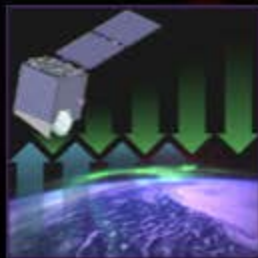


LINKING OBSERVATIONS OF CLIMATE, THE
UPPER ATMOSPHERE AND SPACE WEATHER

LOCUS



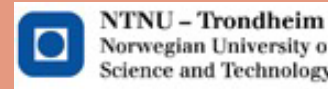
The LOCUS consortium(s)



LOCUS Elegant Breadboard (8th CEOI)



The LOCUS Satellite (EE9 proposal)

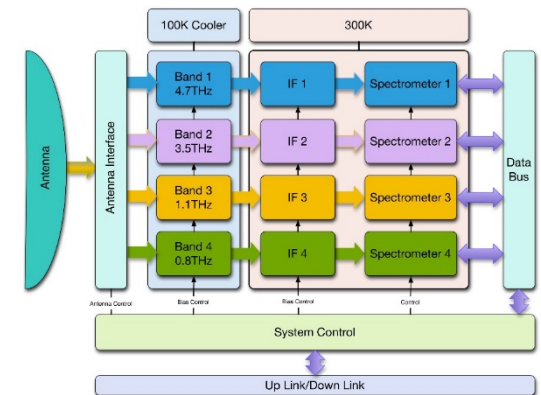
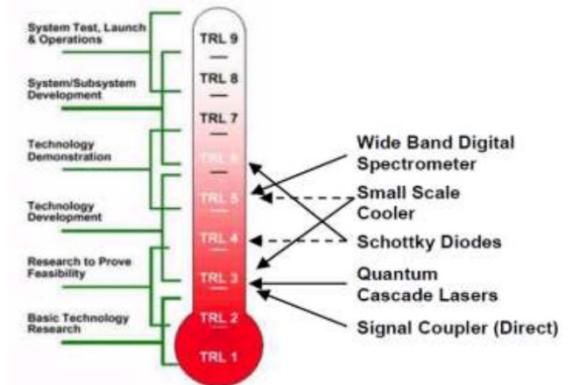


LOCUS payload development

RAL Space / UCL / Uni. of Leeds / STAR Dundee /
Uni. of Huddersfield / Durham Uni. / JCR Systems



- **LOCUS:** A new THz remote sounder for global mapping of the mesosphere/lower thermosphere (MLT). Original concept supported by CEOI
- **CEOI objective:** Raise payload TRL
- **7th call: Receiver TRL advancement**
 - 1.1-THz Schottky diode receiver
 - 3.5 & 4.7-THz Quantum-cascade lasers (QCLs)
 - Wide-band spectrometer enhancement
 - Space compact cryocooler
- **8th call: System elegant breadboarding**
 - Off-axis Cassegrain antenna development
 - Optical bench integration & test
 - Thermal vacuum tests of the integrated antenna



The LOCUS Elegant Breadboard



Brian Ellison,
Daniel Gerber
Simon Rea,
Hui Wang,
Manju Henry,
Martin Crook,
Tom Rawlings

Edmund Linfield,
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Michael Emes,
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Janet Charlton,
Soe Min Tun



