology from RAL Space

*Phasing Topology 1*



*Phasing Topology 2*





## *Development Review: Sub-Harmonic Image-Rejection Mixer*

- SHIRM receiver deployed at High Altitude Research Station Jungfrauoch in Feb 2012
  - Collaboration with Institute of Applied Physics, University of Bern
  - IAP contracted by Omnisys to undertake atmospheric measurements of their DSB receiver for STEAM-R



*Sphinx Observatory: 3480m a.m.s.l*

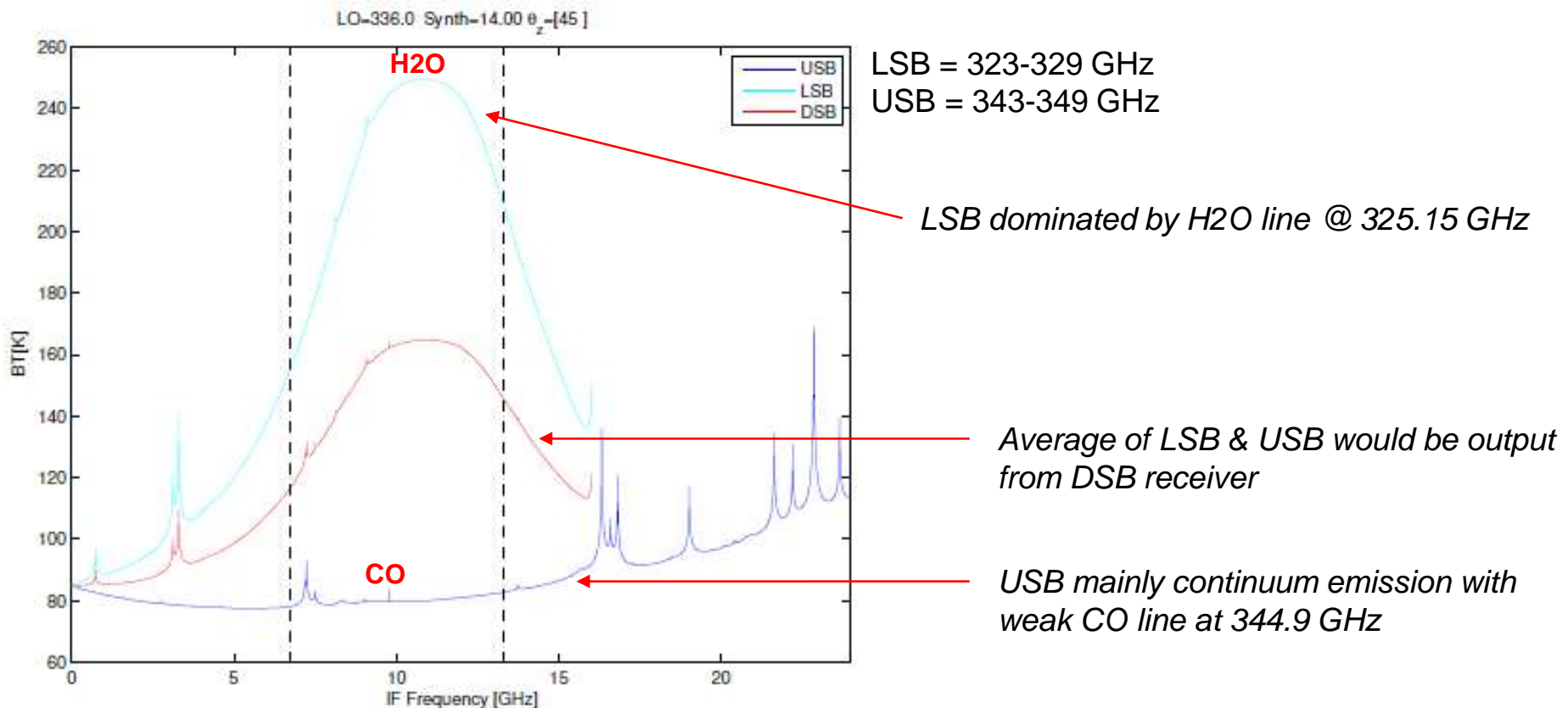


*Test-bed comprising:*

- UK SHIRM receiver or Swedish DSB receiver
- Omnisys digital autocorrelation spectrometer

