



# THz Quantum-Cascade Lasers for Heterodyne Techniques

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# Overview



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- *LOCUS* “supra-terahertz” radiometry channels
- THz quantum-cascade lasers
- Waveguide integrated THz QCL systems

# LOCUS – *Linking Observations of Climate, the Upper-atmosphere and Space-weather*

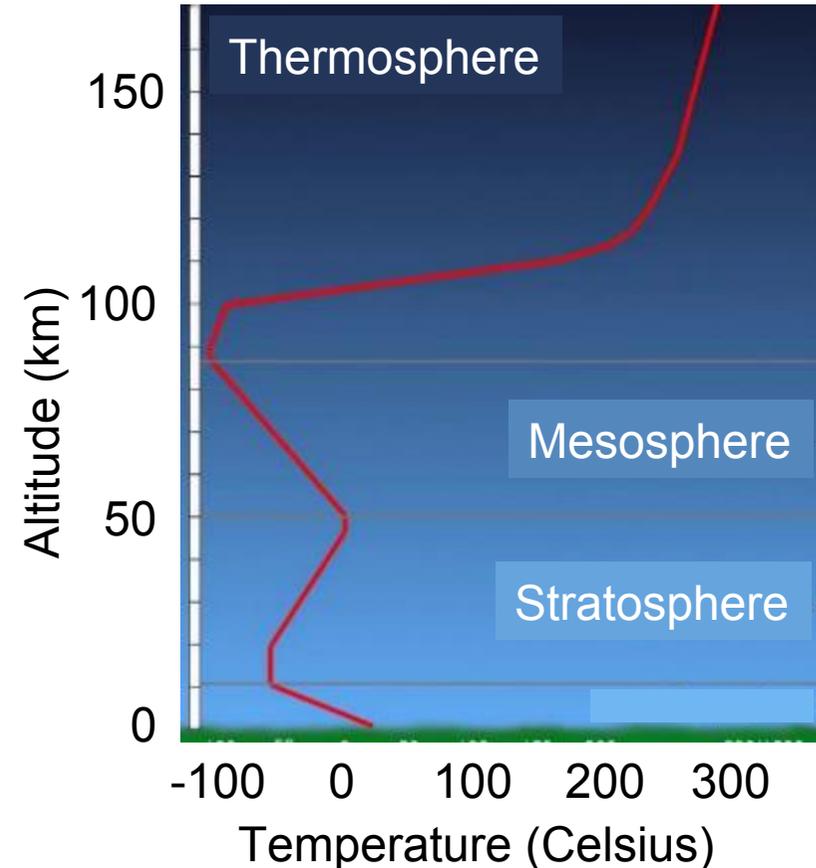
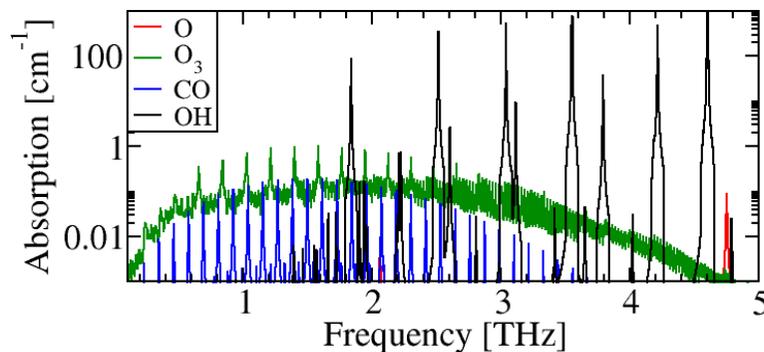


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A breakthrough THz remote sounder

[www.locussatellite.com](http://www.locussatellite.com)

- Compact payload for small satellite
- Measure key species in mesosphere & lower thermosphere
- “Gateway” between Earth atmosphere & near-space
- Increase understanding of natural & anthropogenic effects on climate change

