

STEREOID: an Earth Explorer 10 mission candidate to observe land, ice, and ocean dynamics

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STEREOID for Earth Explorer 10

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Proposing Team			
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15	Tebaldini, Stefano, Assistant Professor	Politecnico di Milano	IT
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19	Esposito, Marco	Cosine	NL

Earth Observation (ACEO),
Proposals submitted in the
the procedures followed by

September 2017 (cf. ESA/PB-

id down in the Call, been

port of two scientific panels.
a total of 32 external experts.
technical and programmatic
ervation Programmes and the
s carried out using the seven

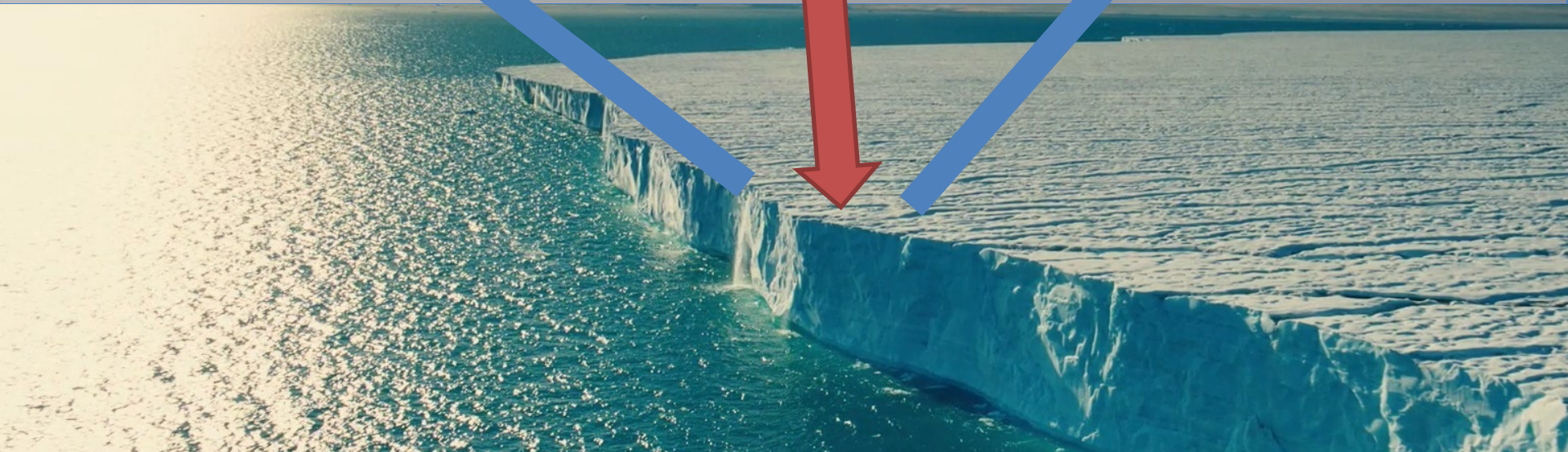
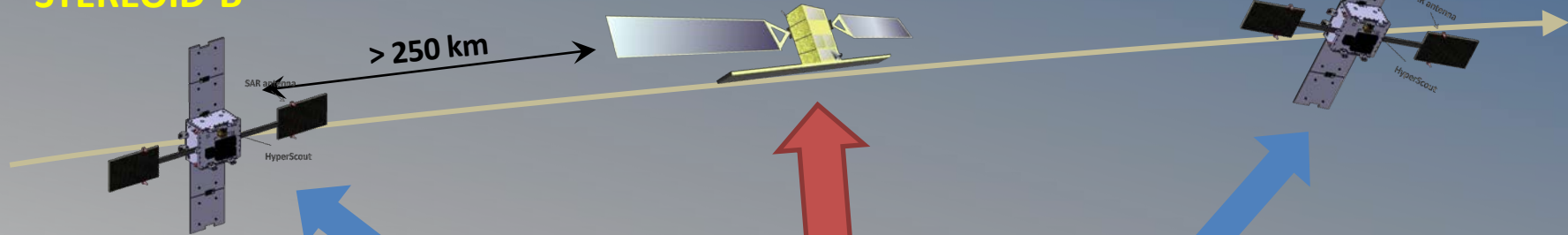
ACEO recommends three Mission Ideas for further assessment studies at pre-feasibility level, i.e. phase 0 study (without order of priority), namely **STEREOID**, Daedalus and G-CLASS:H2O.

Sentinel-1 D

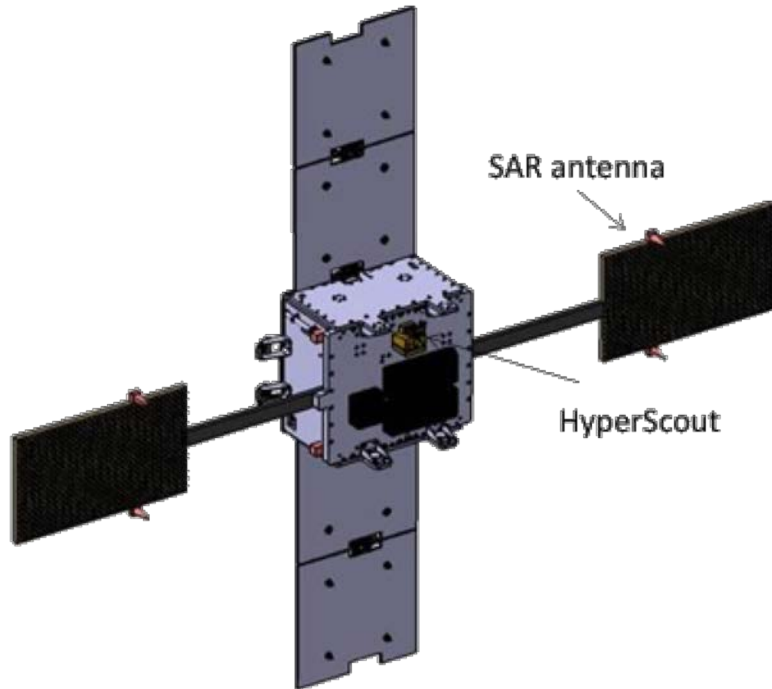
STEREIOD-A

STEREIOD-B

> 250 km



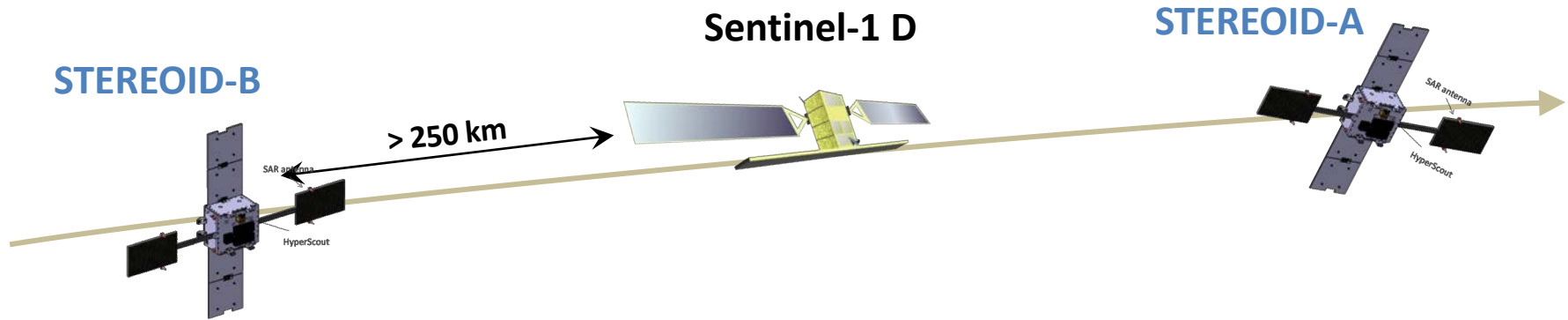
Space-segment (in proposal)



Platform

- < 500 kg
- Which is small for an Earth Explorer

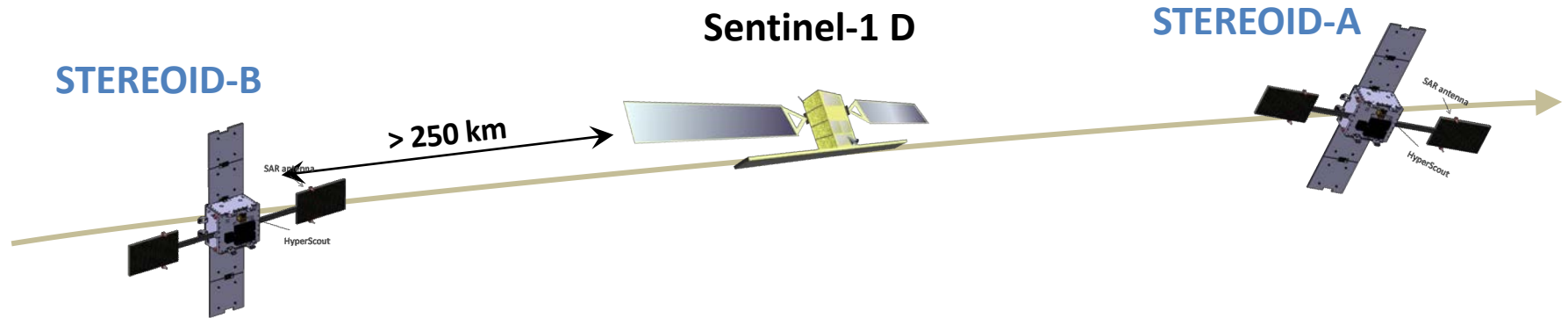
Mission phases: variable formation



- **Stereo formation**

- Maximum line-of-sight diversity
- Best for surface current vectors and 3-D surface deformation