## Development Review: Sub-Harmonic Image-Rejection Mixer

- Demonstrating SHIRM performance through atmospheric observation

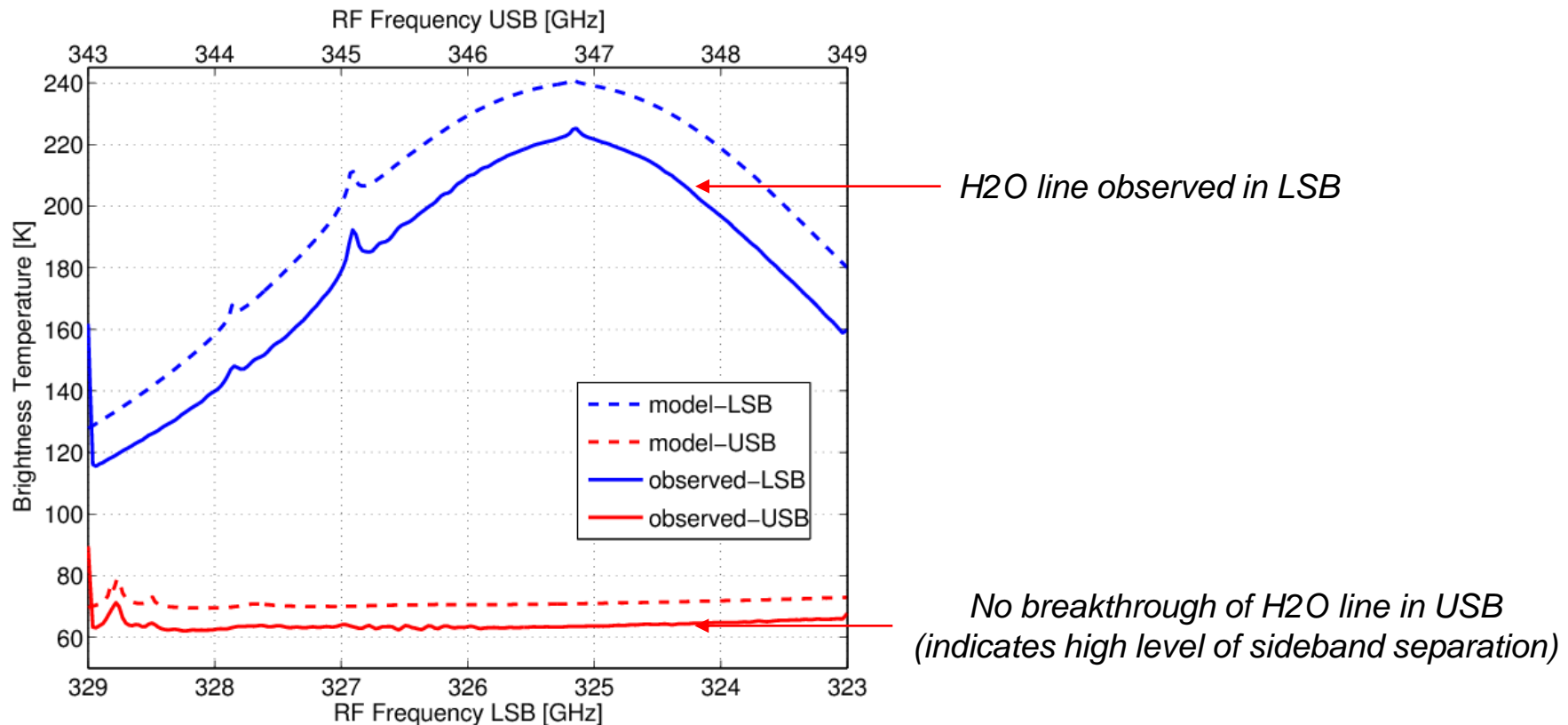


Simple atmospheric simulation at mid-latitudes



## Development Review: Sub-Harmonic Image-Rejection Mixer

- Successful demonstration of sideband separation receiver
  - Sideband separation of 24 dB inferred from atmospheric data



## *Development Review: Wideband Spectrometer*

- Prototype Wideband Spectrometer (WBS) developed by STAR-Dundee Ltd / Astrium Ltd (ESA Wideband Spectrometer Study)
  - Digital FFT-based, real sampling @ 3 Gsamples/s max
  - Bandwidth 1 GHz, Resolution 1 MHz
- Additional work funded in CEOI 4<sup>th</sup> call project
  - Characterisation with mm-wave front-end (MARSCHALS)
  - Options for increasing bandwidth and route to space implementation
  - Recommendations for WBS v2