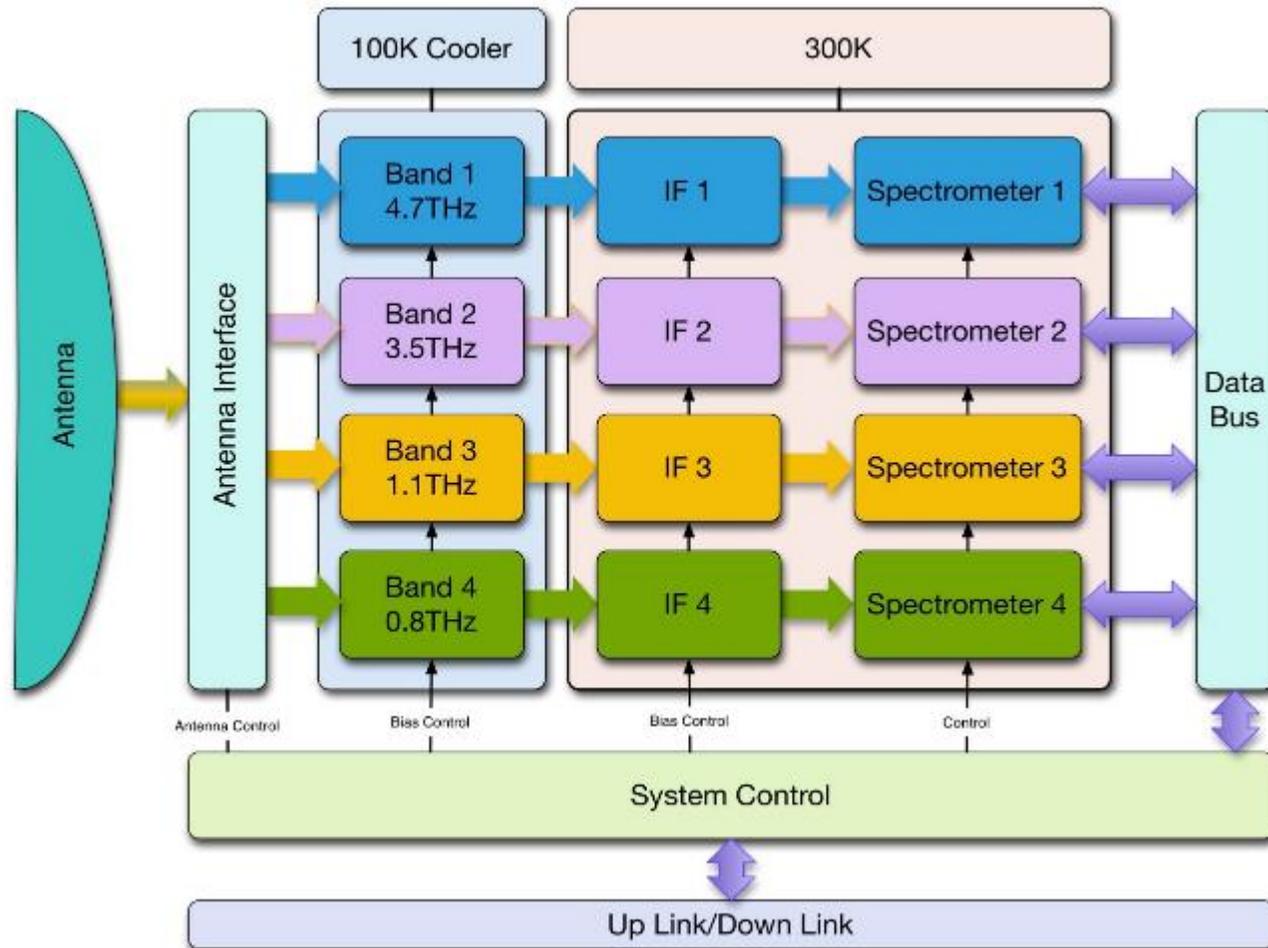


**Integrated, compact and efficient source of THz radiation are needed**

# Radiometry system architecture



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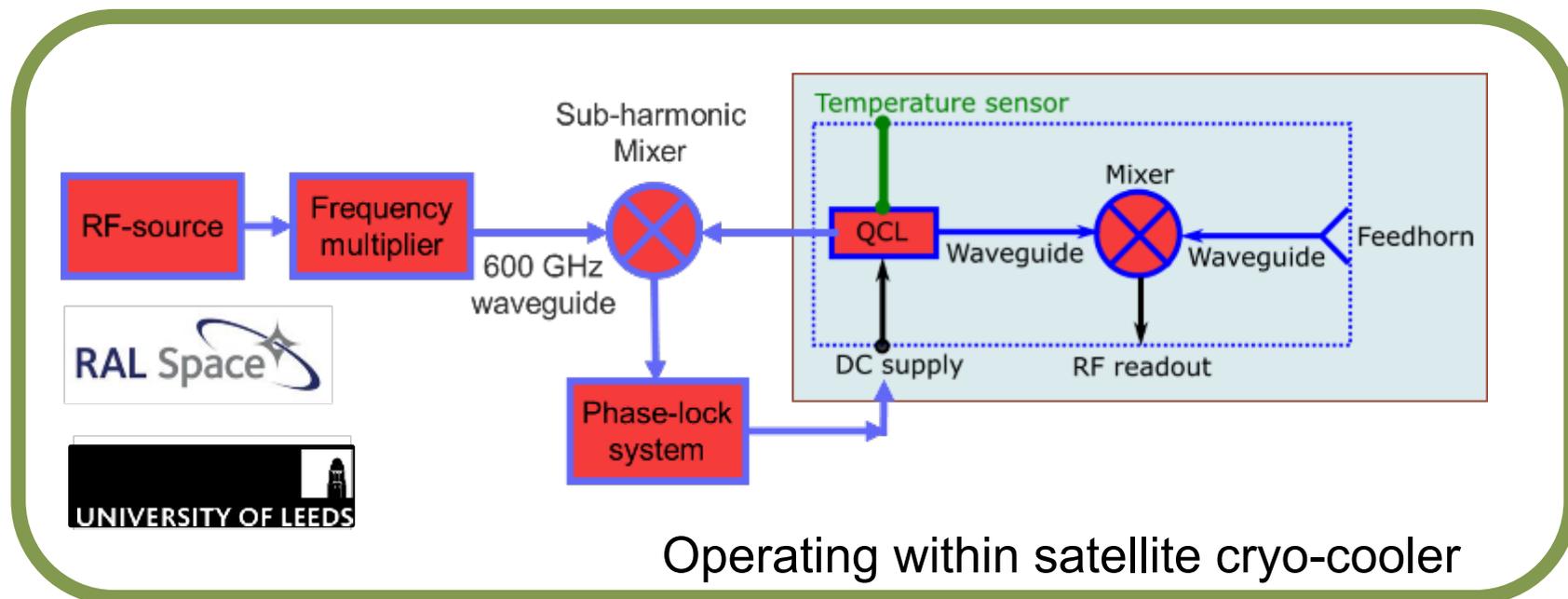
## System schematic

# THz LO requirements



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- 1 mW local-oscillator output power
- Compact, low-mass
- Low input power ( $< 5$  W)
- Fully integrated within satellite-ready cryo-cooler
- 4 GHz tunability

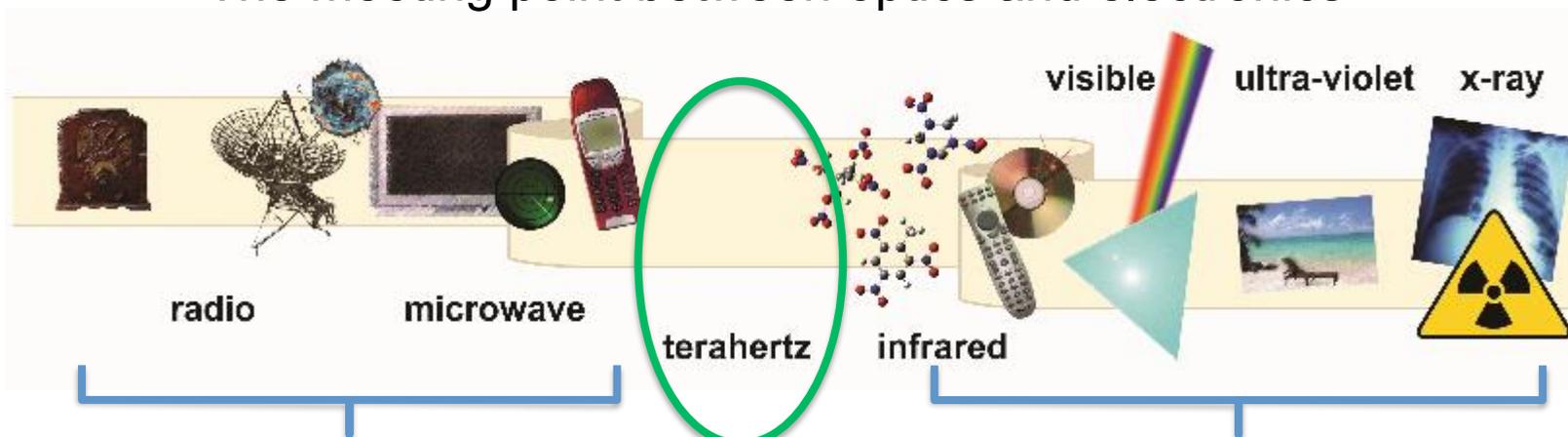


# THz radiation sources



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The meeting point between optics and electronics



## Electronic (classical)

- **oscillators**
- Diodes & harmonics generators
- Limited to low frequencies by transit times

## Optical (quantum)

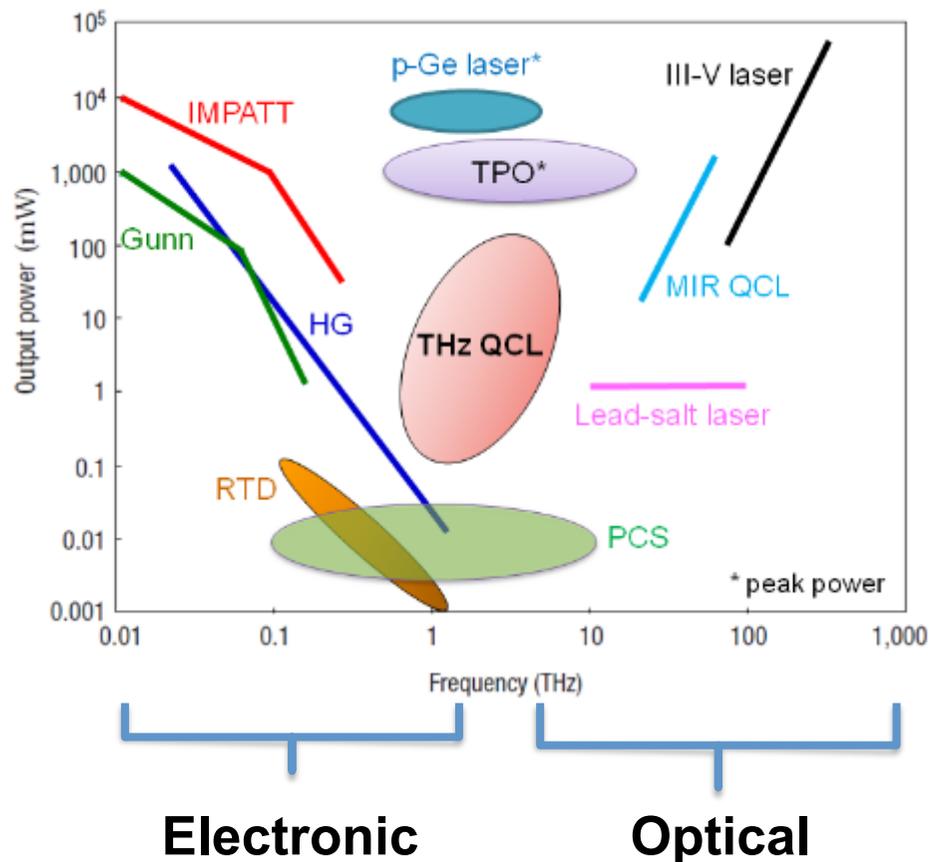
- **oscillators**
- Lasers & atomic transitions
- Limited to high frequencies by energy states in materials

The "THz gap":  
 $f = 2-10 \text{ THz}$

# THz radiation sources



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THz Quantum cascade lasers (QCLs) are the only compact & high-powered coherent THz sources.

- IMPATT – Impact Ionization Avalanche Transit-Time diode
- HG – Harmonic Generation
- RTD – Resonant-Tunnelling Diode
- TPO – THz Parametric Oscillator
- PCS – Photoconductive Switch
- QCL – Quantum Cascade Laser

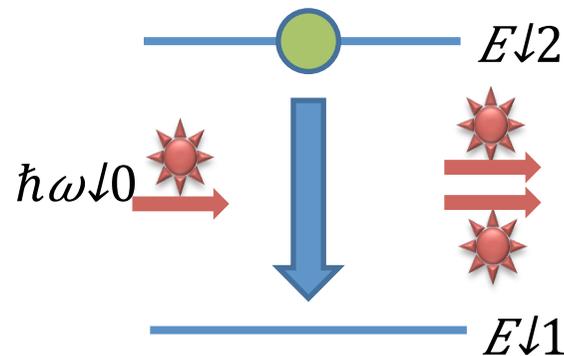
# THz semiconductor lasers



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## Semiconductor device requirements:

- Low enough energy level separation for THz photon emission
- More electrons at high energy than low energy (i.e., a population inversion)



$$h\nu = \hbar\omega_0 = E_2 - E_1$$

$$\begin{aligned} h &= 4.14 \times 10^{-15} \text{ eVs} \\ &= 4.14 \text{ meV / THz} \end{aligned}$$