Mission phases: variable formation

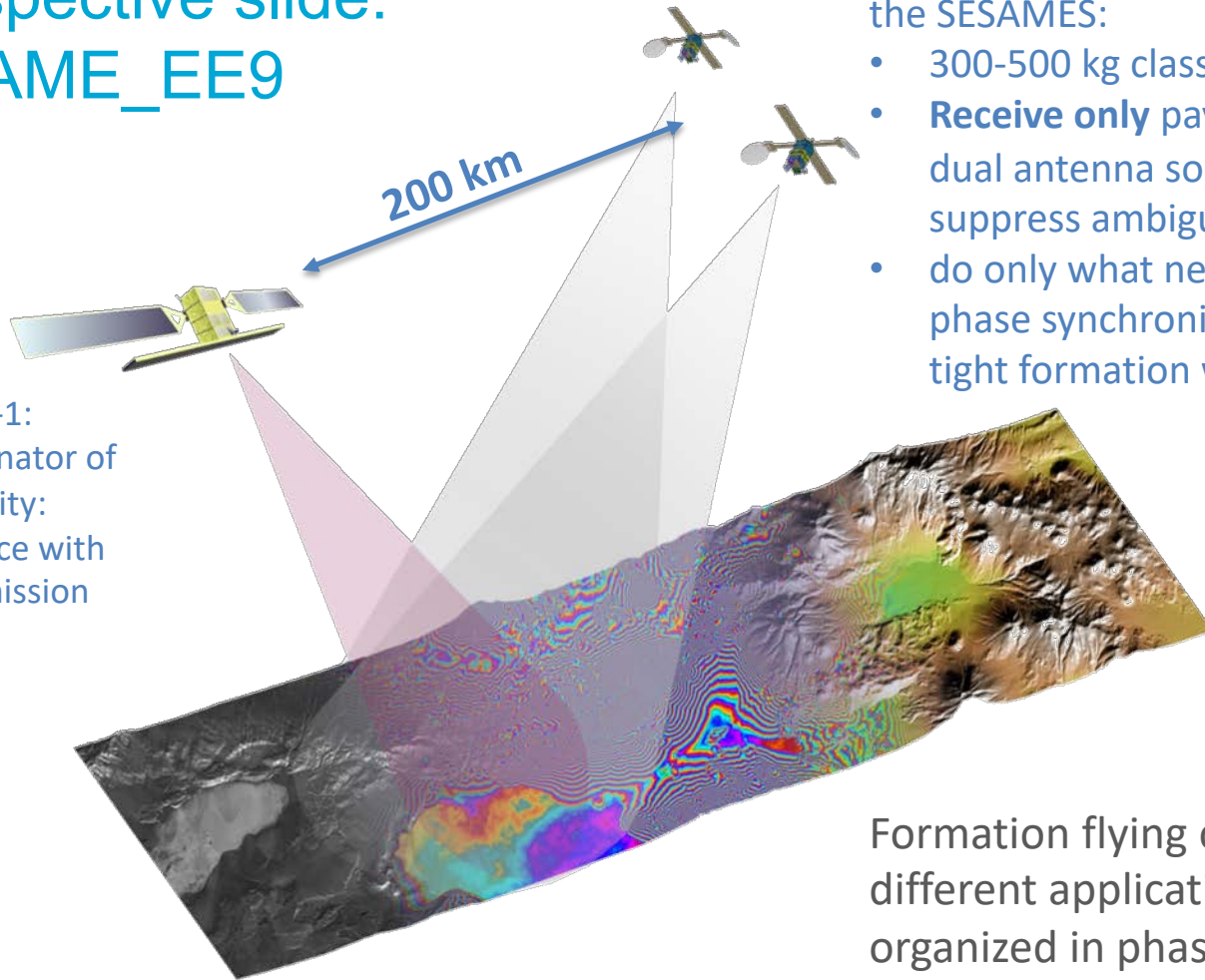


- **XTI formation**
 - Close-formation (TanDEM-X style)
 - Intended for DEM time-series
 - 400 m to 1 km baselines
- **ATI formation**
 - 100 m to 200 m along-track separation

One slide of history

Retrospective slide: #SESAME_EE9

Sentinel-1:
Used as illuminator of
opportunity:
no interference with
Sentinel-1 mission



the SESAMES:

- 300-500 kg class satellites
- **Receive only** payload with dual antenna solution to suppress ambiguities
- do only what needed: no phase synchronization nor tight formation with S1

Formation flying optimizes to different applications/ROIs: mission organized in phases.

Science objectives

Mission objectives (1 min overview)

Solid Earth

- 3D surface deformation (volcanic, seismic, landslides)
- Sudden topographic changes

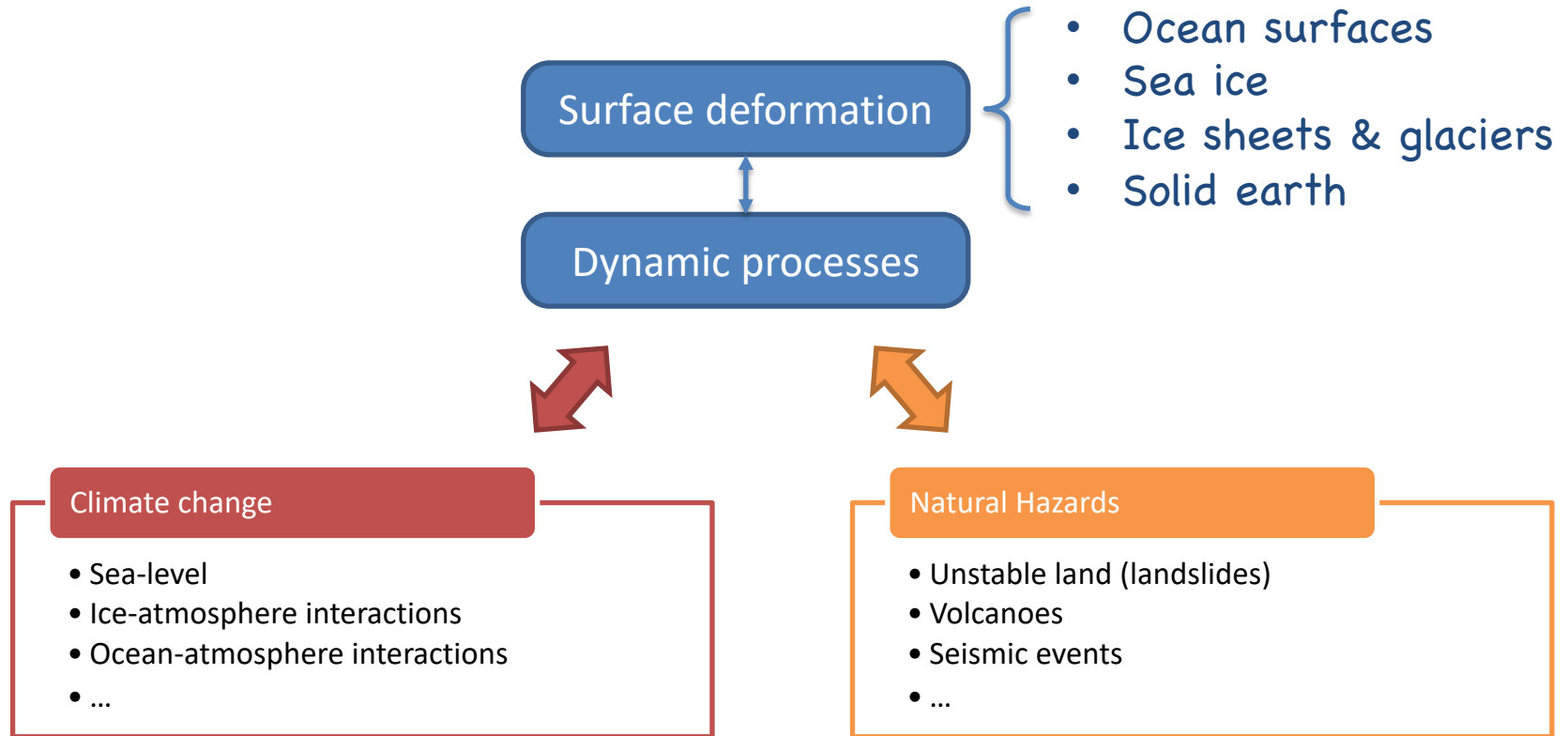
Cryosphere

- Glacier and ice sheets topography/volume/mass change
- High resolution ice flows/deformation
- Sea ice drift and topography
- Marginal Ice Zone variability