*Prototype WBS*



*WBS Characterisation at RAL*

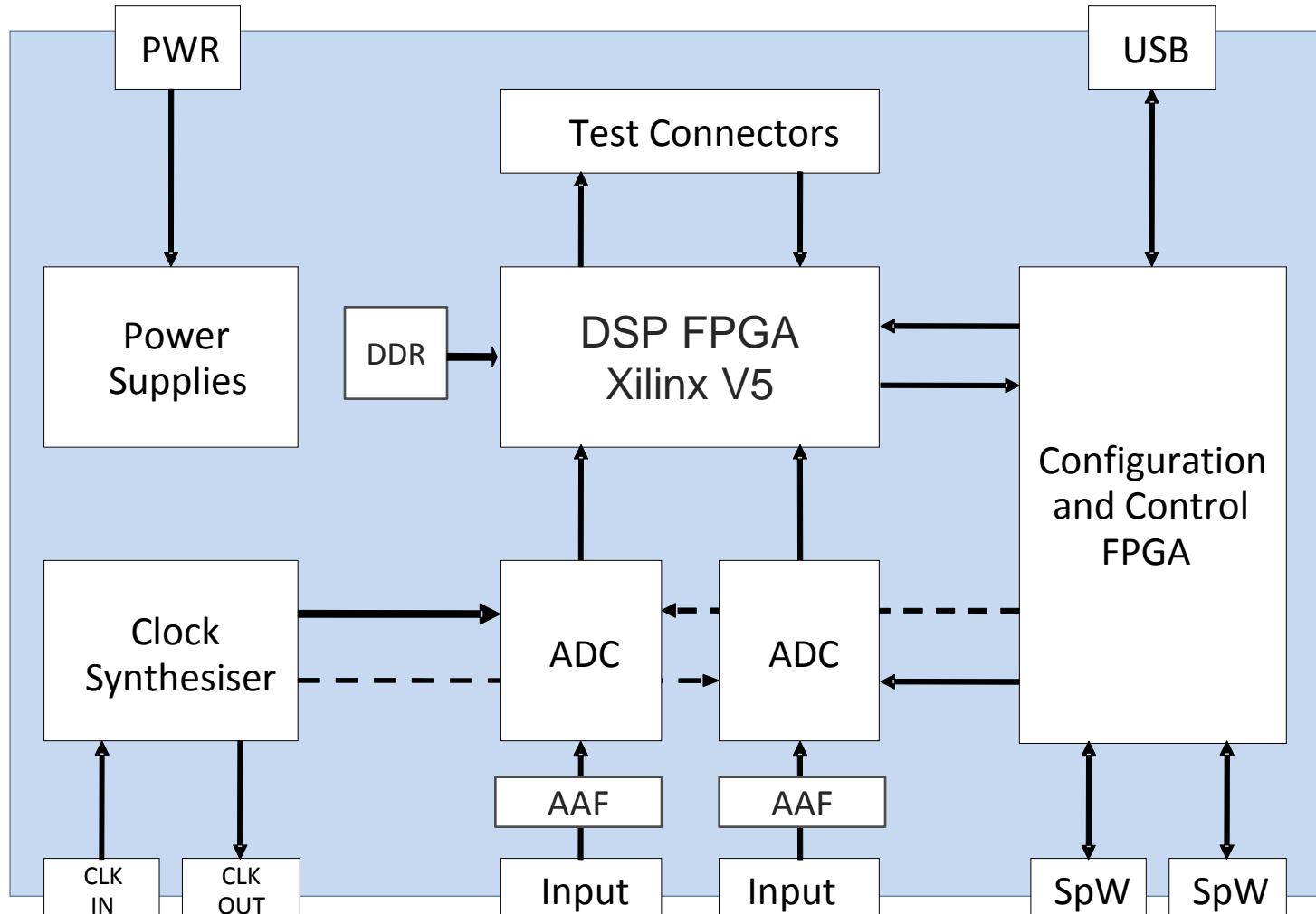
## *Project Status*

- WBS II Development
  - WBS II undergoing final testing at STAR-Dundee Ltd
  - Two units being assembled for delivery to RAL
- SHIRM-WBS II Instrument Breadboard
  - Instrument integration at an advanced stage
  - Test of sub-systems partially completed
  - Awaiting WBS II units
- MARSCHALS Upgrade Study
  - SHIRM front-end integration study complete
  - Awaiting final WBS II performance / budgets as inputs to backend integration study
- Target completion date: End June 2013

## *Wideband Spectrometer II Development*

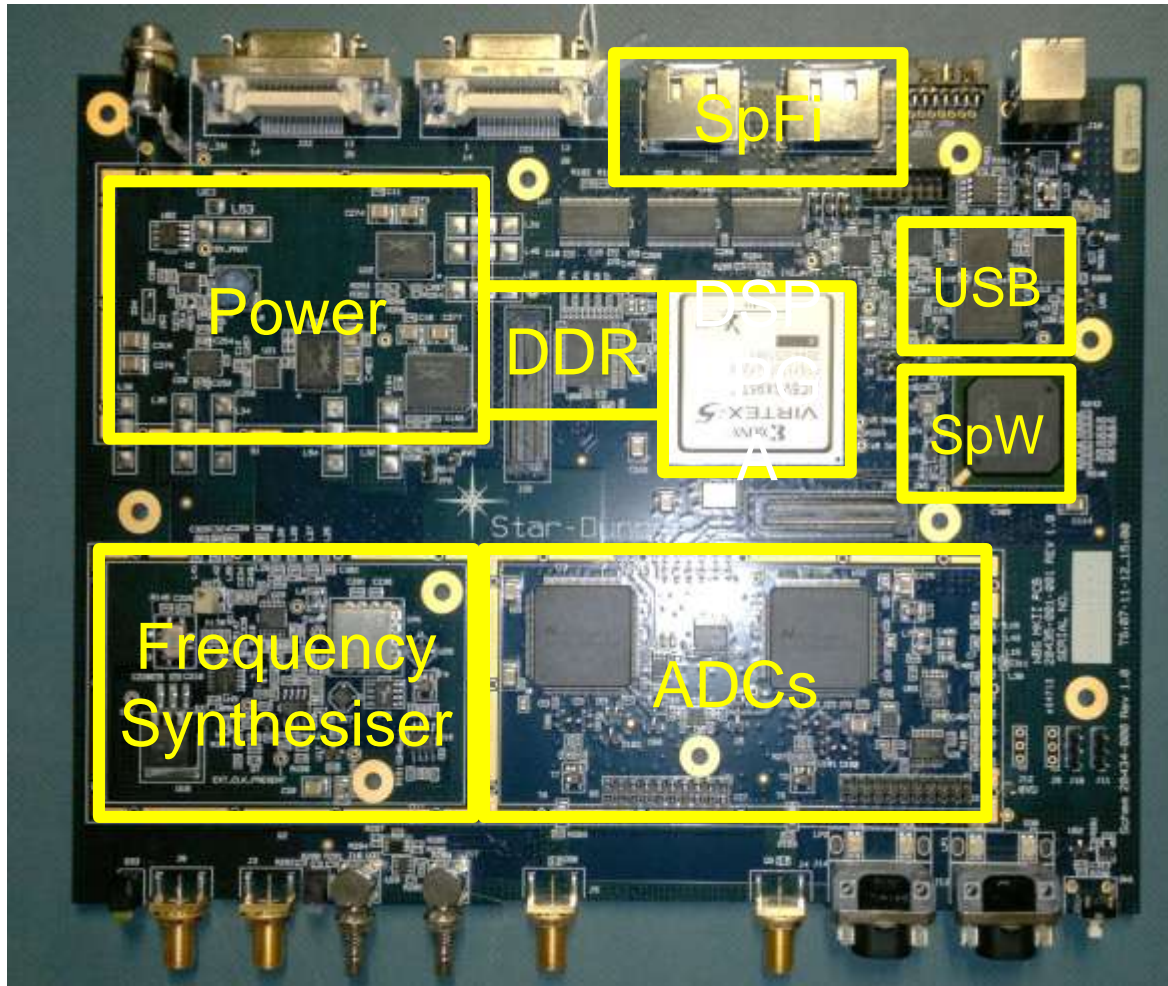
- Two ADCs sampling at 3 Gsamples/s
  - I & Q sampling
  - Resulting signal bandwidth > 2 GHz
- Custom Fast Fourier Transform (FFT) chip design
  - Windowing
  - 2048 point complex FFT at 3 Gsamples/s
    - ~ 1.5 MHz resolution
  - Power detection and accumulation
  - Zero dead-time between data acquisitions
- Hardware triggering interface
  - Start/stop individual acquisitions
  - Start/stop groups of acquisitions of predetermined duration