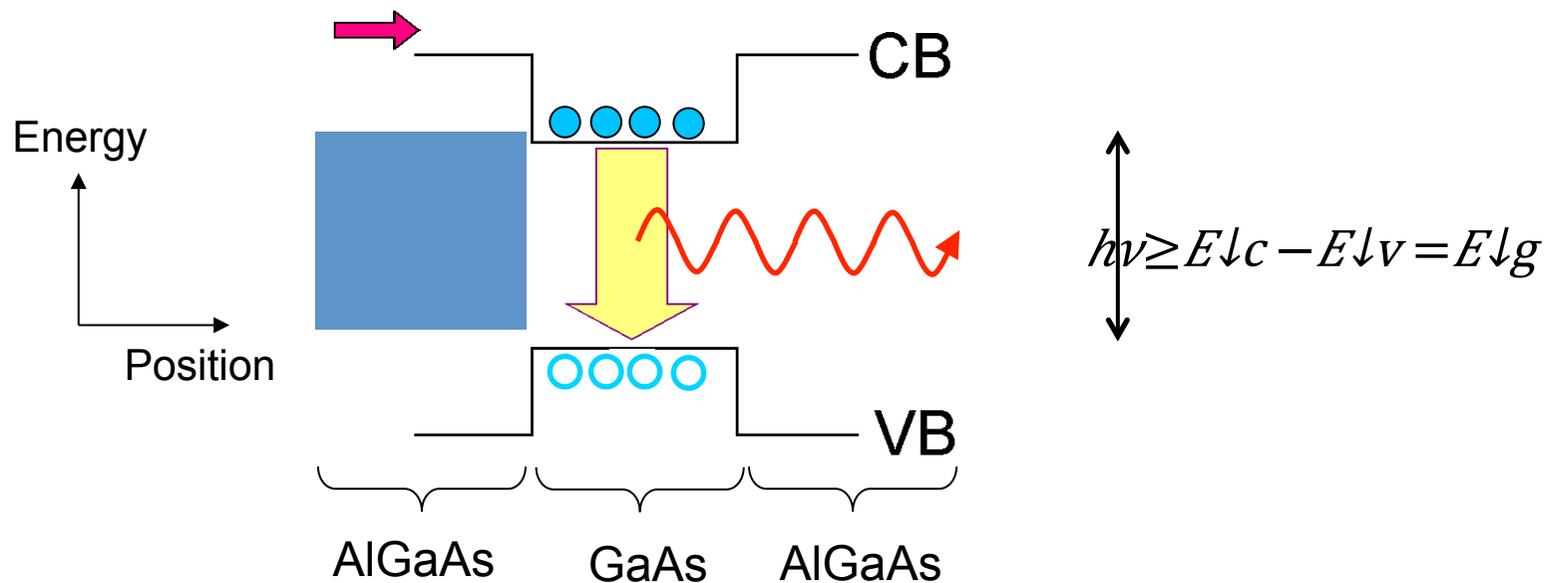
**Absorption occurs:** If particle is initially in low-energy state

**Emission occurs:** If particle is initially in high-energy state

# Conventional laser diodes



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Both electrons and holes involved in (stimulated) emission of photons (bipolar device). One photon per electron (ideally).

**Minimum photon energy = semiconductor bandgap**

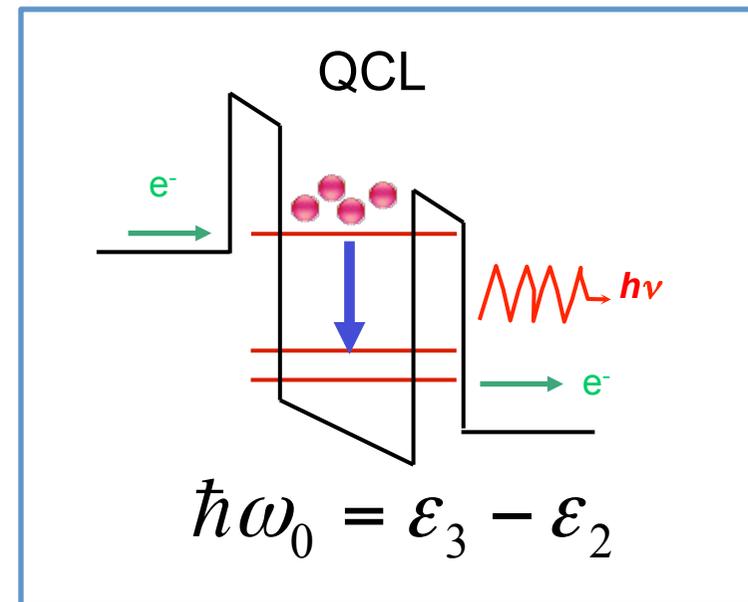
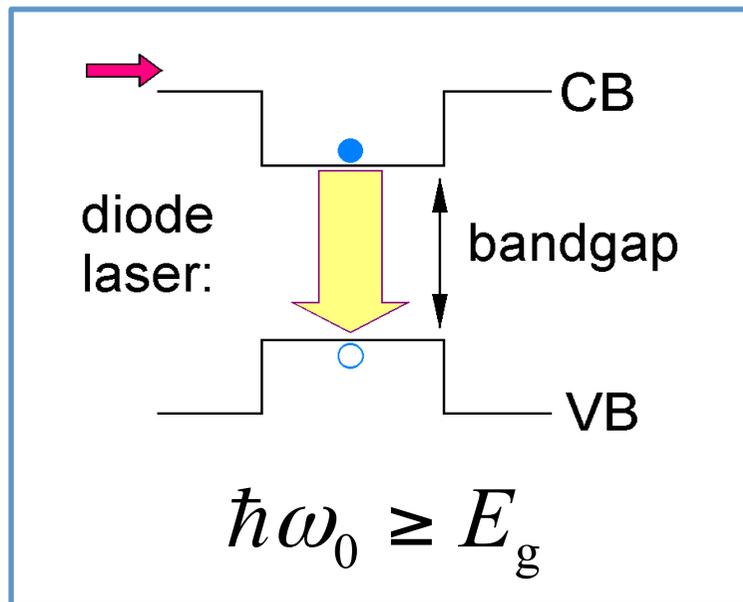
**THz emission impossible with conventional bulk semiconductors**

# Quantum cascade lasers



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**The solution:** use nanoscale layers. Quantum-confinement gives customisable band energies



*Intersubband* transitions in quantum wells

- Long-wavelength emission—Not bandgap limited
- Periodic system: electron 'recycling'

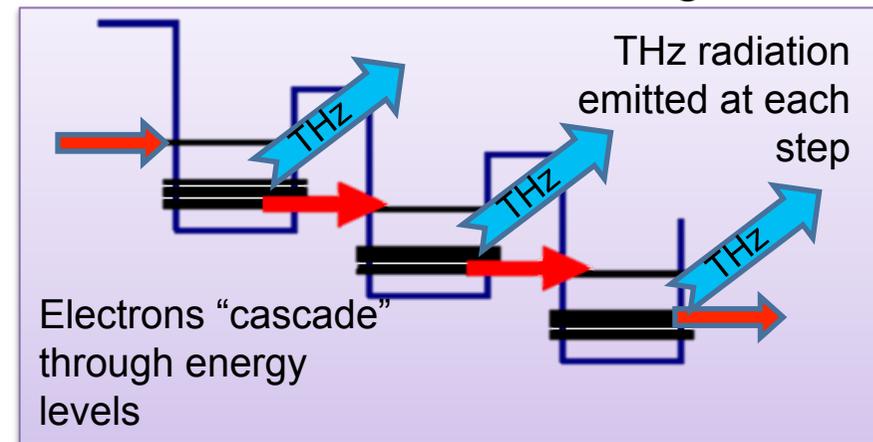
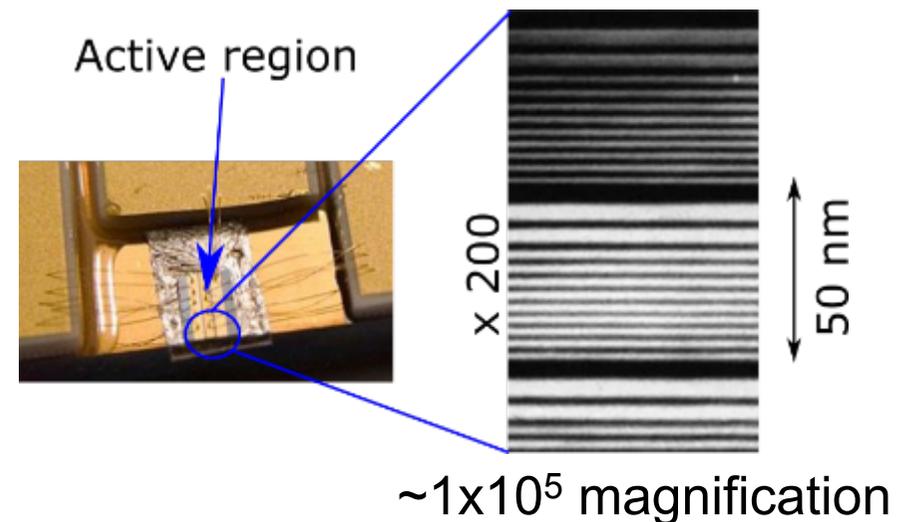
# Quantum cascade lasers