Oceans

- High resolution surface currents and wave data for coastal processes
- Small-scale (100 m to 10 km) ocean dynamics
- Surface deformation field ($\nabla \cdot \text{TSCV}$) Divergence/strain, vorticity, shear
- Extreme weather events

Common theme



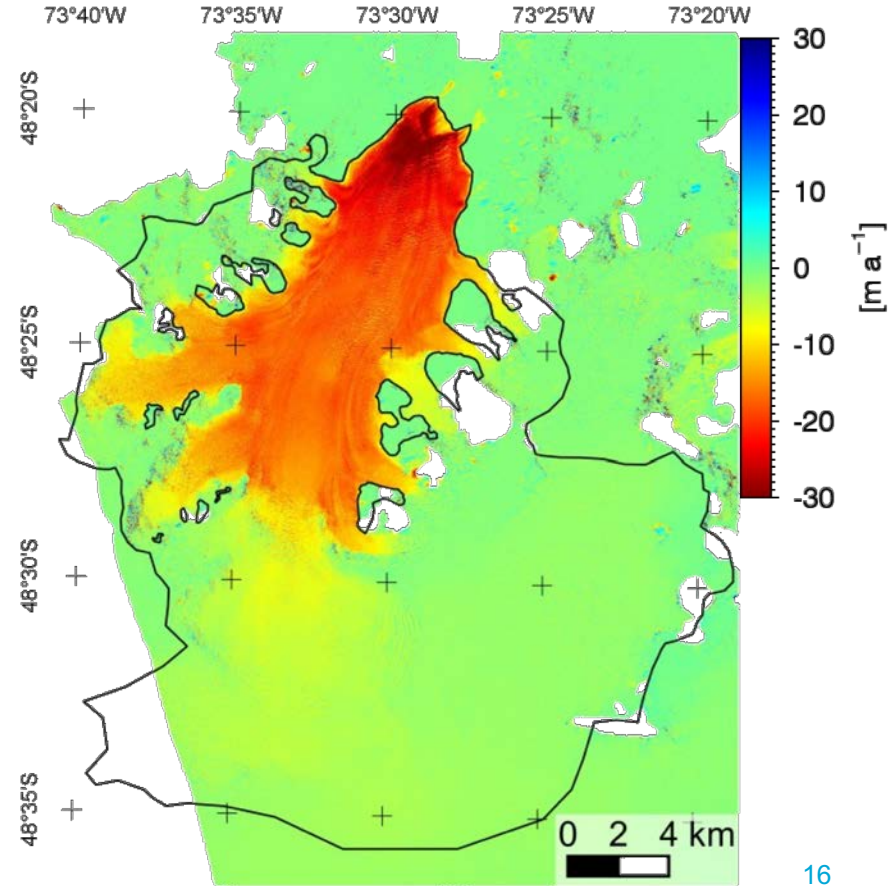
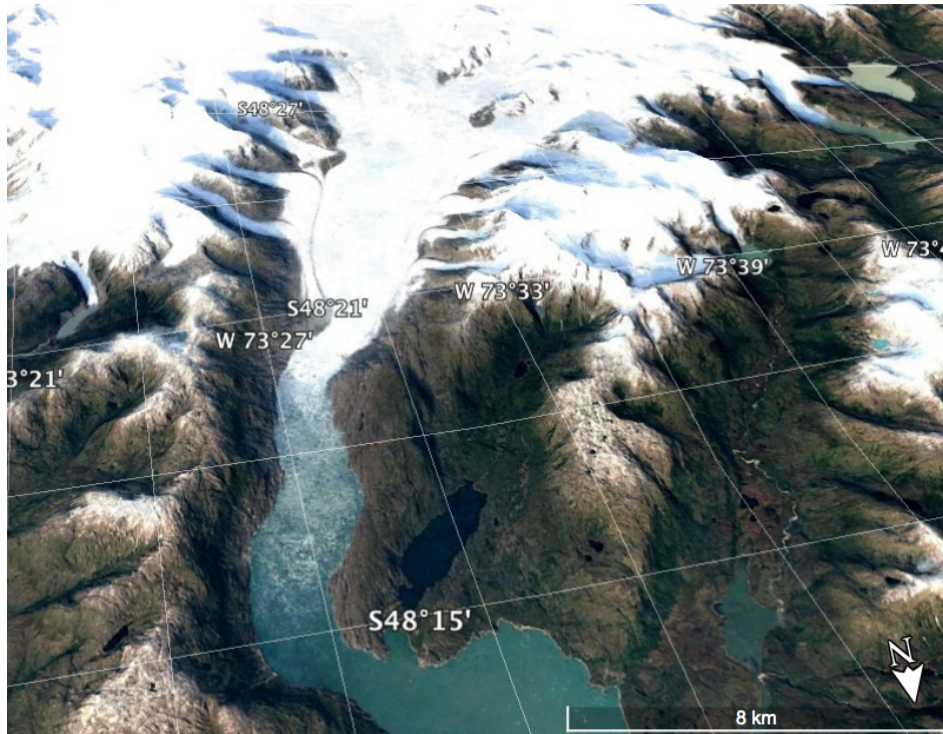
Science objectives: cryosphere

1. Topography change of glaciers, ice caps and outlet glaciers of ice sheets → volume change → contribution to sea level rise
2. 3-D deformation and flow dynamics of ice sheets and glaciers
3. Sea ice
 - Instantaneous drift
 - Thickness and roughness
4. Marginal Ice Zone (MIZ) mesoscale variability

Science objectives: cryosphere (proposal version)

	Short description	ESA challenge
SC-1	Quantify the temporal change of volume and mass of glaciers, ice caps and outlet glaciers of ice sheets and their contribution to sea level change	C2
SC-2	Study the 3-D deformation and flow dynamics of ice sheets and glaciers to better understand their stability and dynamic response to climate change	C1
SC-3	Measure instantaneous sea ice drift velocity and deformation, improve estimations of sea ice roughness and thickness in support of sea ice model development, to improve operational forecast and mass balance studies.	C3
SC-4	Quantify sea ice and Marginal Ice Zone (MIZ) mesoscale variability	C3

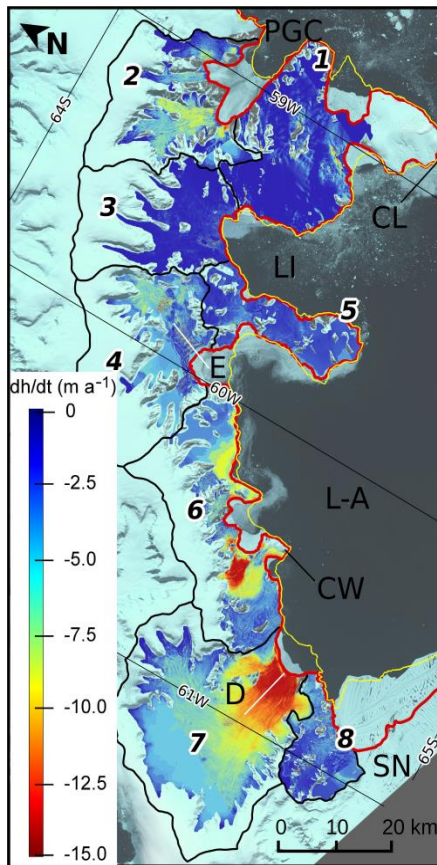
Ice loss in mountain glaciers



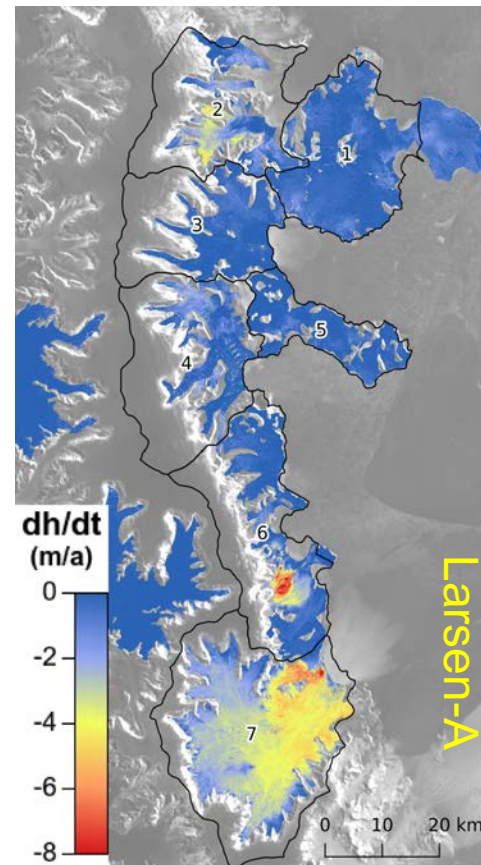
Δ DEMs for mass balance

Annual change in surface elevation (dh/dt) for Antarctic Peninsula outlet glaciers by means of TanDEM-X DEM Differencing