## Wideband Spectrometer II Development





## *Wideband Spectrometer II Development*



## *Wideband Spectrometer II Development*



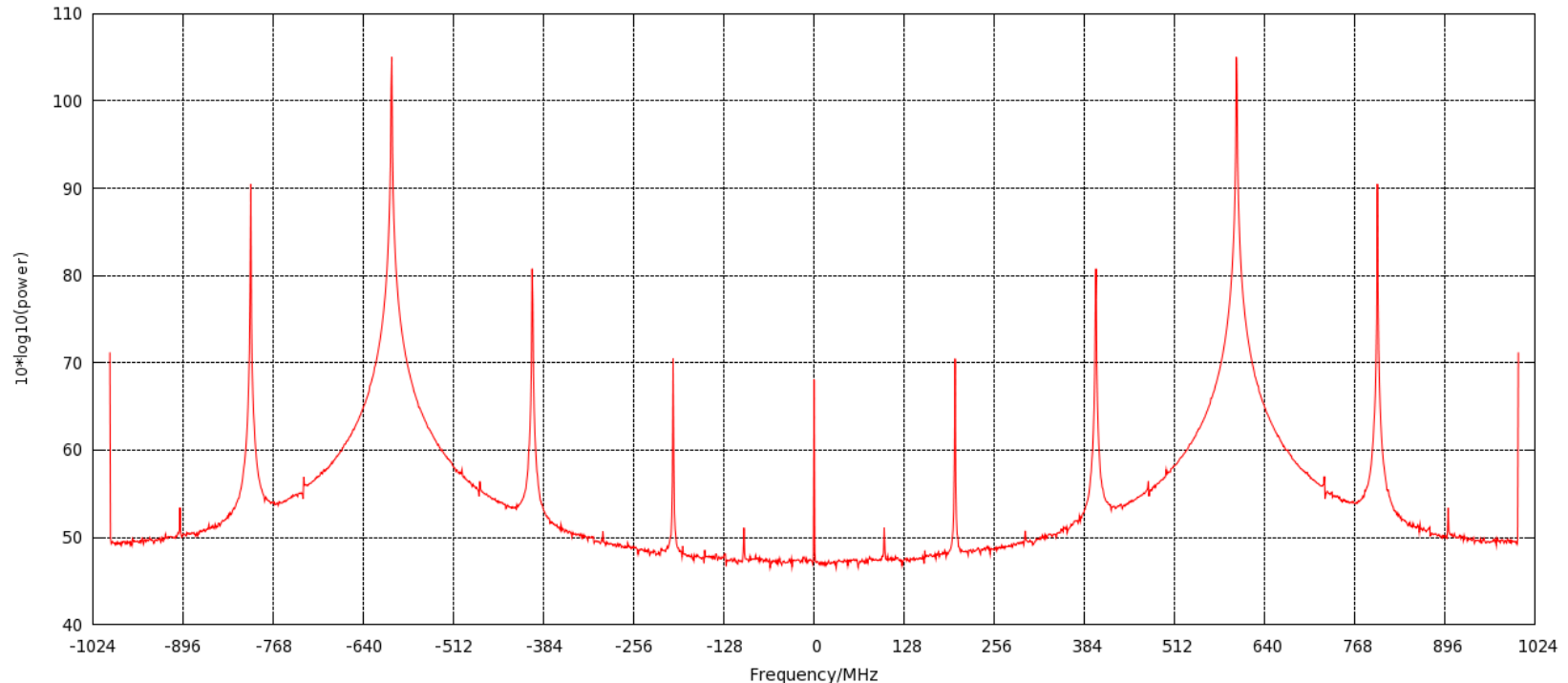
Dimensions  
L= 165 mm  
W= 220 mm  
H= 30 mm



## Wideband Spectrometer II Development

### Preliminary Test Results: FFT running in hardware

Vaunix 600MHz at 0dBm WBS II 2000MHz, n=10000 (no window)