Olivier
Auriacombe



Steve Parkes,



Guoyu Yu,
David Walker,
Hongyu Li



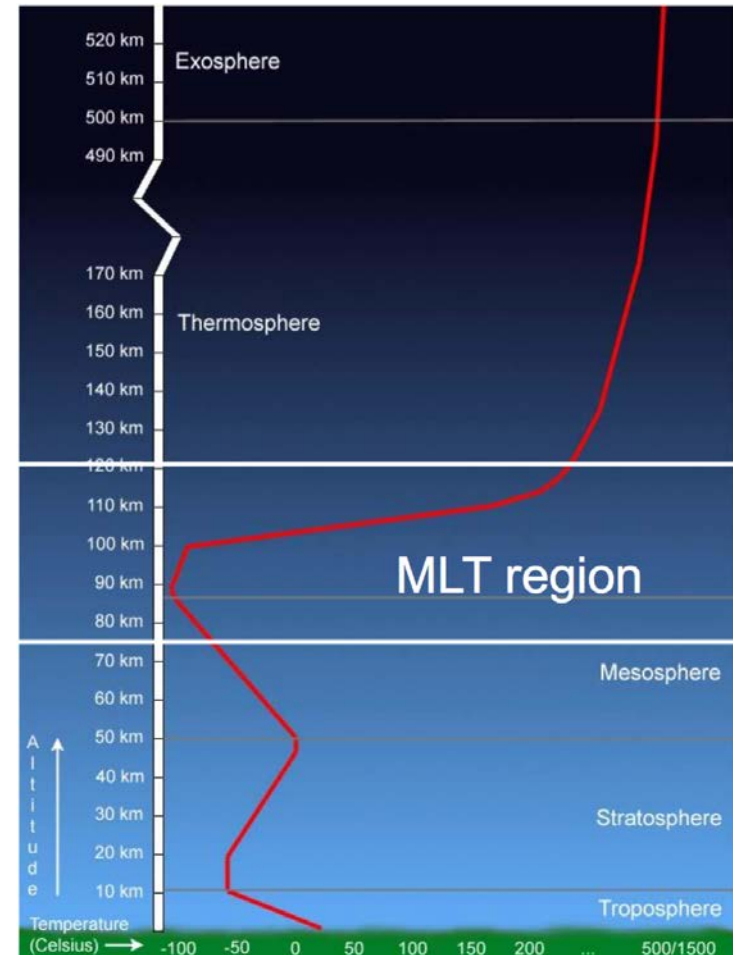
Cyril Bourgenot
David J Robertson



Christopher
Saunders,
Jonathan Friend

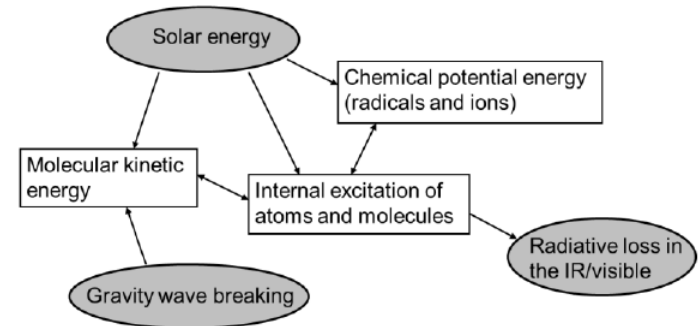


- The MLT (Mesosphere, lower Thermosphere) is a “gateway” between Earth’s atmosphere and near space environment
- Indicative of climate change through:
 - Increased cooling rates
[Beig et al., JGR, 2011](#)
 - Increase of mesospheric clouds
[DeLand et al., JGR, 2015](#)
- Not well explored, because it is:
 - Too high for balloons; too low for orbiters
 - Many key species only detectable at THz frequencies



- Key MLT observables are:

- O: Understanding thermal balance and chemistry of Upper Atmosphere (O not globally measured; drives cooling actively & through quenching rates)
- NO (formed in the MLT) is believed to be transported down, impacting on O₃ chemistry. (relevant to Climate Change)
- OH, HO₂, CO: Understanding Upper Atmospheric chemistry and dynamics
- Formation of noctilucent clouds (correlation with atomic Oxygen depletion)

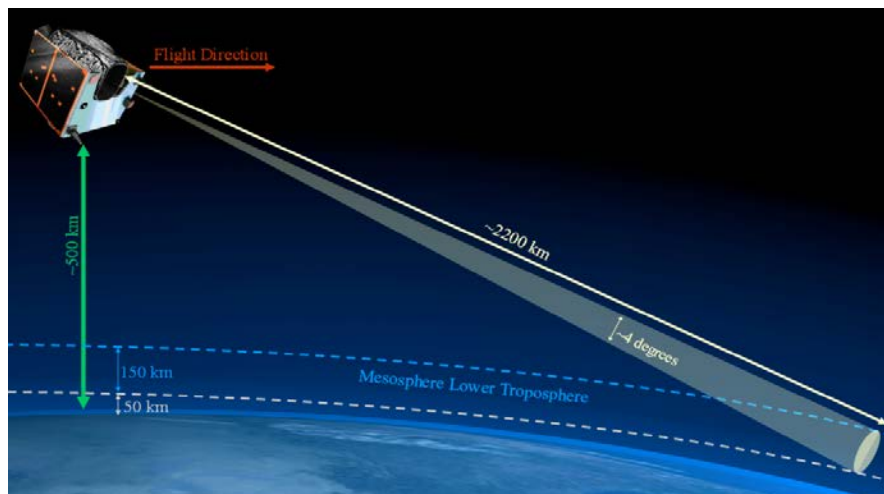


The LOCUS concept

A breakthrough THz remote sounder: www.locussatellite.com



- A limb scanning observatory with THz and IR channels to identify the relevant species.
 - The IR radiative fluxes (which have already been “accessed” in the past allows us to better understand the thermal equilibrium of the MLT but not its chemical composition
 - Simultaneous measurements of IR emission from vibrationally excited CO₂, NO, O₃ and OH will provide complementary information on thermal structure, polar mesospheric clouds (PMC), energetics and chemical processes.