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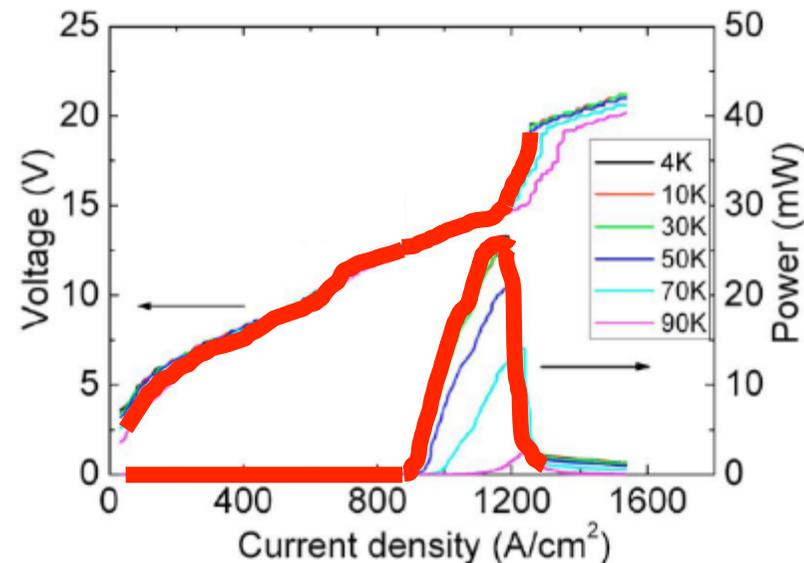
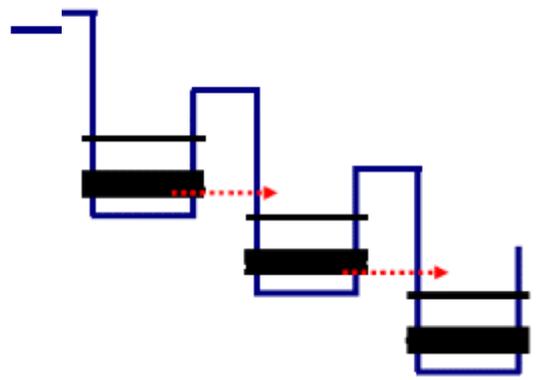
The first powerful and compact continuous-wave THz source:

- ~1000 semiconductor layers, grown using molecular-beam epitaxy
- “Electron-recycling” → efficient THz generation
- 1 W pulsed THz power; ~100 mW continuous-wave
- 1–5 THz range



Peak THz power corresponds to efficient injection of current:

- Lower “upstream” energy bands align with upper “downstream” bands
- Population inversion yields THz gain

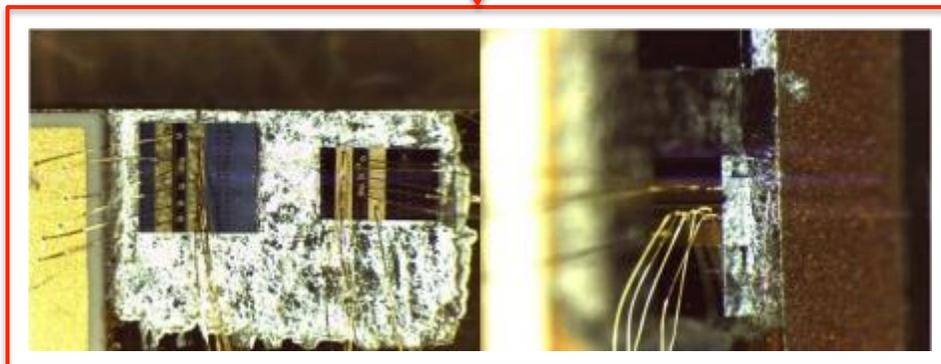


# LOCUS Core Technology

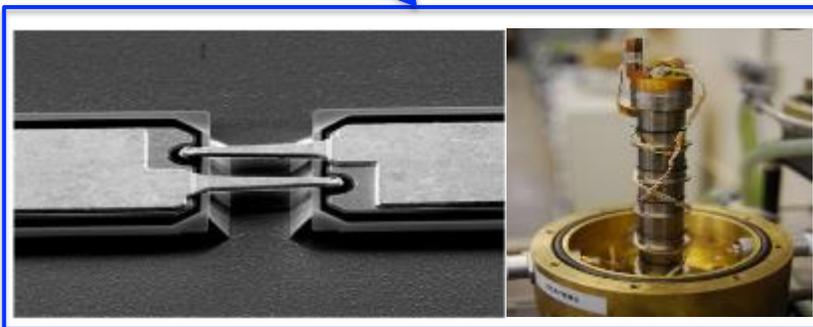


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3.5 & 4.7 THz QCL  
Local Oscillators  
University of Leeds



Schottky Barrier Diode  
& Space Coolers RAL



Digital Spectrometer  
STAR-Dundee



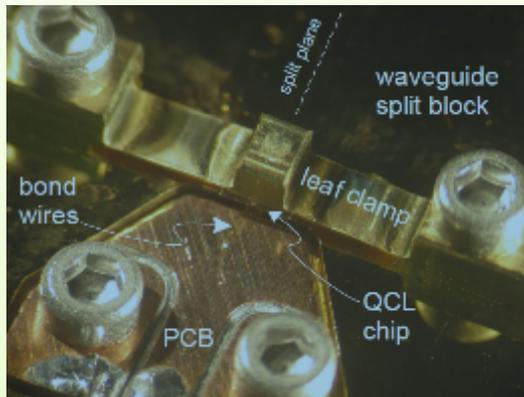
UK also leading LOCUS science definition via  
Leeds, UCL and RAL

Small Satellite  
Surrey Satellites Ltd

# Recent integration approaches

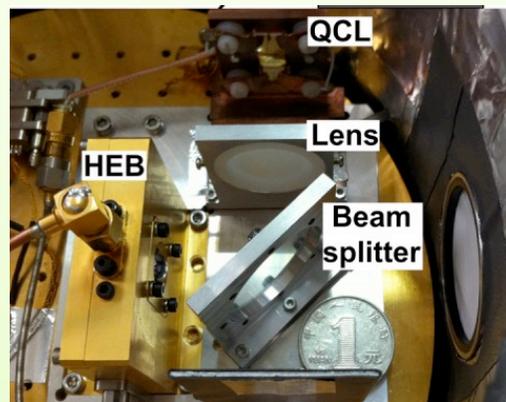


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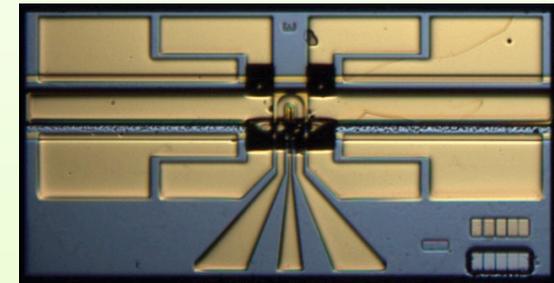
**QCL**  
**+ waveguide**  
**+ horn antenna**

Justen et al., 26<sup>th</sup> Int.  
Symp. Space THz Tech  
(2015)



**QCL**  
**+ HEB mixer**

Miao et al., *Opt. Express*  
**23**, 4453 (2015)



**QCL**  
**+ Schottky mixer**  
**(monolithic)**

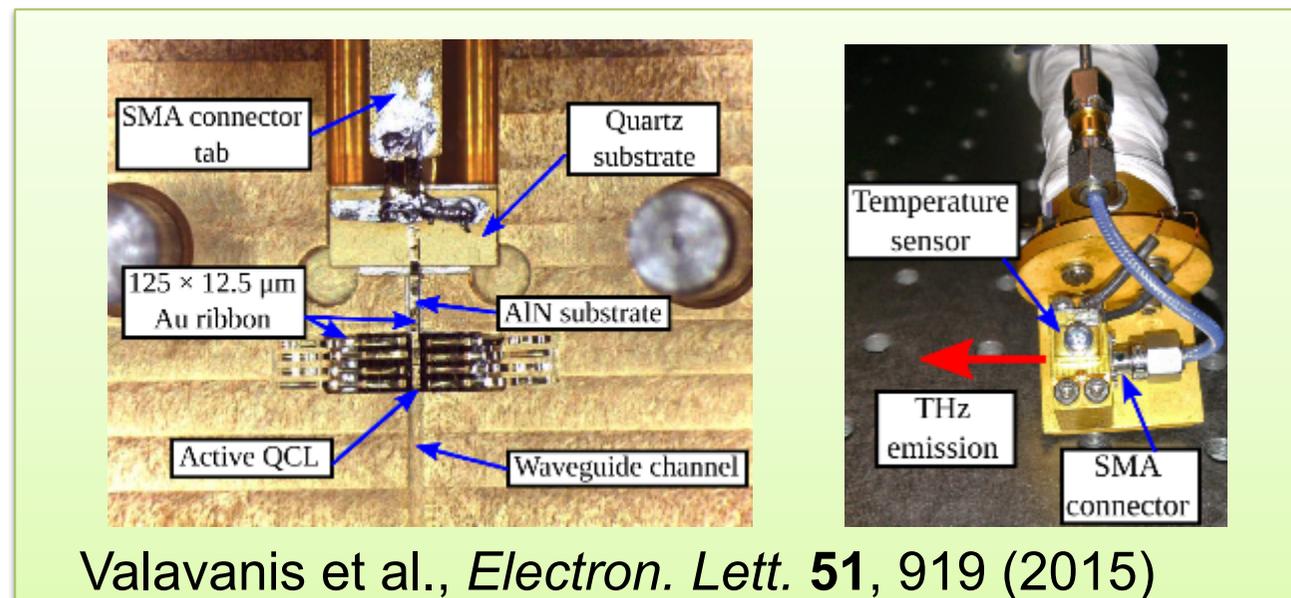
Wanke et al., *Nat. Photon.* **4**, 565 (2010)