Signal fed into I input, Q input = 0

Averaged for 10,000 spectra

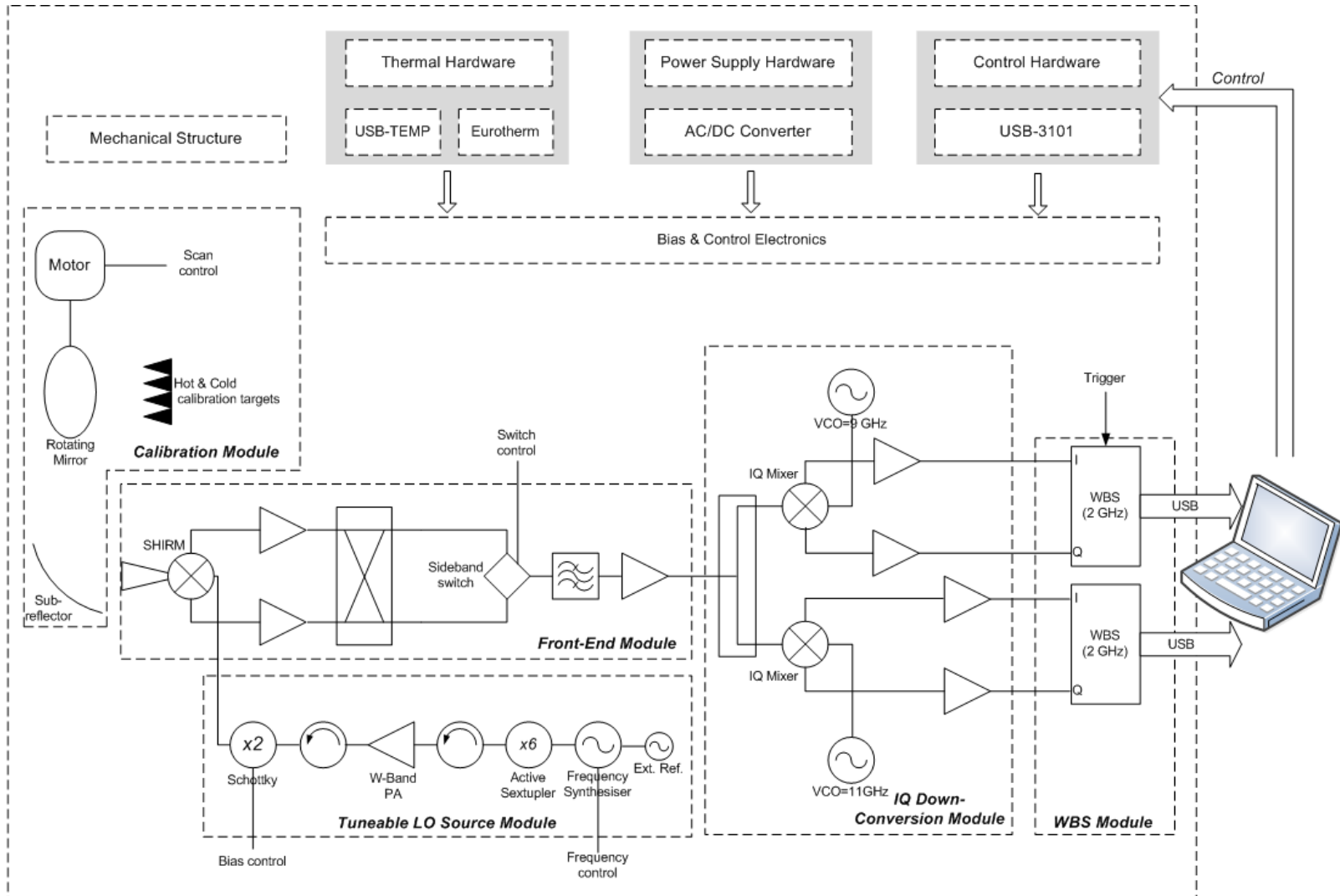
Primary signal at 600 MHz (spurs at ~200, 400 & 800)

Clock breakthrough is low (@ 100 MHz)