2011 to 2013

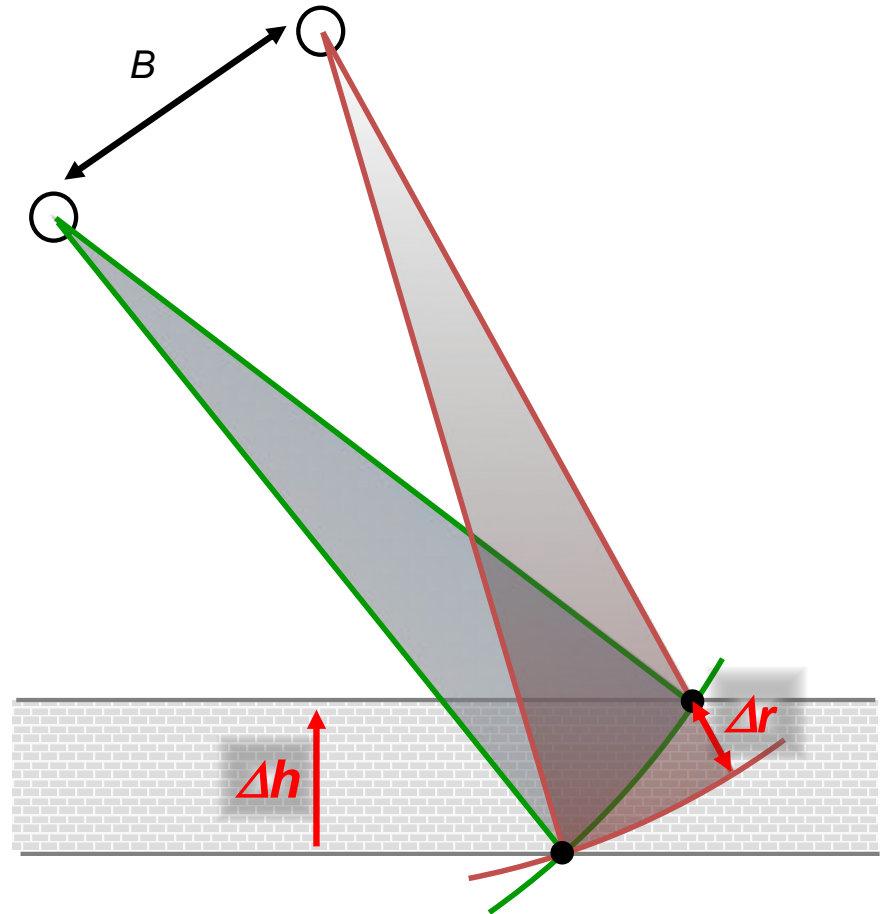


2013 to 2016



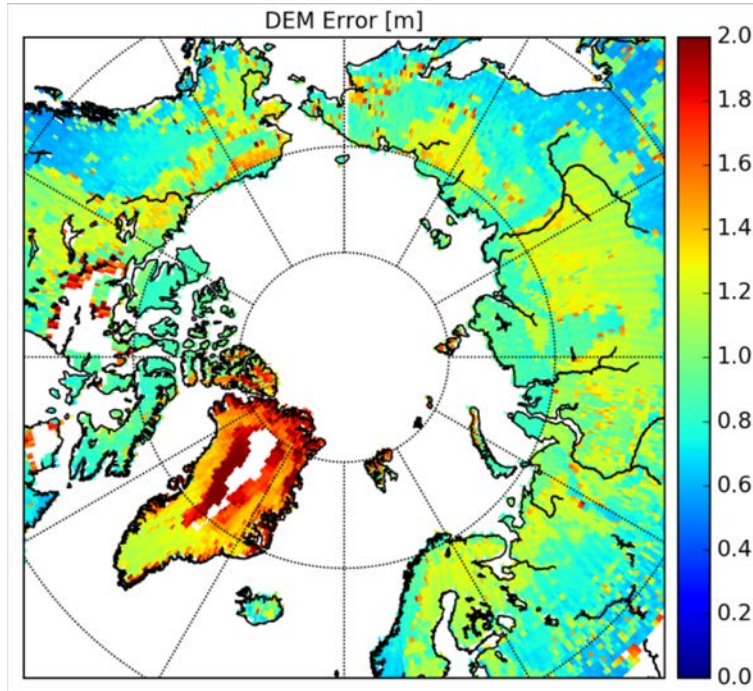
Measurement principle

- Single-pass cross-track interferometry
 - No temporal decorrelation
 - Virtually no atmospheric errors
 - Large and reconfigurable baselines
- Challenges/limits
 - Baseline changes along orbit
 - Precise baseline knowledge
 - System-level phase errors: synchronization
 - Penetration depth