# LOCUS integration design



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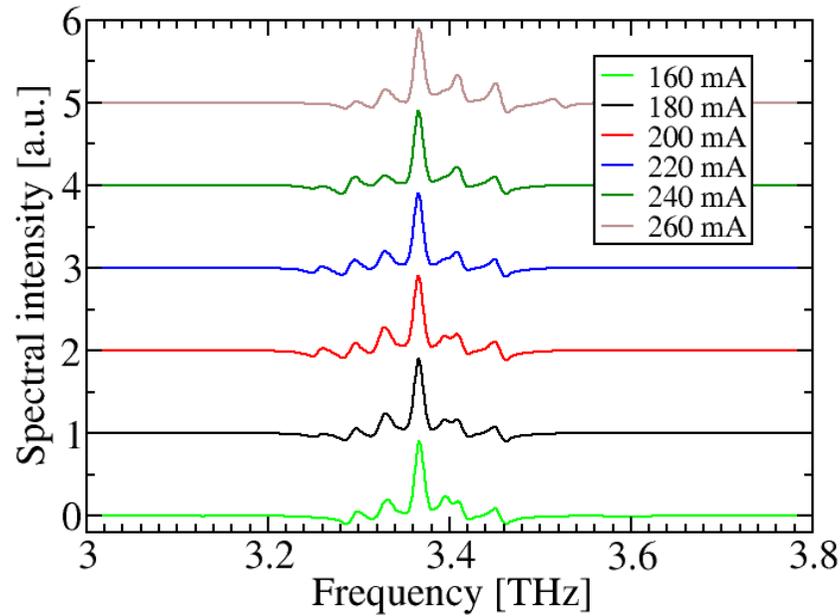
- Double metal 3.5 THz QCL
- Precision-micromachined  $300 \times 150 \mu\text{m}$  Cu waveguide
- High-frequency electronic ribbon-bonding + SMA
- Integrated temperature sensor