## *SHIRM-WBS II Breadboard Development*

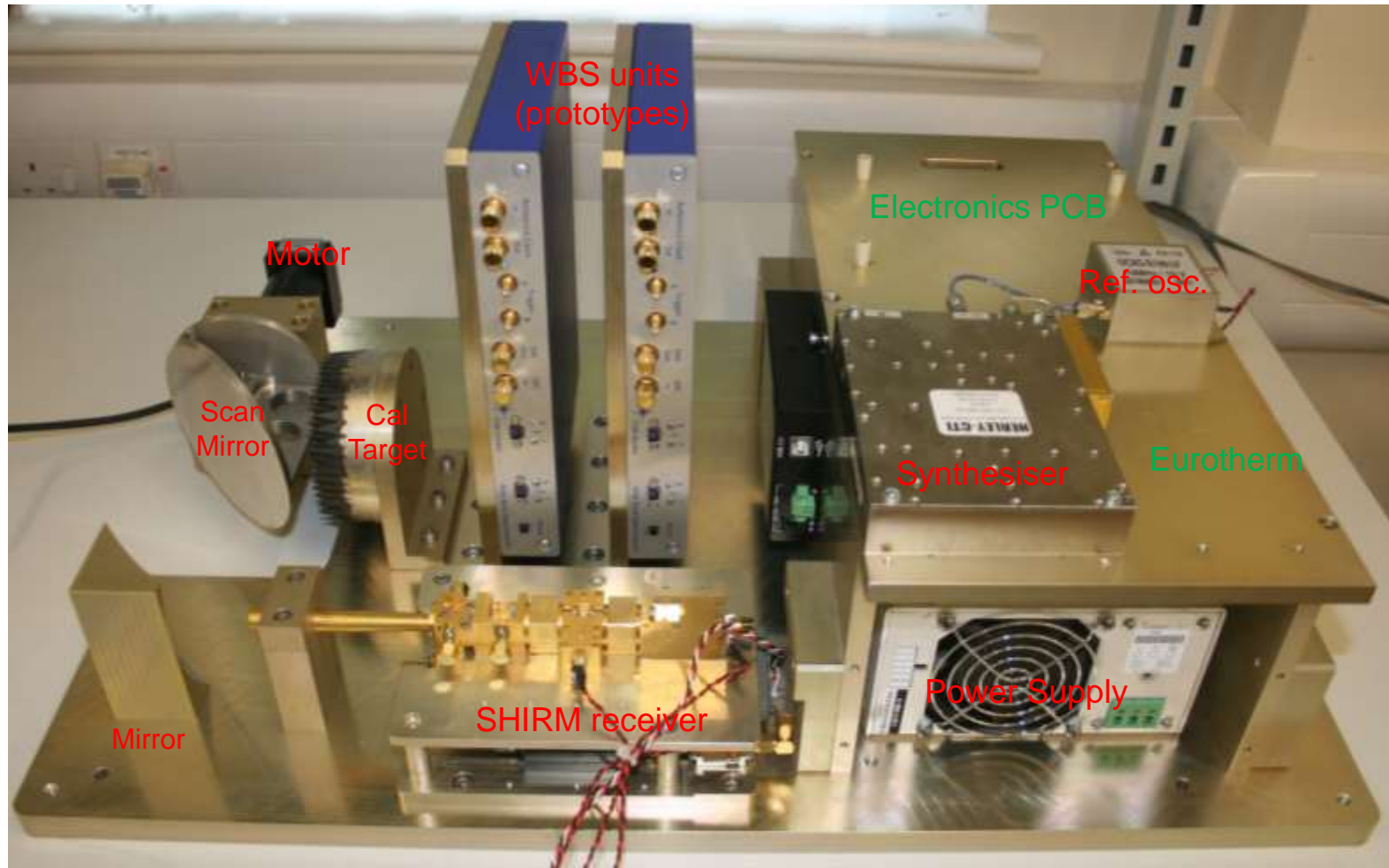
- Development of total-power radiometer comprising
  - 340 GHz sideband-separating receiver
  - 2x WBS II units providing 4 GHz bandwidth
- Characterisation of receiver
  - Noise temperature, sideband separation, stability, spectral response
  - Knowledge & understanding of system-induced artefacts
  - Risk mitigation in preparation for MARSCHALS upgrade
- Potential for future deployment at ground-based observatory
  - Return to Jungfraujoch