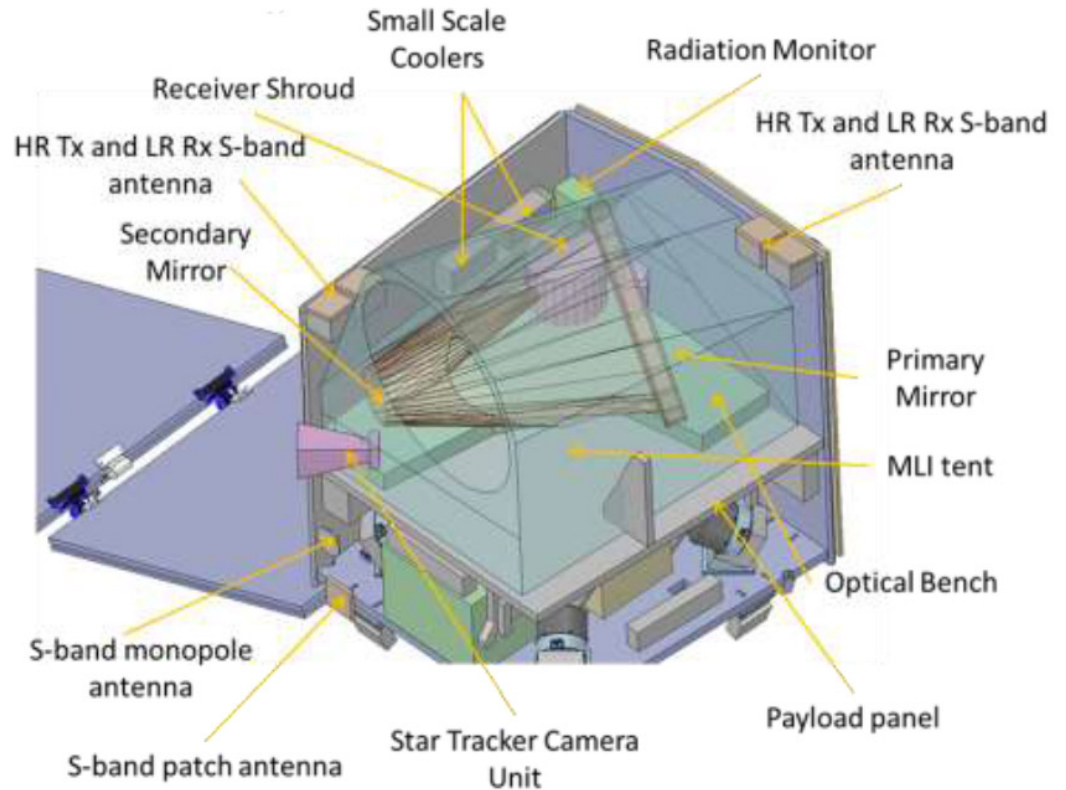
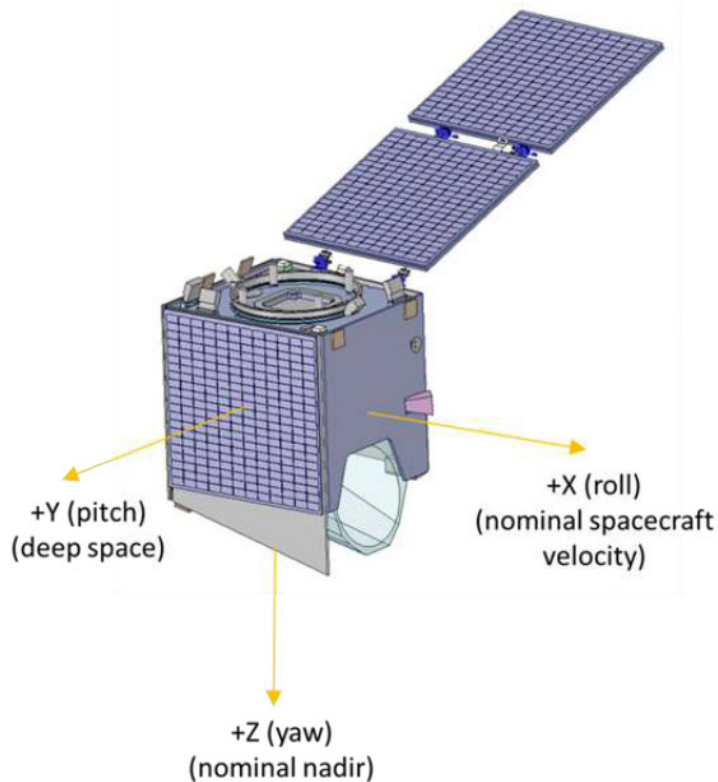
Band	Centre (THz)	ν	Species Covered	Predicted Performance (NETD K)
1	4.7		O, O ₃	46
2	3.5		OH, CO, HO ₂	12
3	1.15		NO, O ₃	4
4	0.8		O ₂ , O ₃	3