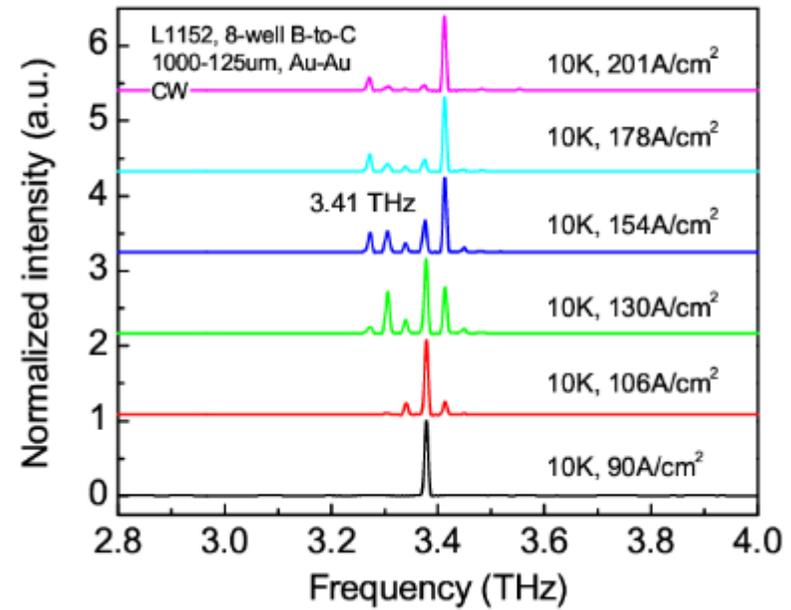
# CW characterisation



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Mounted

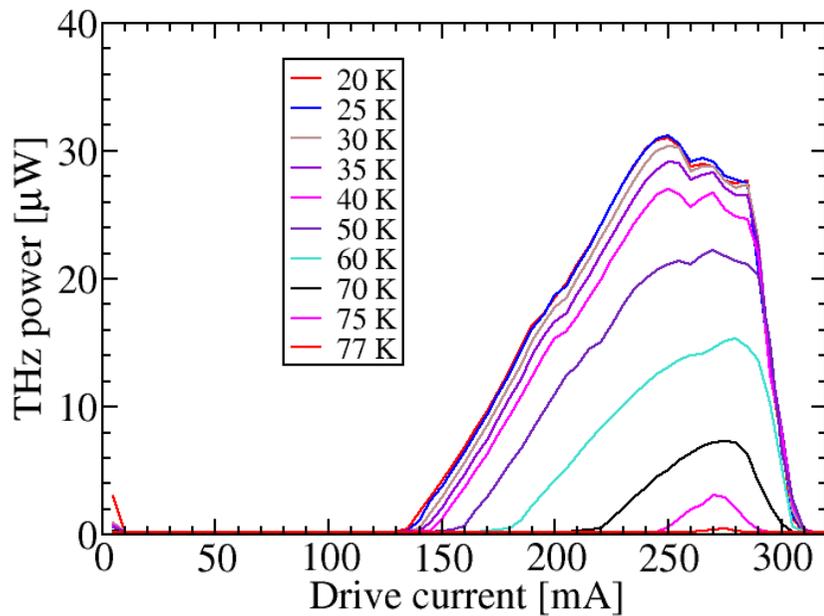


Unmounted

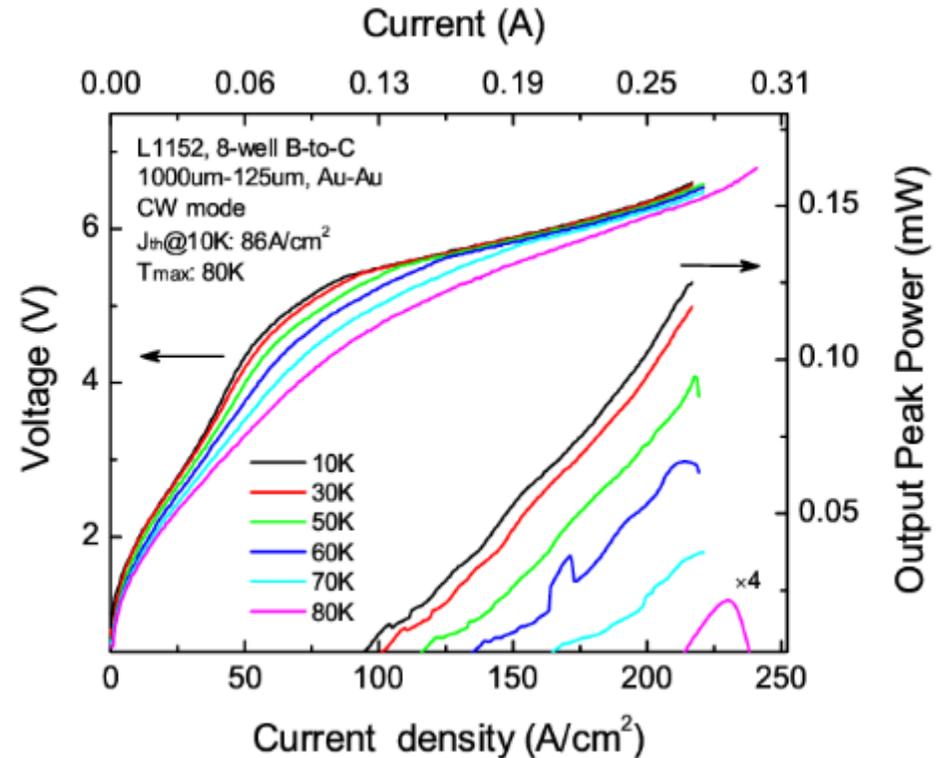
# CW characterisation



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Mounted



Unmounted

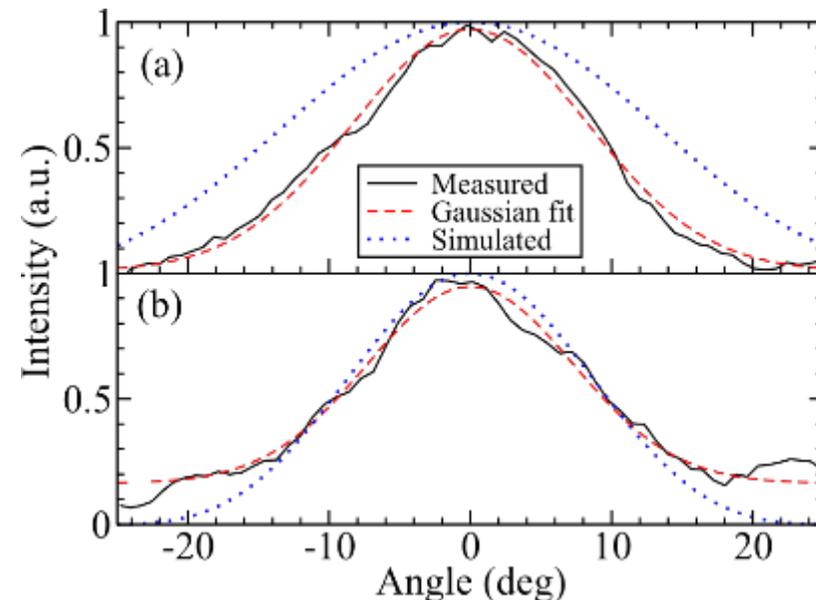
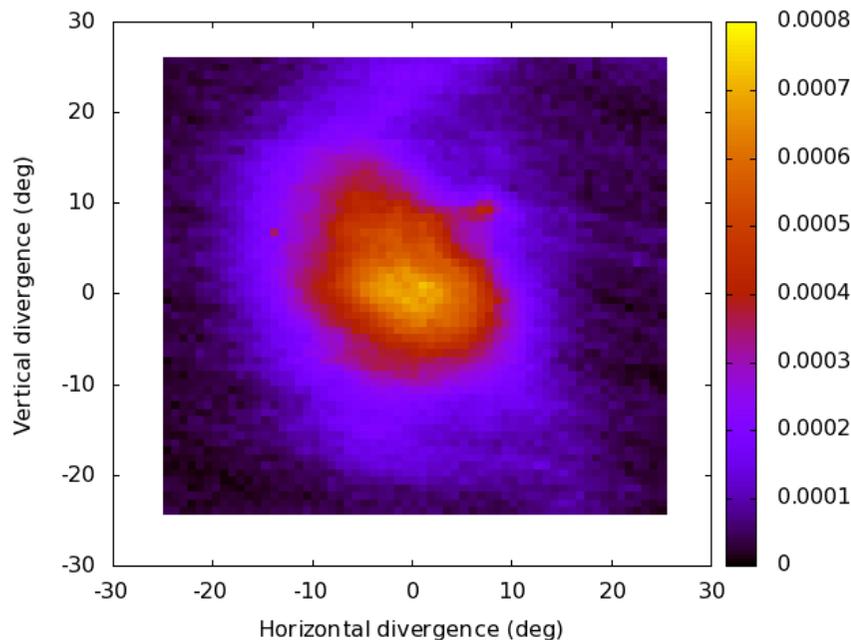
Block integration concept works! Minimal change in threshold current or maximum operating temperature.

Collected THz power reduced to  $\sim 20\%$ ... Optimisation needed!

# Waveguide integrated QCLs



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## Far-field THz beam-pattern significantly improved:

- Almost Gaussian profile
- Divergence = 17.1-deg (in-plane) / 19.7-deg (growth direction)
- Dramatic improvement over unmounted QCL (~120-deg)
- Underpins future systems with **no external optics**

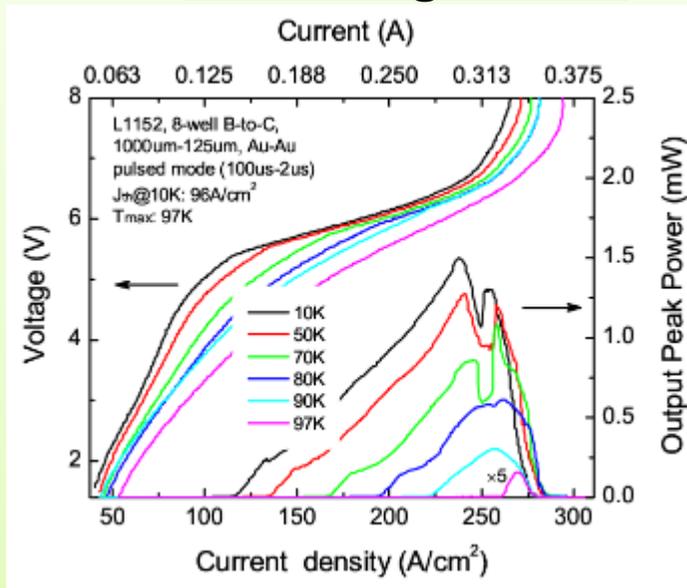
Valavanis et al., *Electron. Lett.* **51**, 919 (2015)

# QCL optimisation



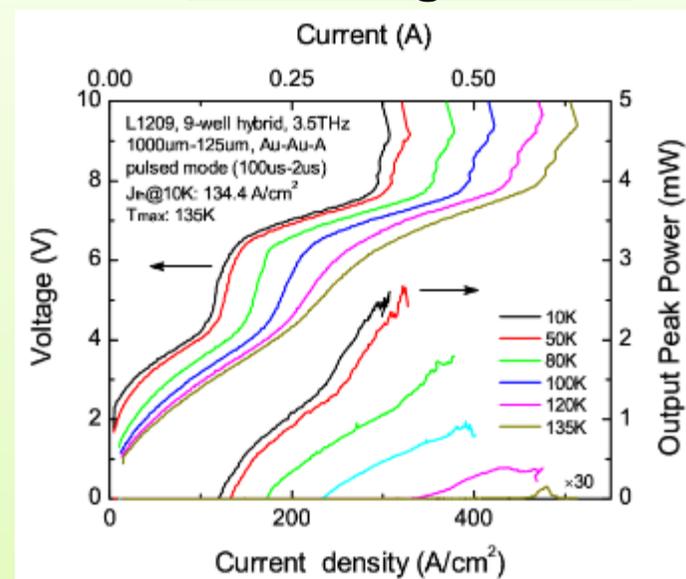
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## Active region "A"



G. Scalari et al., APL 82, 3165 (2003)

## Active region "B"



M. Wienold et al., *Electron. Lett.* **45**, 1030 (2009) [rescaled 3.1 to 3.5 THz]

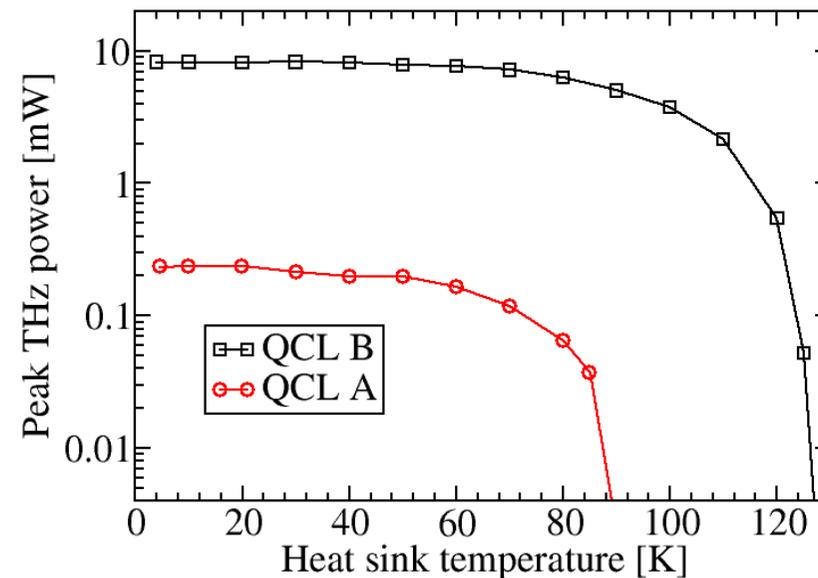
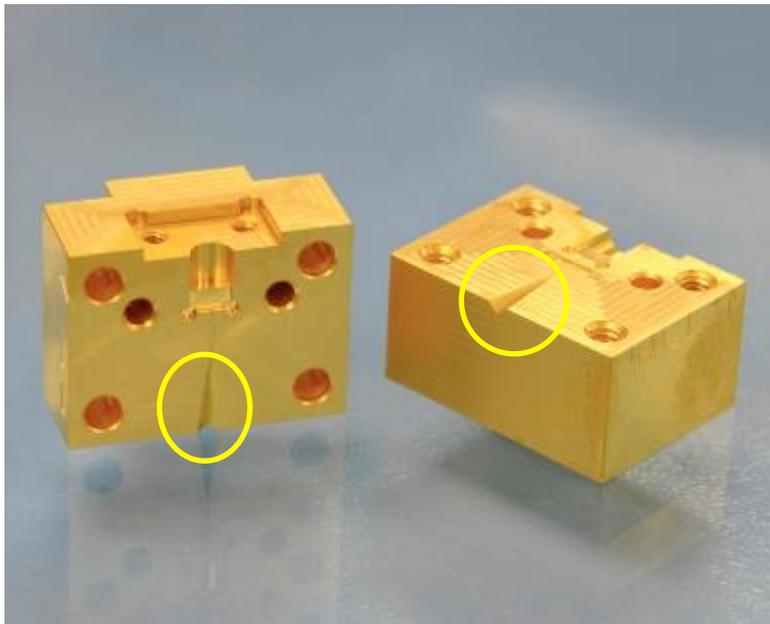
System	$f$ (THz)	$T_{max}$ (K) (pulsed/cw)	$J_{th}$ (A/cm <sup>2</sup> , 10K) (pulsed/cw)	$P_{max}$ (mW, 10K) (pulsed/cw)	$P_{dis}$ (W, 10K) (pulsed/cw)
A	3.27–3.45	97/80	<b>96/86</b>	1.5/0.12	<b>1.79</b>
B	3.31–3.58	<b>135/86</b>	134/133	<b>2.6/0.41</b>	3.10