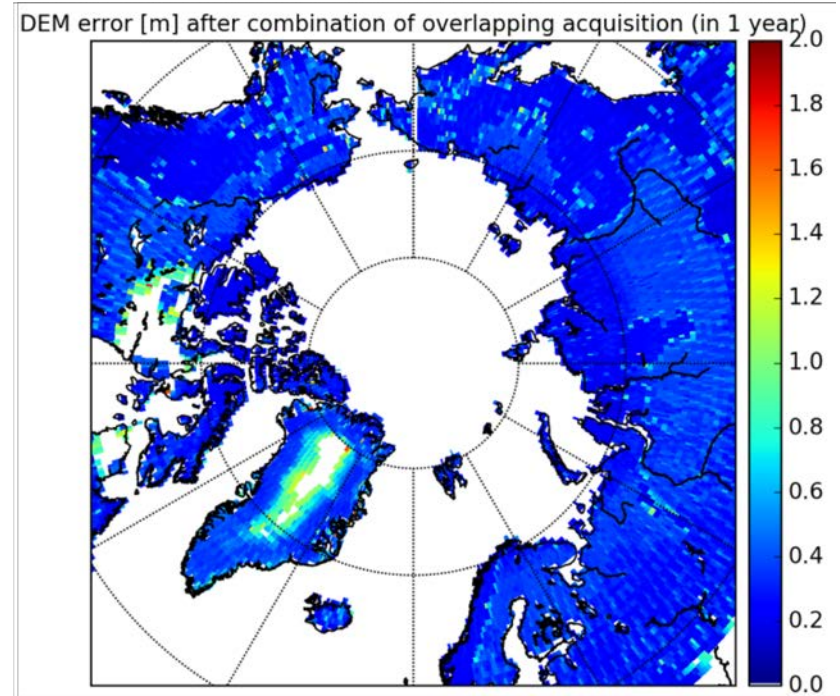


DEM performance example: SESAME, 100m resolution

Single Acquisition



Yearly average



Main open issue (remote sensing wise): penetration depth

Ice is semi-transparent

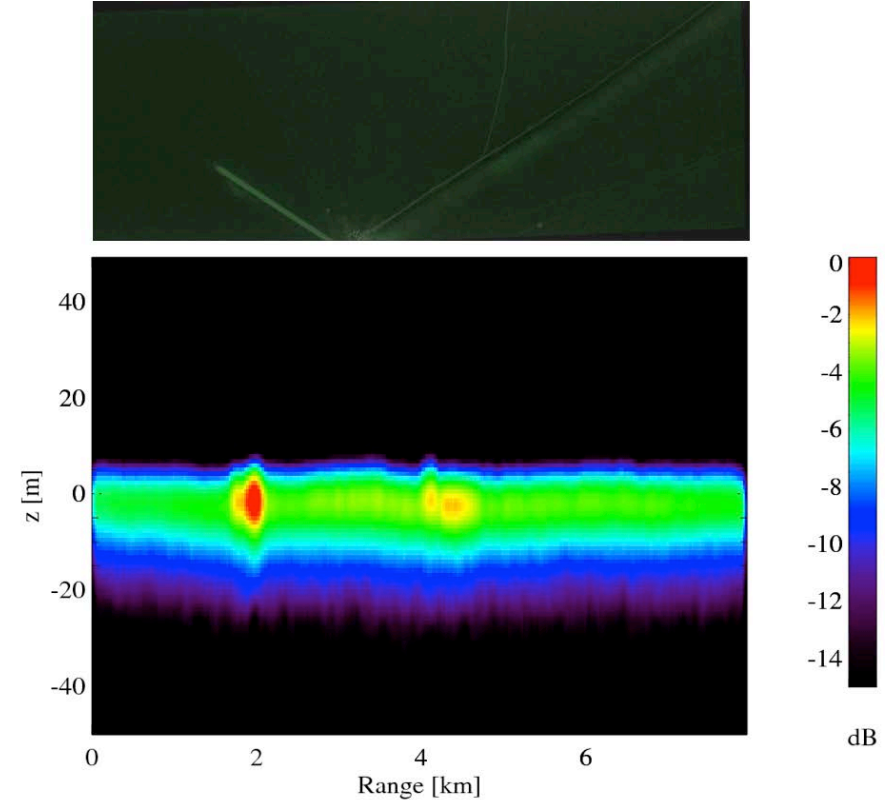


InSAR phase-center height below ice surface



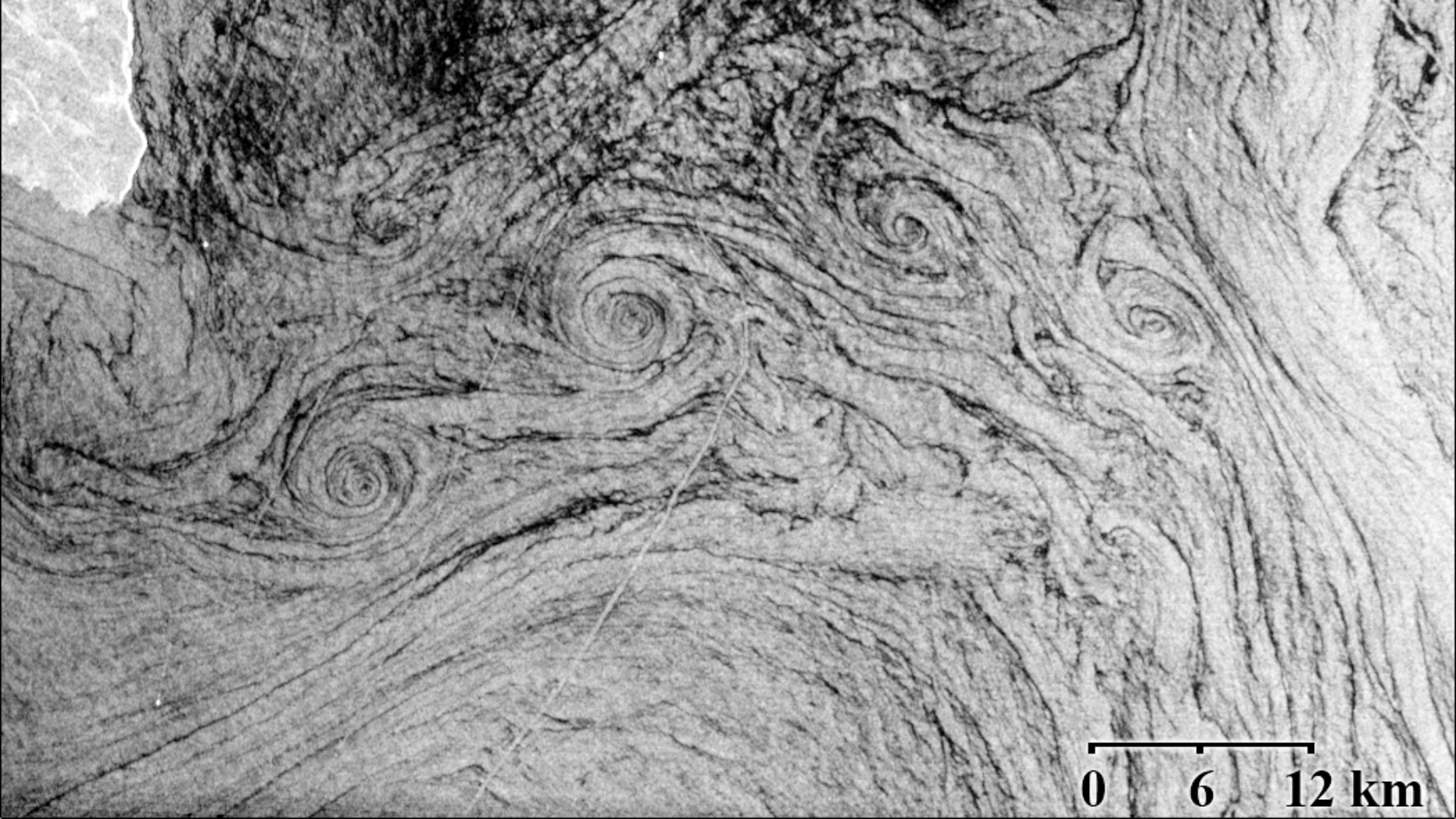
Problem if penetration depth varies

TSX tomographic image over lake Vostok



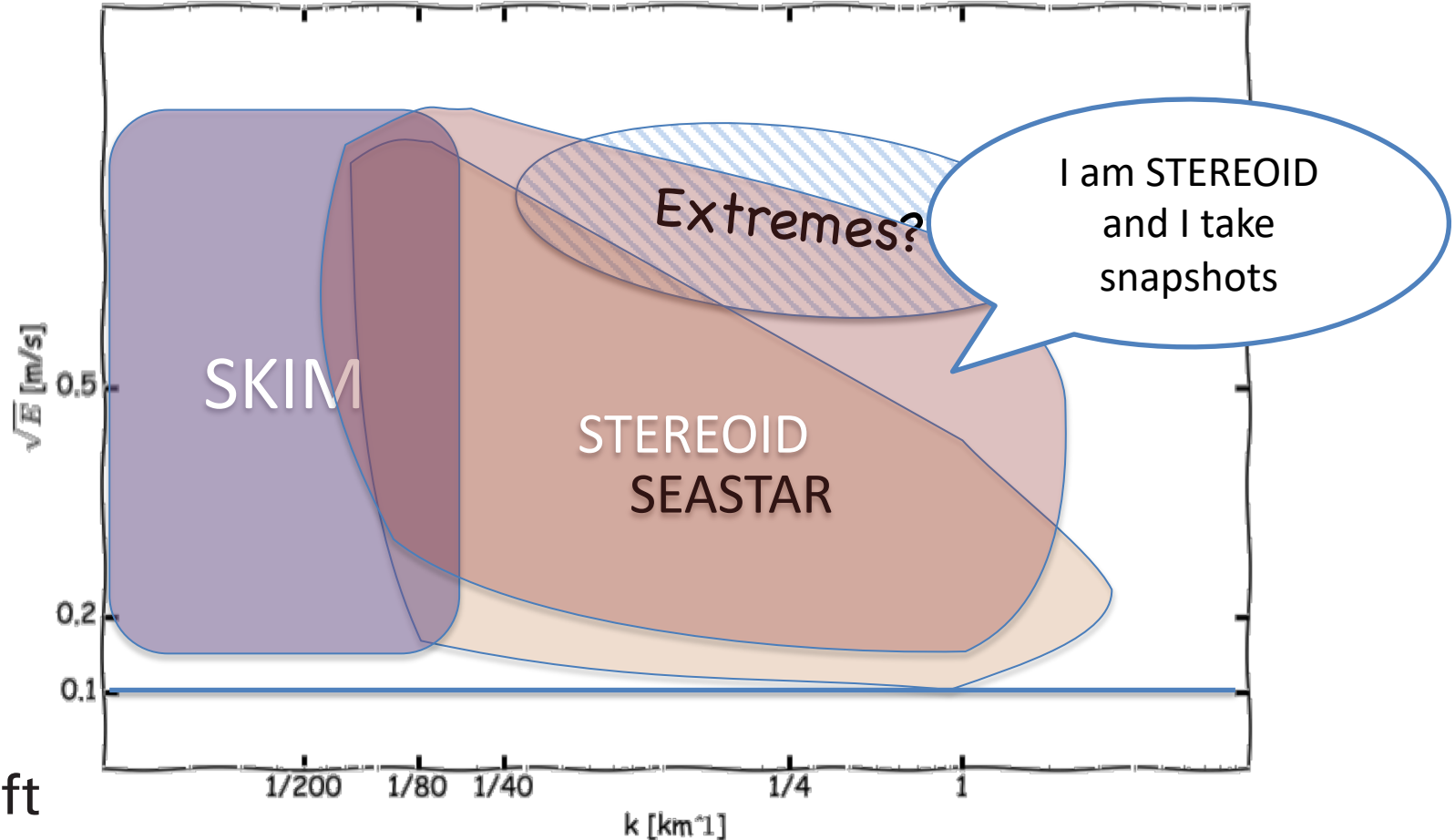
Science objectives: oceans

1. High resolution surface currents and wave data to validate and develop models for coastal processes
2. Characterize and quantify small-scale (100 m to 10 km) ocean dynamics
3. Measure the surface deformation field (gradients TSCV)
 - Divergence/strain
 - Vorticity
 - Shear
4. Improve our understanding of air-sea interactions
5. Extreme weather events