## SHIRM-WBS II Breadboard Development

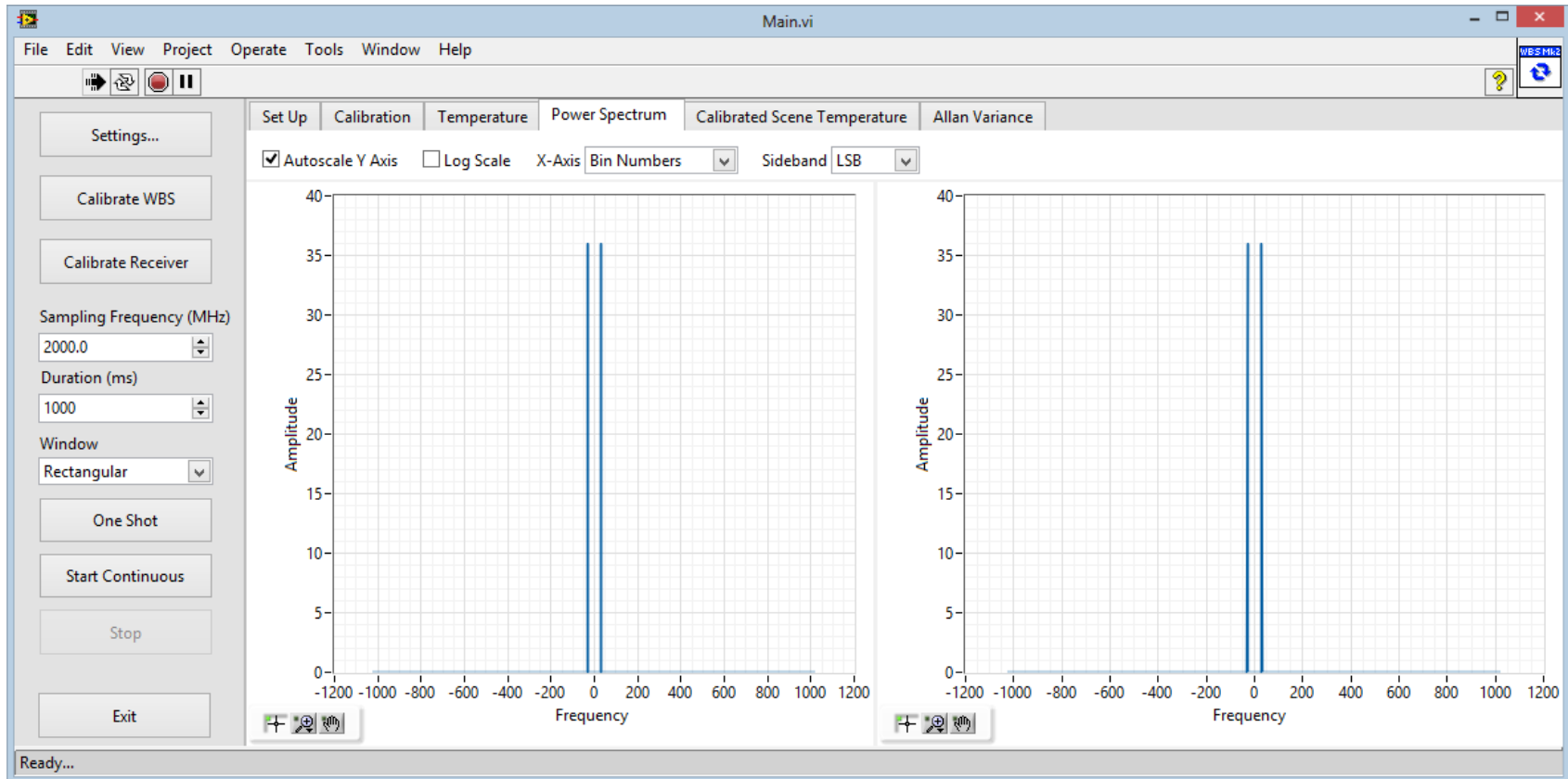




## *SHIRM-WBS II Breadboard Development*



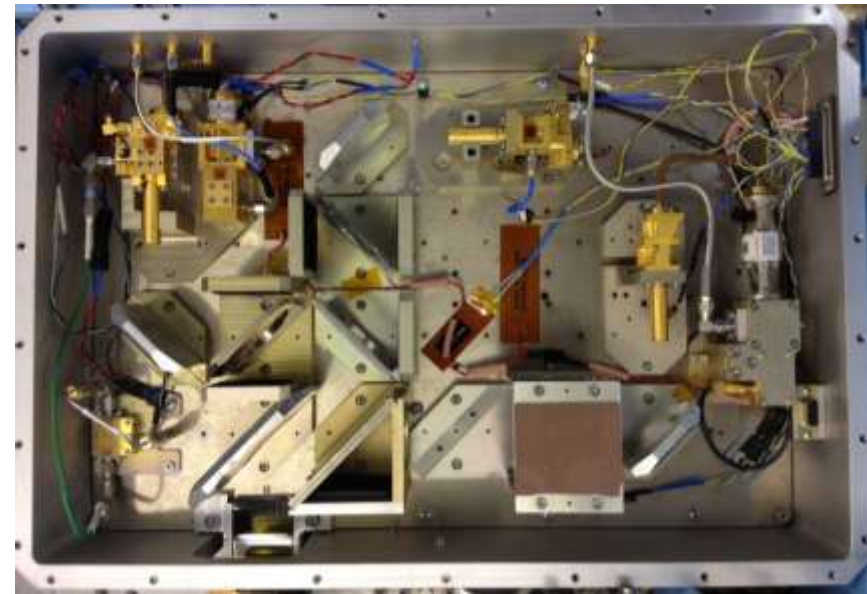
## *SHIRM-WBS II Breadboard Development*



Screenshot of software showing two spectra from two WBS units.  
 Tabs indicate different views of the data.  
 Configuration and calibration buttons down on the left hand side.

## *MARSCHALS Upgrade Study*

- Preliminary assessment for integration of
  - 340 GHz sideband-separating receiver
  - 6x WBS II units providing 12 GHz bandwidth
- Front-End assessment conclusions
  - Existing receivers in MARSCHALS
    - Band B (294-305.5 GHz)
    - Band C (316.5-325.5 GHz)
    - Band D (342.2-348.8 GHz)
  - Replace Band D with SHIRM receiver
  - New QO diplexer required



*MARSCHALS QO Layout*

## *MARSCHALS Upgrade Study*

- Preliminary assessment for integration of
  - 340 GHz sideband-separating receiver
  - 6x WBS II units providing 12 GHz bandwidth
- Back-End assessment on-going
  - Expect to retain original filterbank spectrometers (12 GHz @ 200 MHz res.)
    - well characterised and flight-proven
  - Must find additional space for WBS II units (space identified on MARSCHALS)



*Identified space for  
high-res spectrometer*

# Achievements against Goals

- Project not completed, however significant hardware development undertaken in short timeframe
- Development of WBS II
  - Xilinx FFT core too slow
  - Custom FFT FPGA core had to be designed
- Development of SHIRM-WBS II breadboard instrument
  - System design of complete instrument
    - mm-wave receiver
    - Mechanical design
    - Power supply
    - Thermal control
    - Bias/control electronics
    - Instrument control software
    - Scan control
    - Calibration



# Issues and problems encountered

- WBS II
  - New FFT core design was challenging – difficult to fit logic into FPGA
  - Synchronisation of data from ADCs was problematic
- SHIRM-WBS Instrument Breadboard
  - No major issues yet (although development has taken longer than expected)
  - Instrument characterisation has not started yet – issues to be expected