# Feedhorn integration



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Diagonal horn-antenna with optimised QCL + waveguide



Preliminary results:

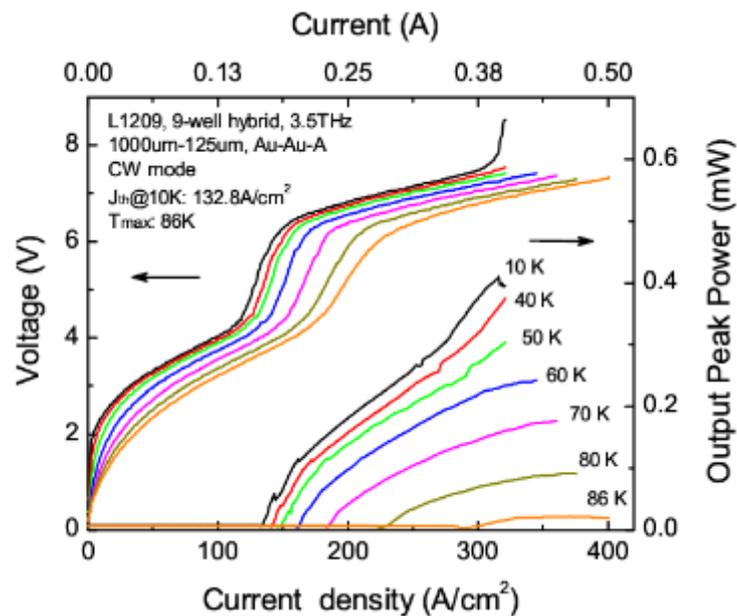
- 8.2 mW output power
- 127 K pulsed operation (80 K, cw)
- 6.2 mW @ 77 K (1 mW, cw)

# Thin QCL substrates

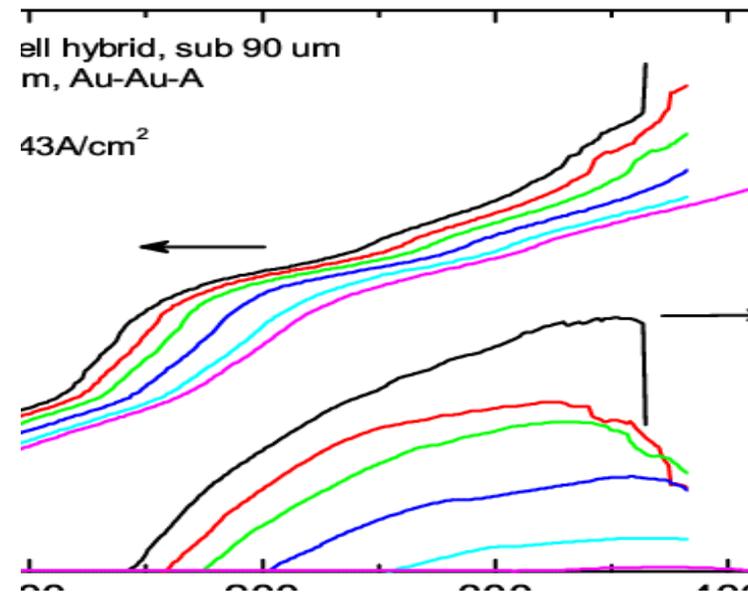


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Reducing substrate thickness allows direct mounting of QCL in waveguide channel, and reduced power dissipation.



**Device:** 1000 $\mu\text{m}$ ×125 $\mu\text{m}$ ×15 $\mu\text{m}$   
**Substrate:** 180  $\mu\text{m}$   
**T\_max:** 86 K  
**P\_dis:** 3.1 W



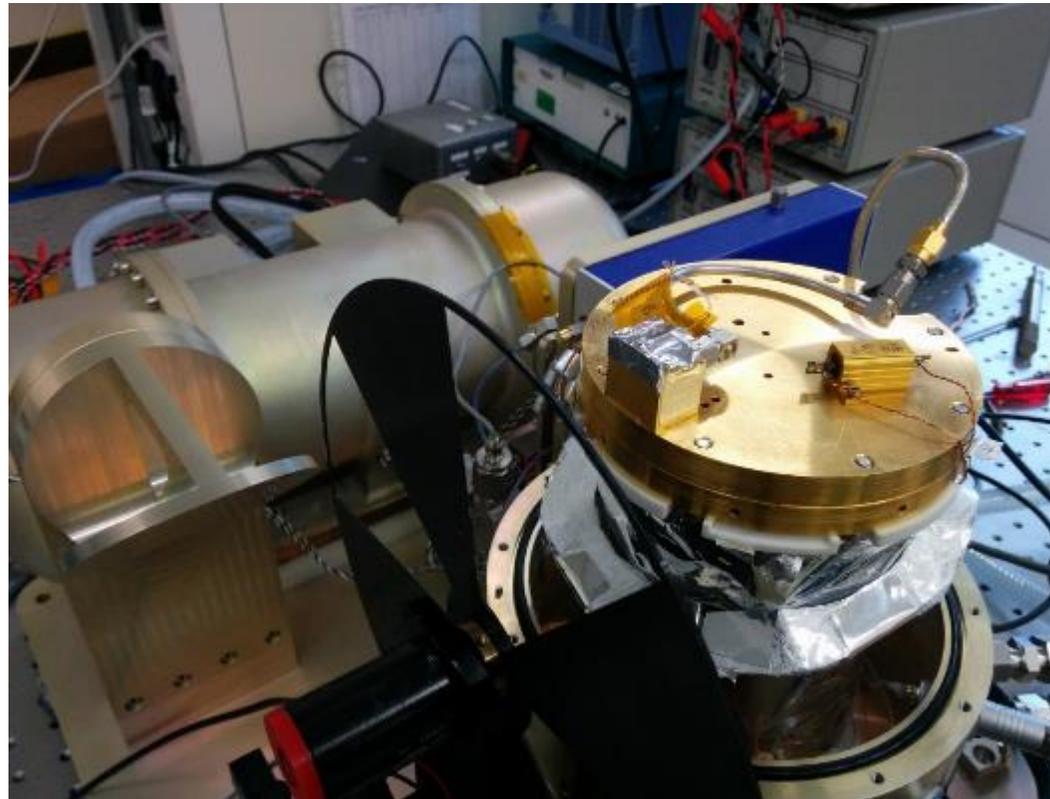
**Device:** 980 $\mu\text{m}$ ×75 $\mu\text{m}$ ×15 $\mu\text{m}$   
**Substrate:** 90  $\mu\text{m}$   
**T\_max:** 85 K  
**P\_dis:** 2.3 W

# Cryo-cooler performance



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3.5-THz QCL mounted in RAL Sterling cooler

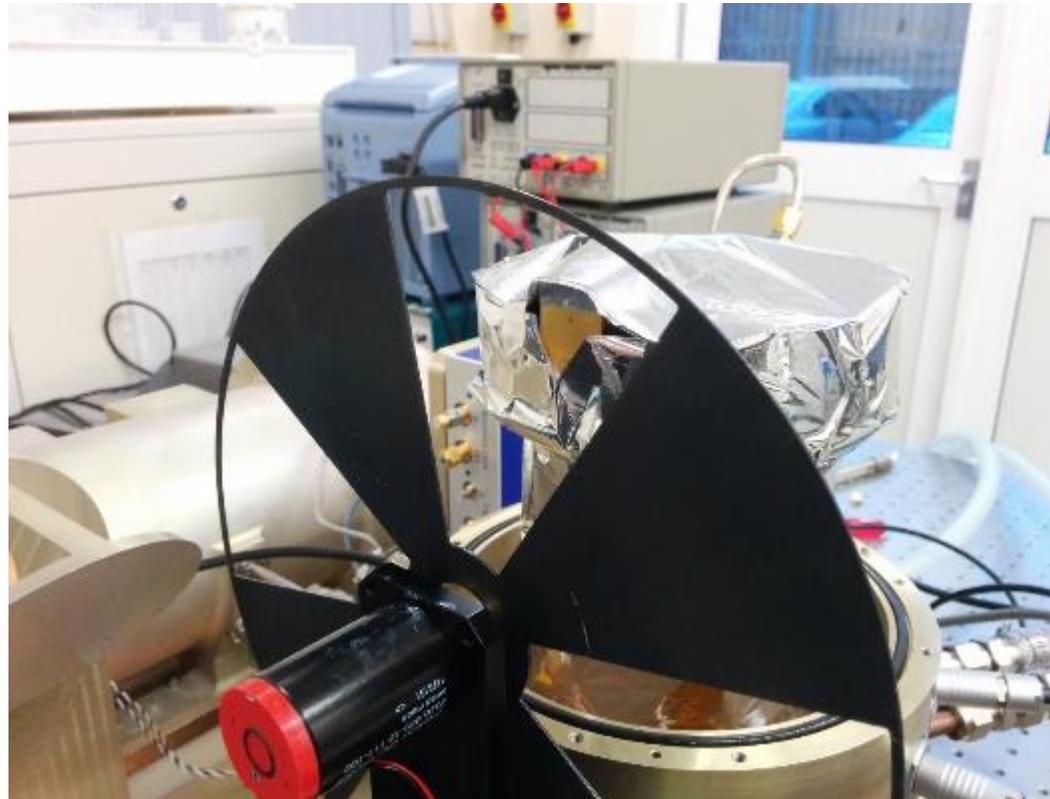


# Cryo-cooler performance



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3.5-THz QCL mounted in RAL Sterling cooler