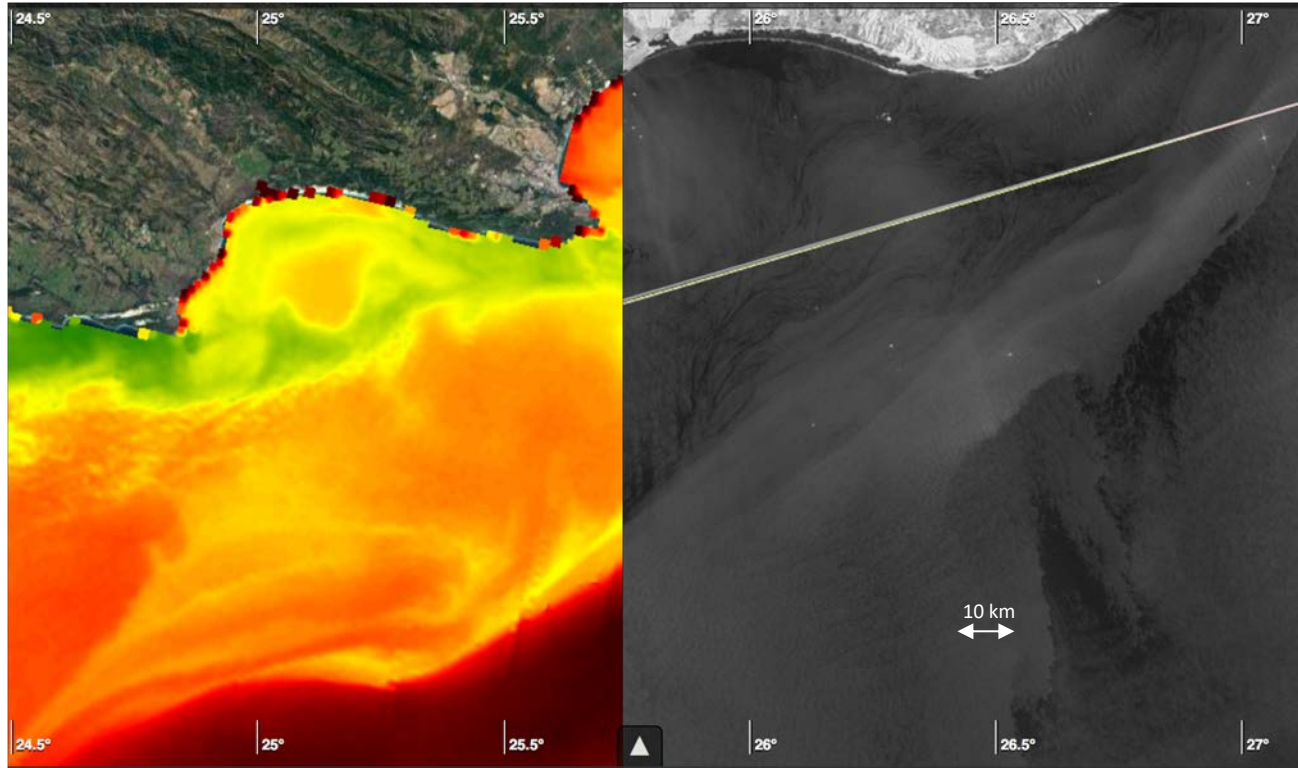


0 6 12 km

And attempt to contextualize STEREOID (TSCCV)

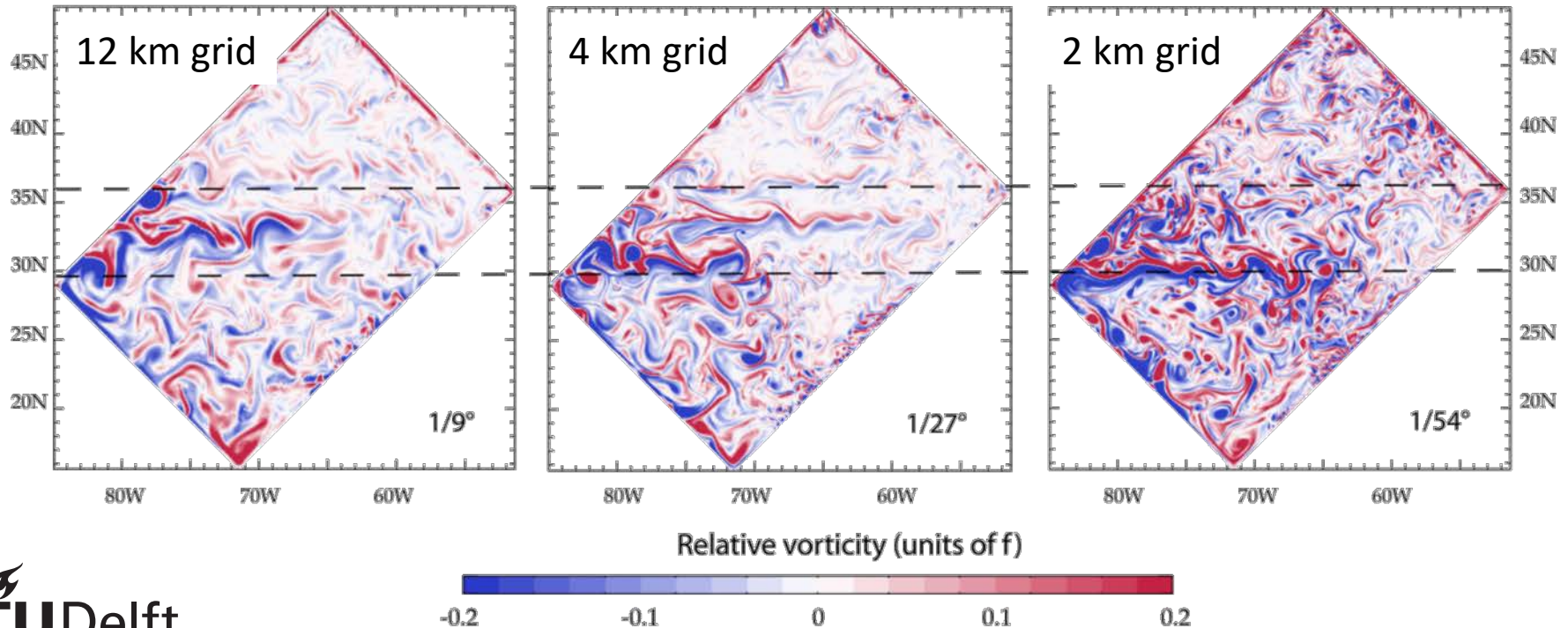


TIR + radar rationale

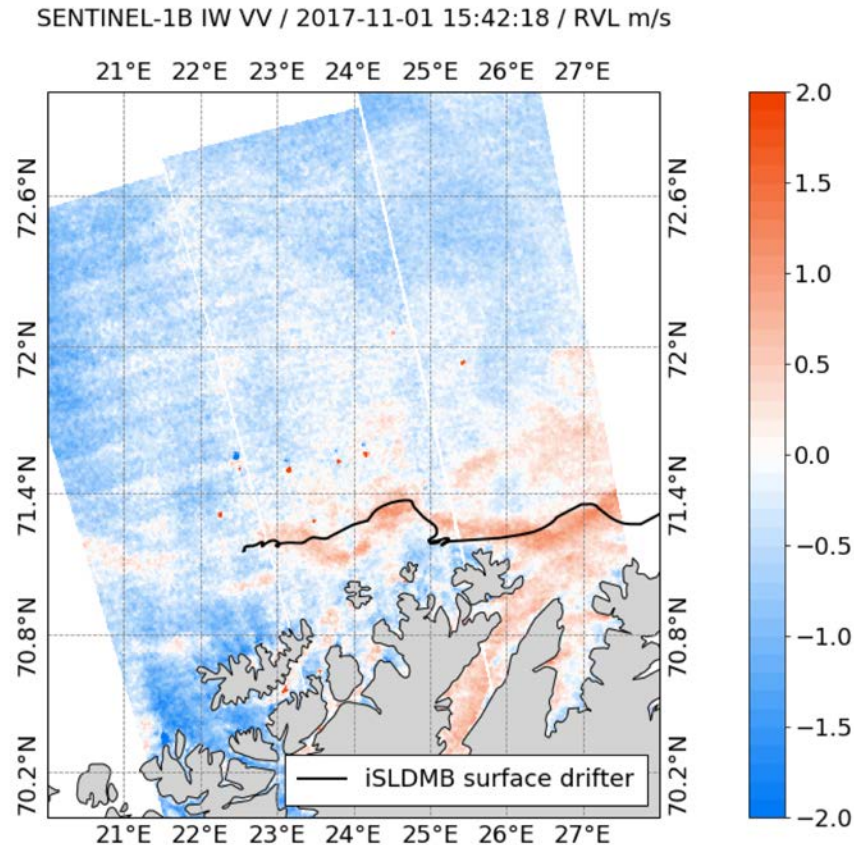


The value of resolution (for example, vorticity)