TABLE I. THz RECEIVER BAND DESIGNATION

Centre λ (μm)	Bandwidth (μm)	Species Covered	Required Detectability ($W m^{-2} sr^{-1}$)
15	5.2	CO ₂	1×10^{-3}
9.3	1.74	O ₃	3×10^{-4}
5.3	0.41	NO	1×10^{-5}
2.07	0.54	OH	1×10^{-5}

TABLE II. IR RECEIVER BAND DESIGNATION

LOCUS Satellite & Payload

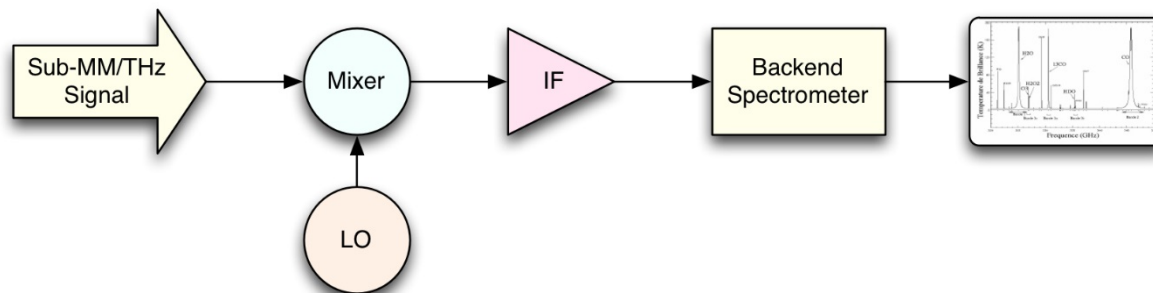


SSTL150 satellite and payload concept from SSTL led ESA IOD.

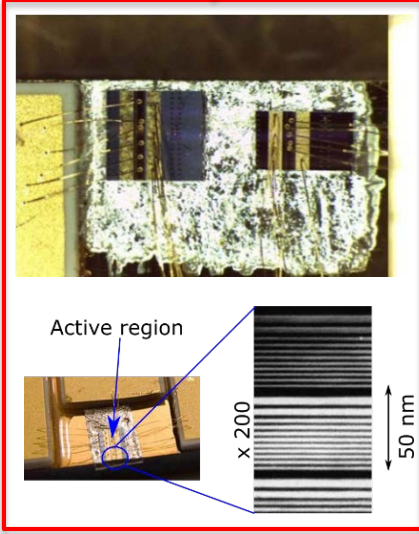
Basic heterodyne system concept



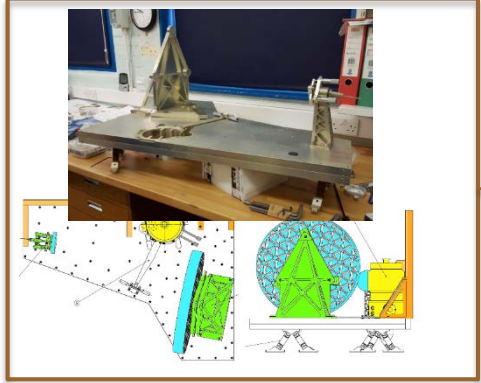
- LOCUS uses heterodyne frequency down-conversion technology (mixers).
- Quantum cascade laser (QCL) provides local oscillator (LO).
- Small active coolers required to cool QCLs.
- Achieves ~ 1 MHz spectral resolution via high-speed digital sampling.