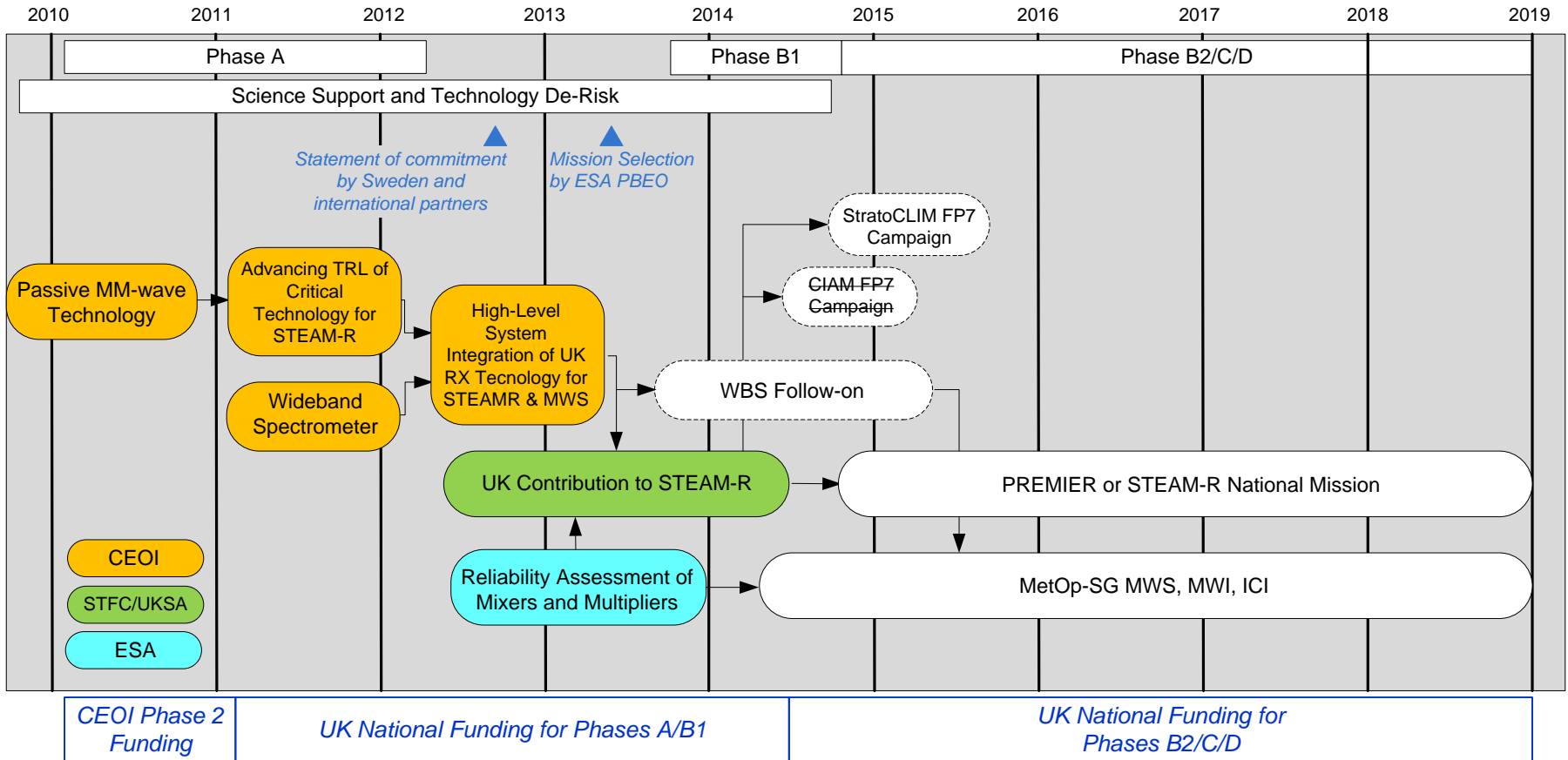
# Positioning Achieved

- Leverage achieved
  - STFC/UKSA funding for UK support to STEAM-R through FYs 12/13, 13/14: £500,000 for upgrade of MARSCHALS
- Collaborations forged
  - Work performed by RAL Space, Astrium Ltd and STAR-Dundee Ltd
  - Support from IAP, University of Bern
  - Support from International Foundation High Altitude Research Stations Jungfrauoch and Gornergrat (HJSJG)

# Other Achievements

- Training and knowledge exchange
  - Sandwich Placement Student at RAL Space: Sean Woodley (University of Bath)
  - Supporting development of SHIRM-WBS II breadboard
  
- UK Capability enhancement
  - Development of digital spectrometer with wide range of applications
    - Passive radiometers
    - Synthetic aperture radar
    - Image processing
    - On-board data processing
  
  - Demonstration of end-to-end performance of sideband-separating receiver with high-resolution backend using UK technology
    - strengthens case for inclusion of SHIRM receiver in STEAM-R

# Roadmap



# Future Activities

- Future steps
  - Await final decision on EE7
  - Engage with Swedish partners in event of PREMIER de-selection
- Continue with current funded projects
  - CEOI 5<sup>th</sup> Call
  - UK Support to STEAM-R (STFC/UKSA)
- WBS follow-on project (dependent on target mission)
  - Increase bandwidth further (higher speed ADCs)
  - Implement IQ imbalance correction
  - Develop flight-representative unit (radiation tolerant FPGA or ASIC)