M. Lévy et al./Ocean Modelling 34 (2010) 1–15



Radial velocities measured by Sentinel-1



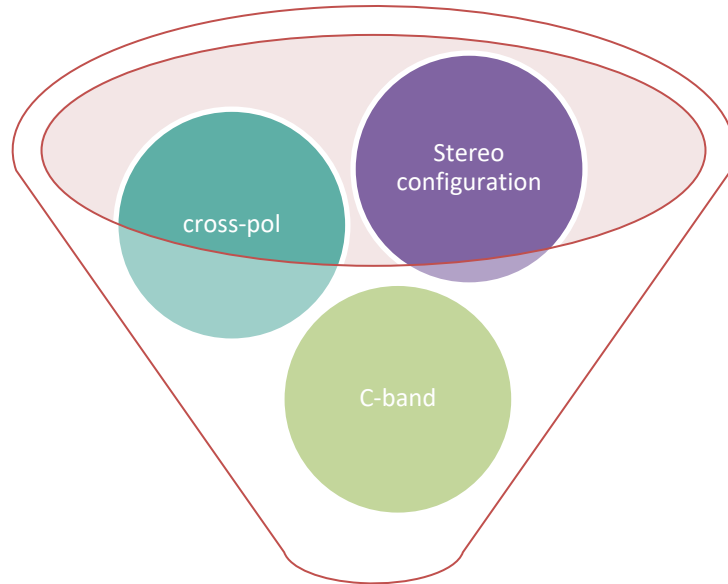
Main open issue: (wind) wave bias

- Problem:

$$v_{\text{Doppler}} = v_{\text{Wave-Bias}} + v_{\text{TSC}}$$

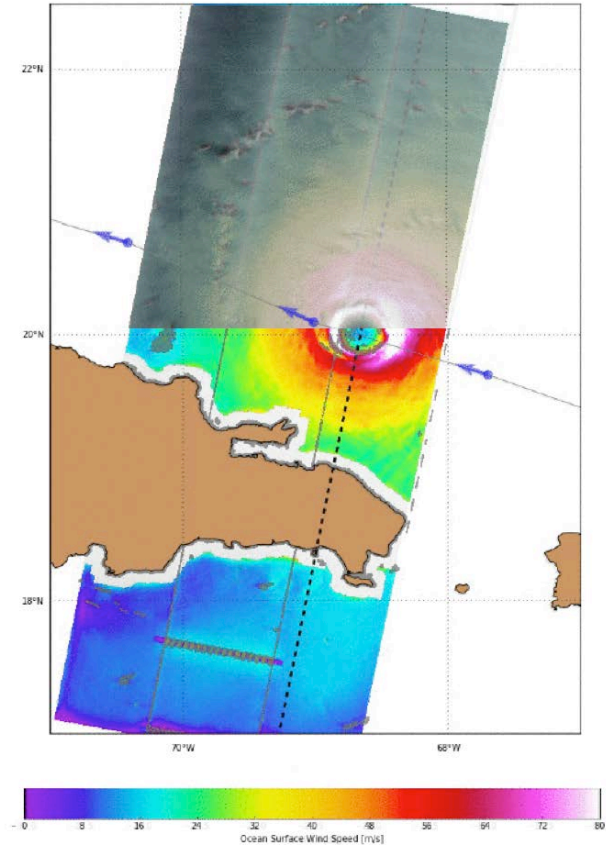
- Approaches being studied
 - Measure waves → estimate bias
 - Estimate wind from backscatter → model waves → estimate bias
 - Exploit polarimetric dependency of wave-bias

Extreme weather [Stereo]



Surface winds +
TSCV

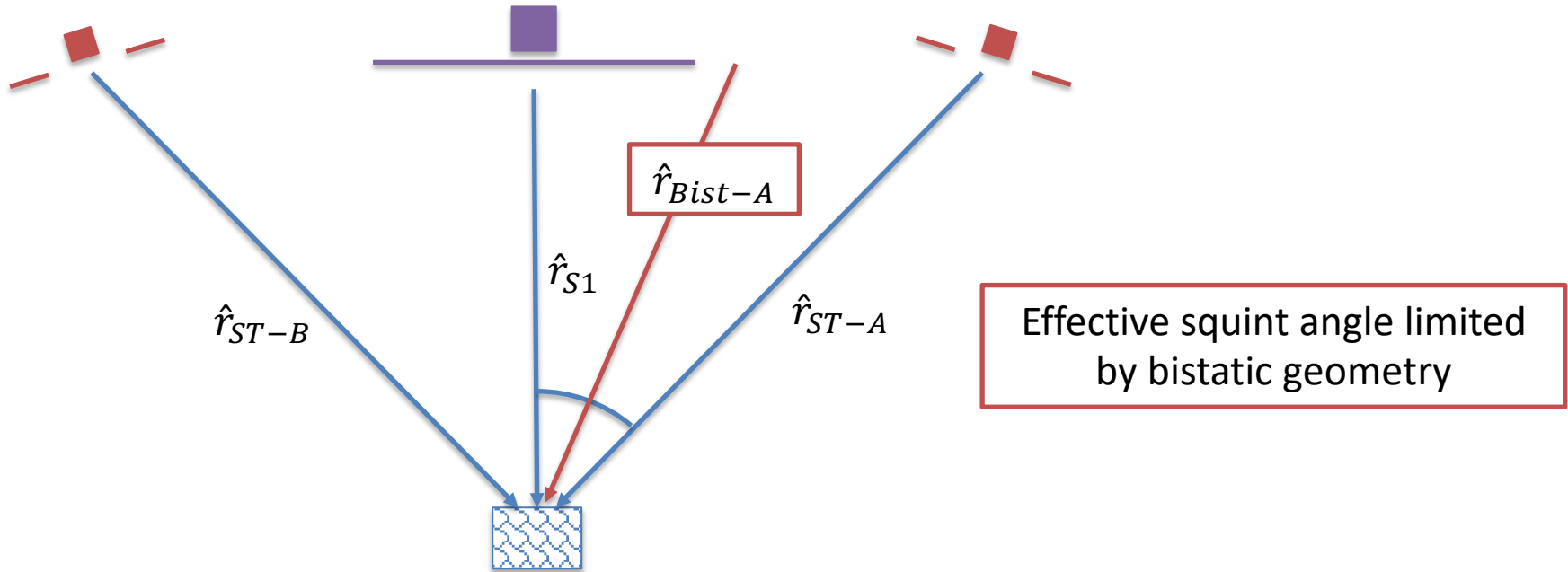
#IrmaHurricane2017



Sentinel-1

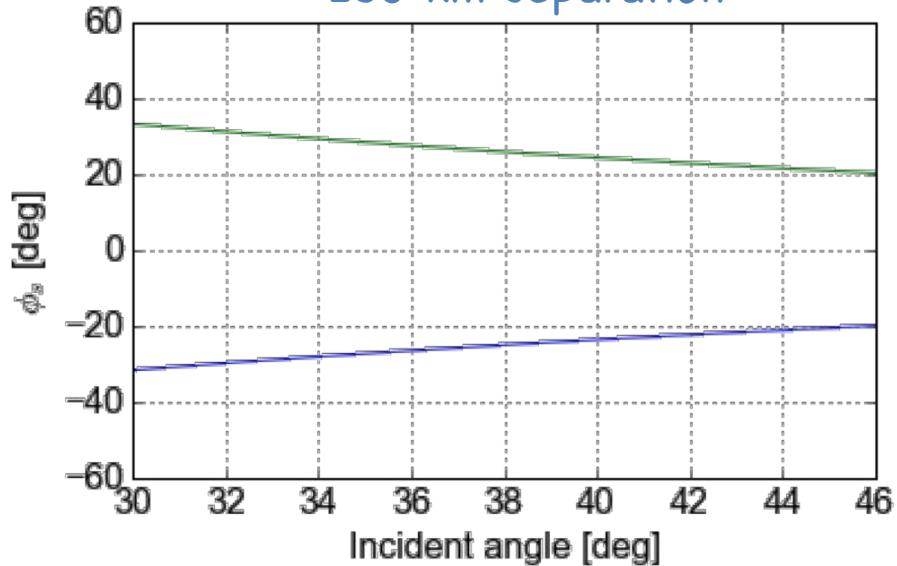
Contains modified Copernicus
Sentinel data (2017)

Squint angles (top view)