LOCUS Critical Technologies

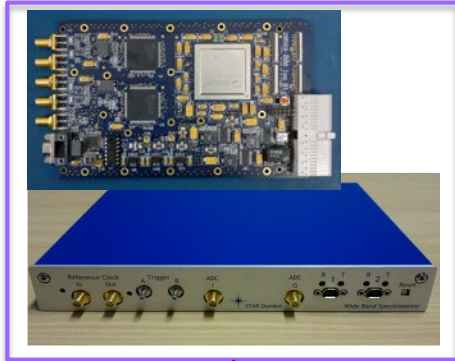
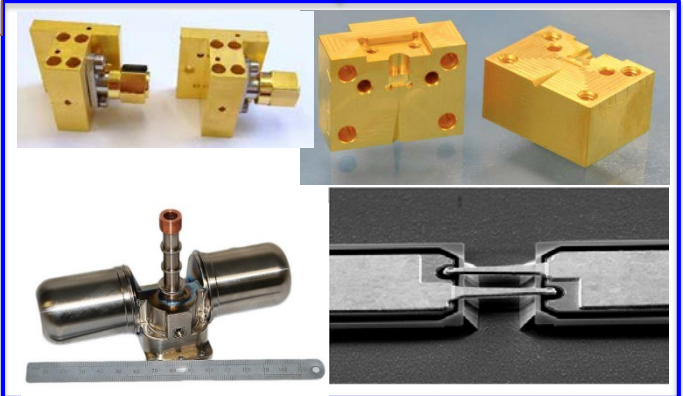


3.5 & 4.7 THz QCL Local Oscillators
University of Leeds

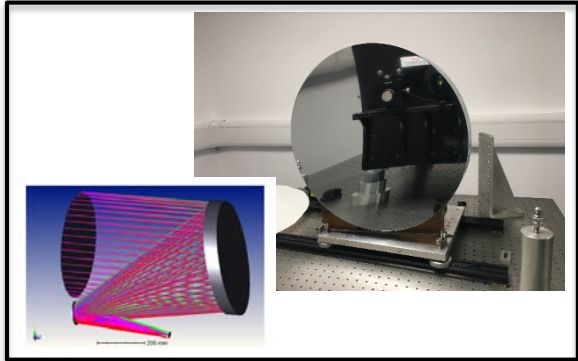


Optical Bench, telescope mounts, thermal model (UCL-MSSL)

Schottky Barrier Diode & Space Coolers RAL



Wide-band Spectrometer STAR-Dundee



All-Aluminium compact Cassegrain telescope, IR pixel (UCL, Glyndwr Uni., Huddersfield Uni., Durham Uni.)

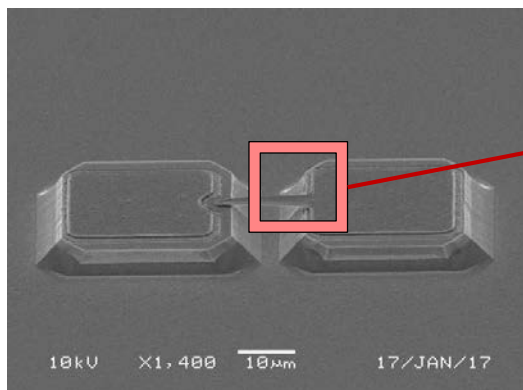


Small Satellite Surrey Satellites Ltd

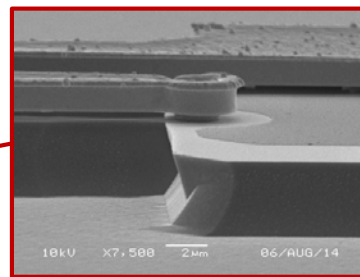
Critical TRL: Supra-THz Mixers and LOs



- Developing core mixer technology at 1.15THz (to 4.7THz via GSTP).
- Use electron beam lithography to define GaAs Schottky diode – reduce shunt capacitance to improve responsivity at THz frequencies
- Move to low bias InGaAs devices to reduce LO power.
- For 1.15THz & below, use harmonic LO chain.