# Cryo-cooler performance



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3.5-THz QCL mounted in RAL Sterling cooler



- Stable QCL operation at 65 K
- 1 mW THz power out-coupled from cryo-cooler into detector

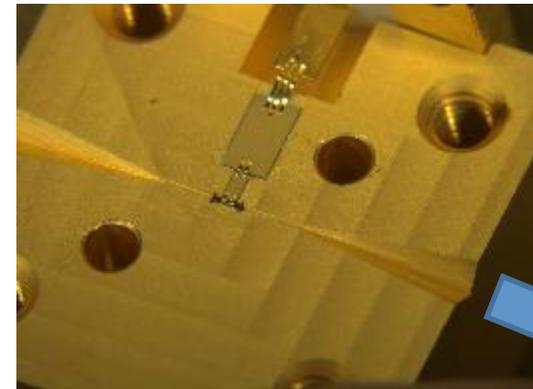
# Dual-feedhorn design



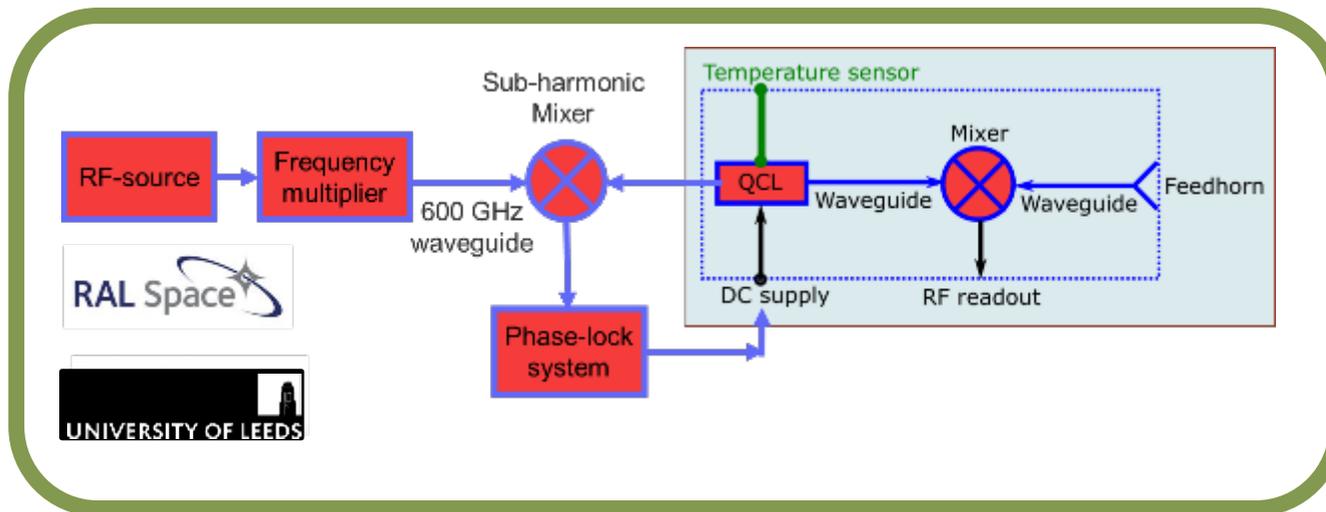
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- Dual-feedhorn design enables simultaneous access to **both facets** of QCL
- Will enable coupling with mixer and stabilisation subsystem

Stabilisation subsystem



THz mixer



RAL Space

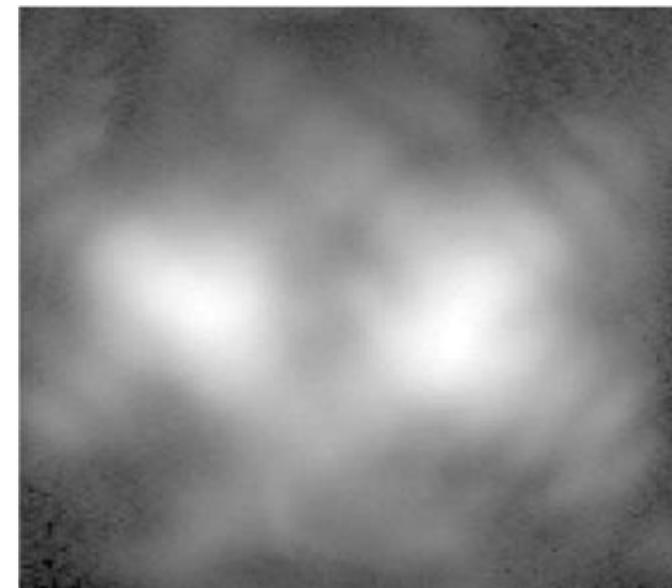
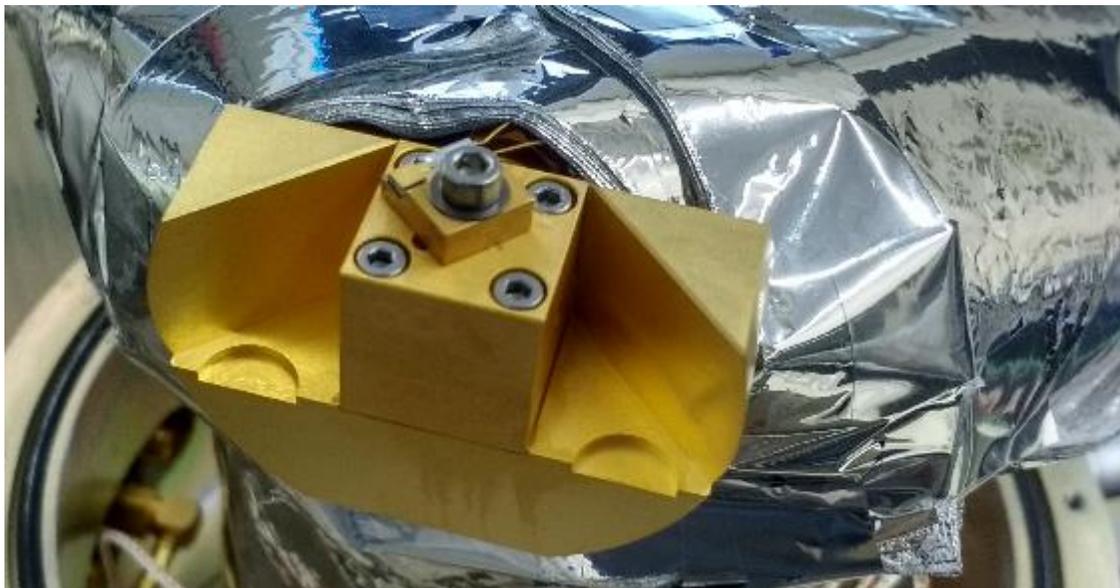
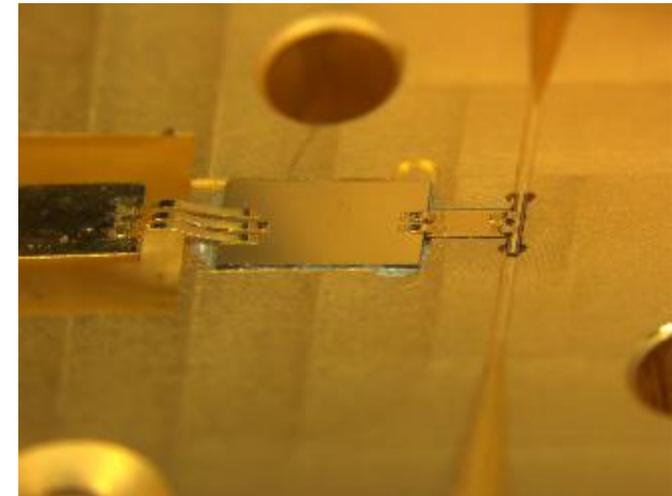
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# Dual-feedhorn design



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- Cryo-cooler operation demonstrated
- Diffraction/interference pattern observed
  - Coherent beam collection from both facets



# Future perspectives



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- Successful waveguide integration of QCLs has been demonstrated
- Key subsequent development steps:
  - Complete system breadboarding
  - Stabilisation subsystem integration
  - Mixer integration
  - Airborne/in-orbit demonstration

# Summary



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- **LOCUS:** A breakthrough THz limb sounder concept
- **THz QCLs:** The first compact, yet powerful direct THz sources
- **Waveguide integration of QCLs:** Progress towards complete THz radiometry systems in compact waveguide blocks, underpinning future satellite applications.

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- CEOI-ST (7<sup>th</sup>, 8<sup>th</sup> and 10<sup>th</sup> call)
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- ESA GSTP
- Royal Society & Wolfson foundation
- ERC
- STFC Centre for Instrumentation

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- University of Leeds (*Y. J. Han, J. R. Freeman, L. H. Li, E. H. Linfield, A. G. Davies*)