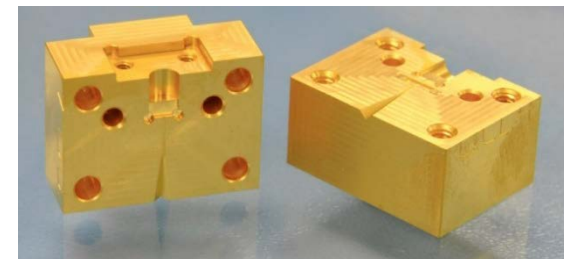
Flip-chip diode



Under-cut walls
near
Schottky Contact



7th Call 1.15THz mixers



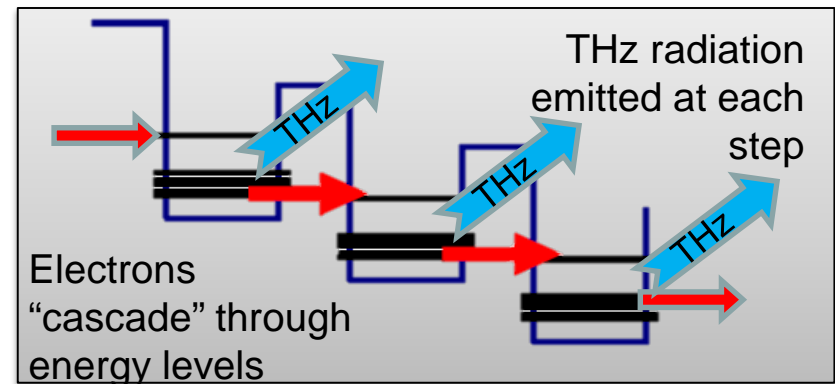
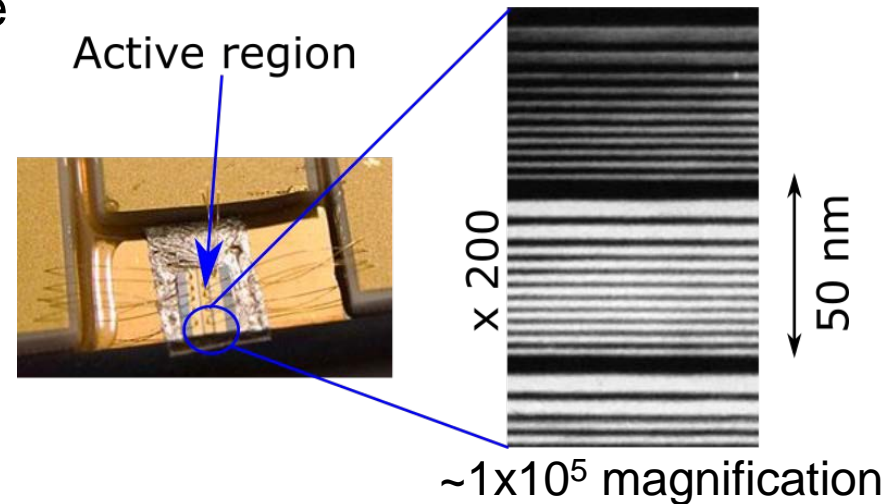
Split block receiver with horn

- For 3.5THz / 4.7THz, integrate with QCL in single waveguide block.

Critical TRL: Quantum cascade lasers

Powerful and compact continuous-wave THz source for LO:

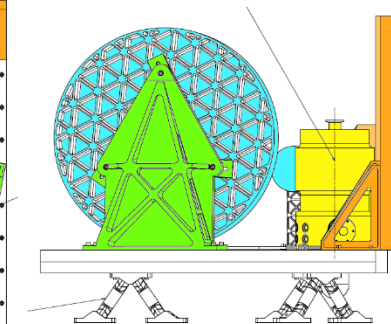
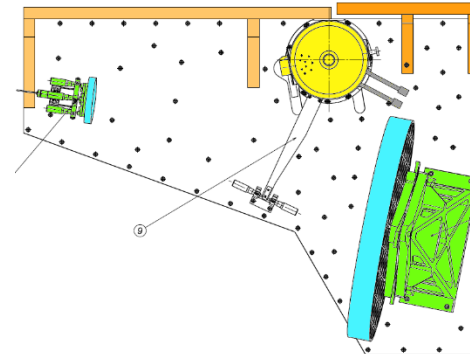
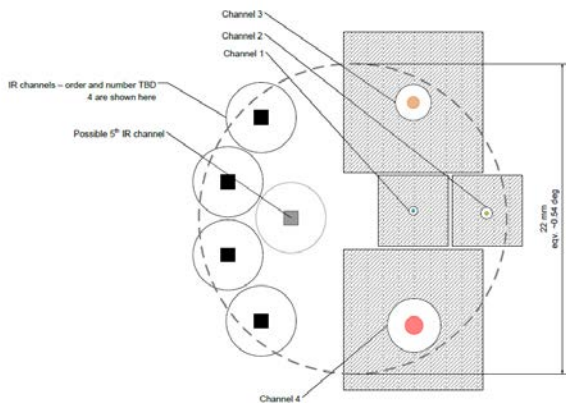
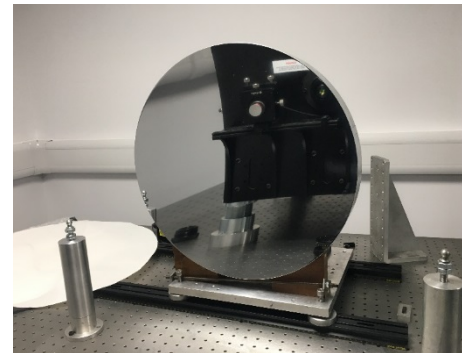
- “Electron-recycling” → efficient THz generation.
- Fabricated using molecular-beam epitaxy.
- Provided mW output at key LOCUS frequencies – 3.5/4.7THz.
- Waveguide integration successfully achieved.
- Operation demonstrated in an active space-cooler environment (~60K).



LOCUS breadboard and optics

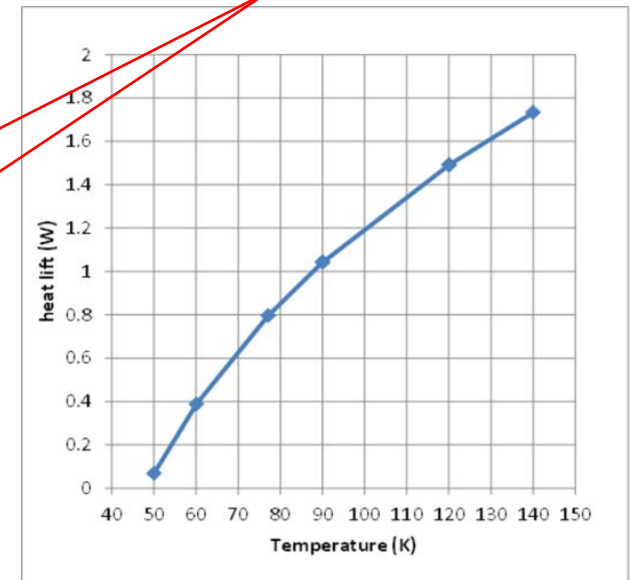
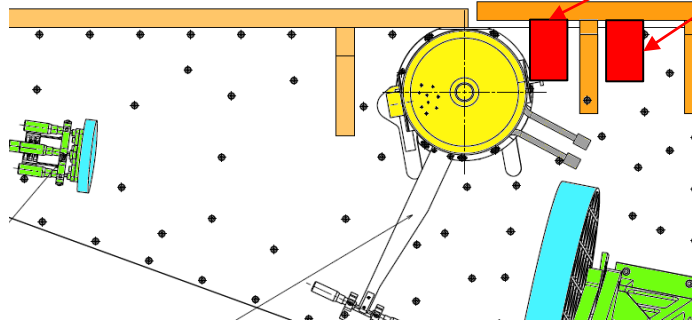
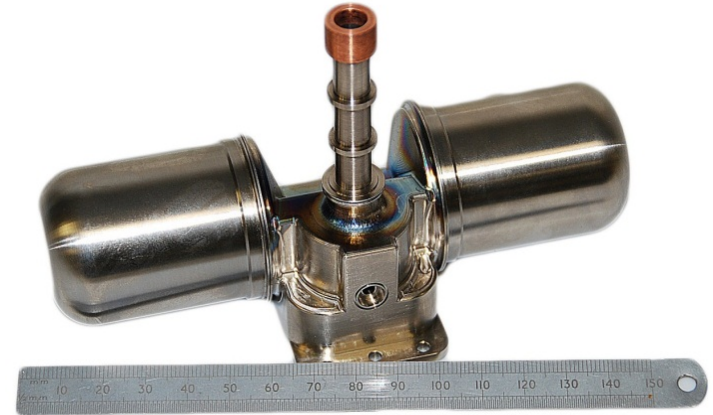


- Off-axis Cassegrain antenna design.
- Optics & receiver mounted onto an aluminium optical bench.
- Compact cryocooler provides 50 K stage for QCL operation.
- IR observation capability.
- Activity will raise to TRL 4/5.



Tactical Space Coolers

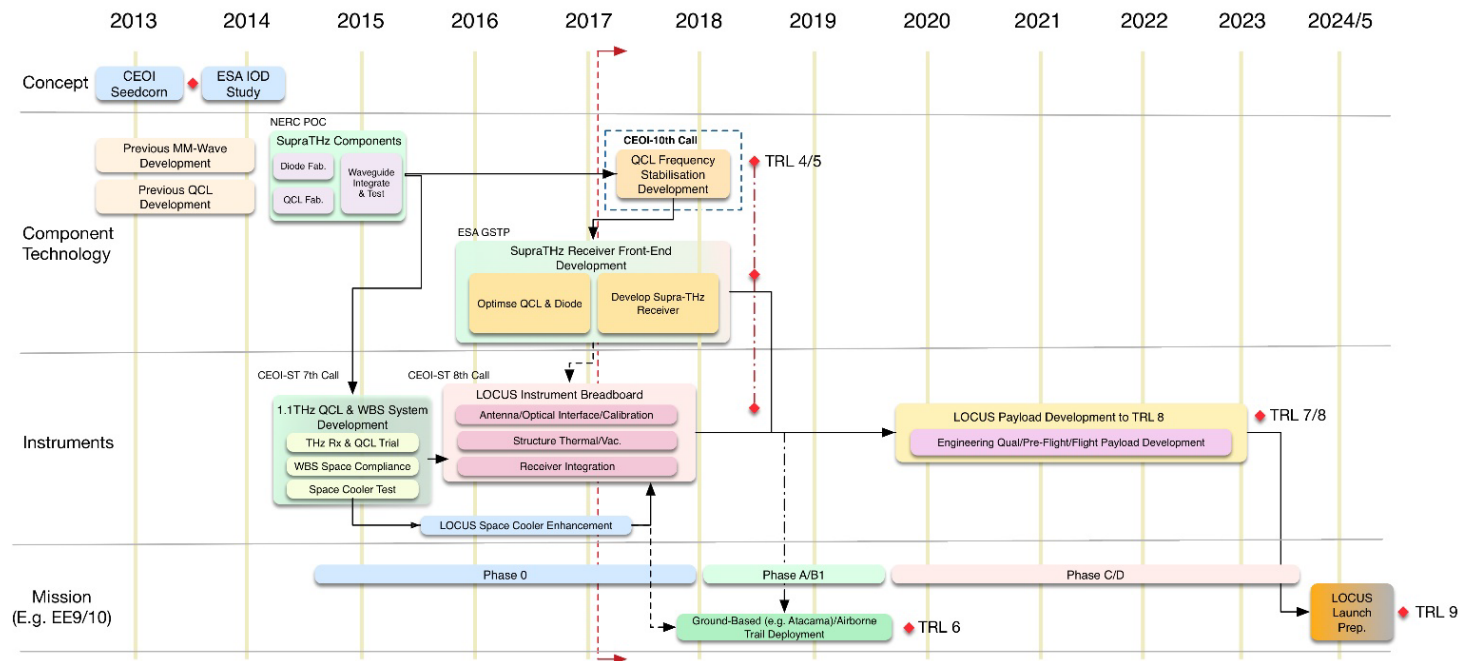
- The RAL tactical mini-cooler provides 1W of cooling at 100K
- LOCUS would be the first Small-class satellite in Low-Earth Orbit (to our knowledge) with an active cooler
- A purposely built radiator to which the cooler is connected, will dump the power generated ~ 23 W/cooler at the hot end



LOCUS roadmap



- CEOI 7/8th call advances receiver/system to TRL 4/5
- IOD concept study completed for ESA.
- Airborne demonstration proposed for CEOI 10th Call (unsuccessful)
- Future candidate for EE10 mission



The LOCUS Elegant Breadboard



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Christopher
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Jonathan Friend



Janet Charlton,
Soe Min Tun

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Cyril Bourgenot
David J Robertson



Thanks for listening

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- THz Receivers: brian.ellison@stfc.ac.uk
- QCL technology: a.valavanis@leeds.ac.uk
- Telescope and Optics: g.savini@ucl.ac.uk