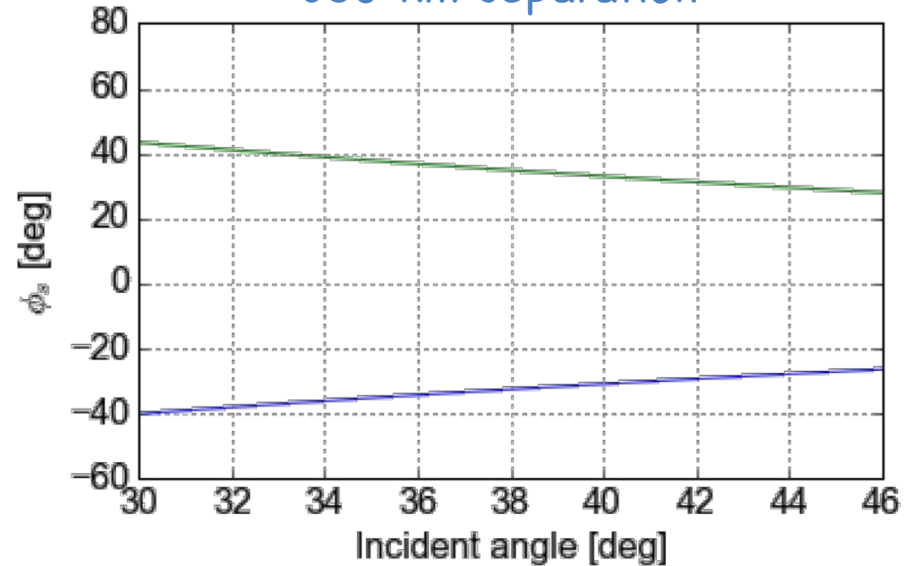


Squint angles

250 km separation



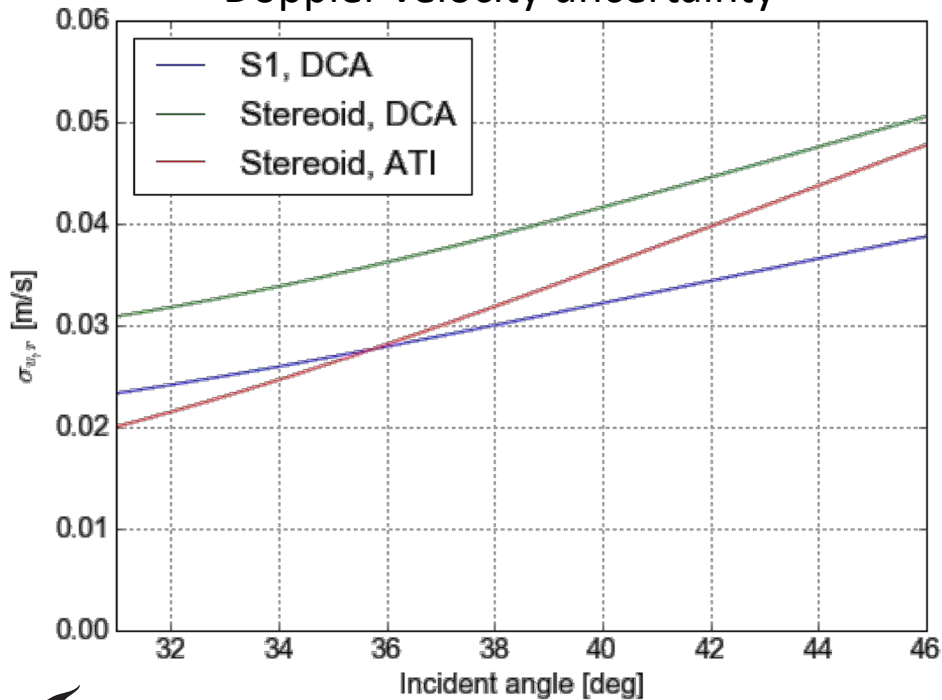
350 km separation



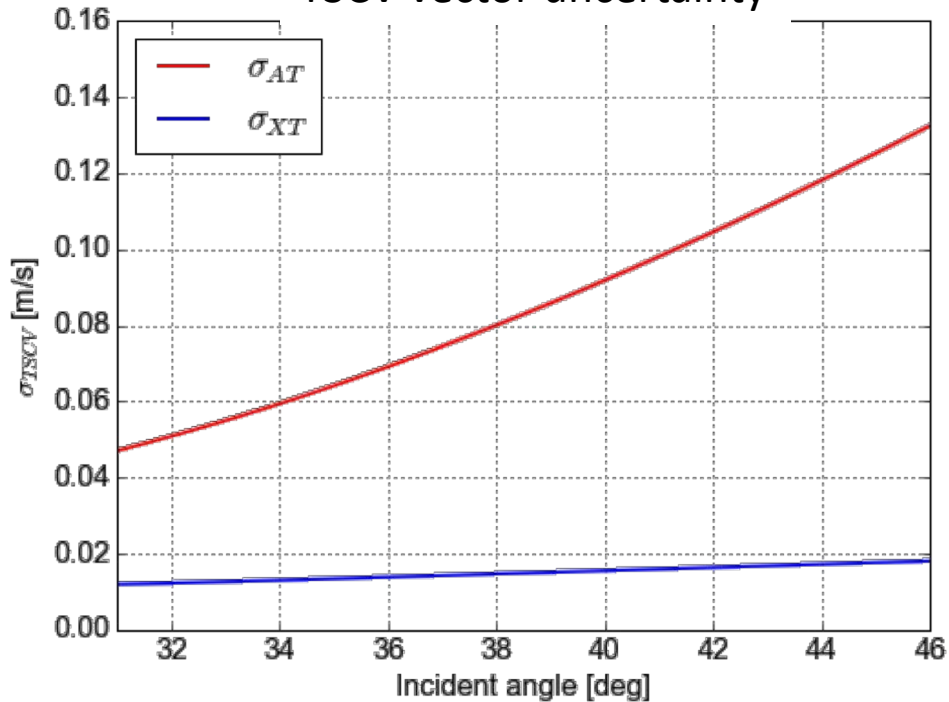
System Sensitivity (ideal retrieval)

- 3 km resolution
- 6 m/s wind
- 250 km separation

Doppler velocity uncertainty



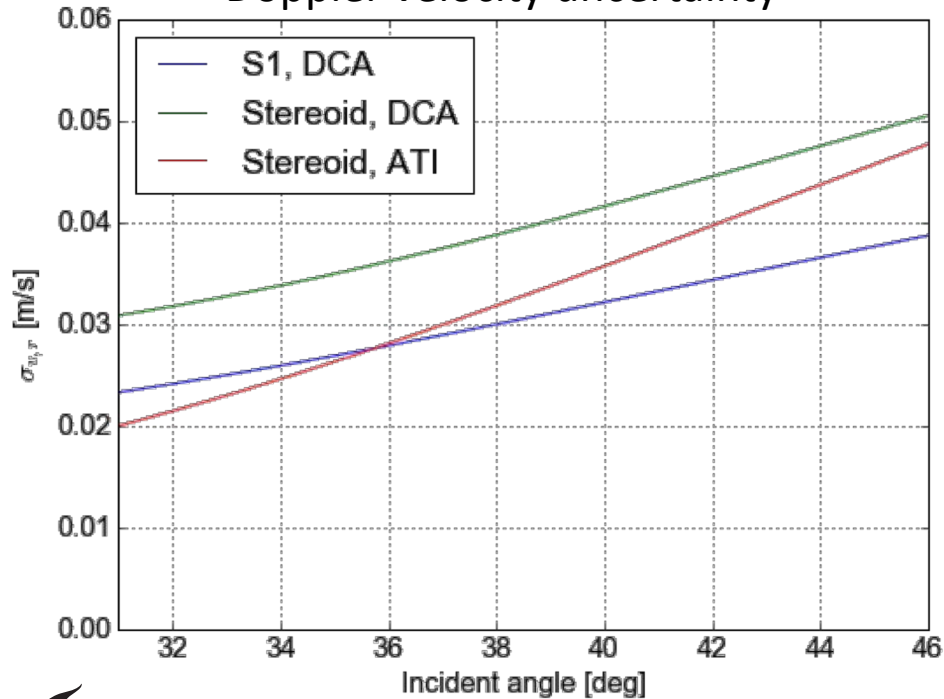
TSCV vector uncertainty



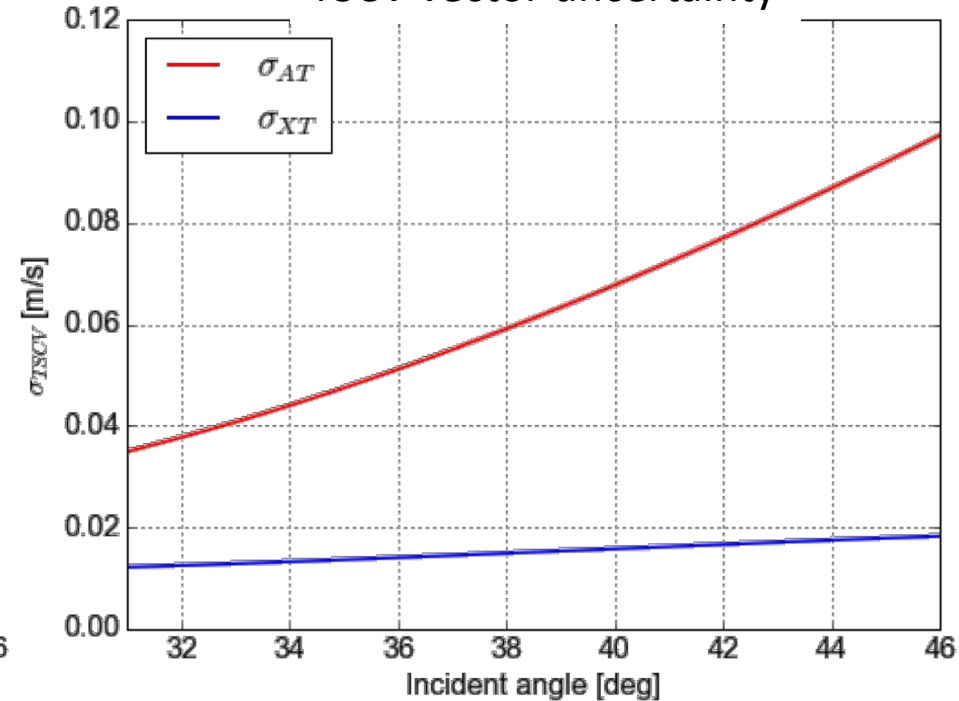
System Sensitivity (ideal retrieval)

- 3 km resolution
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- **350 km separation**

Doppler velocity uncertainty



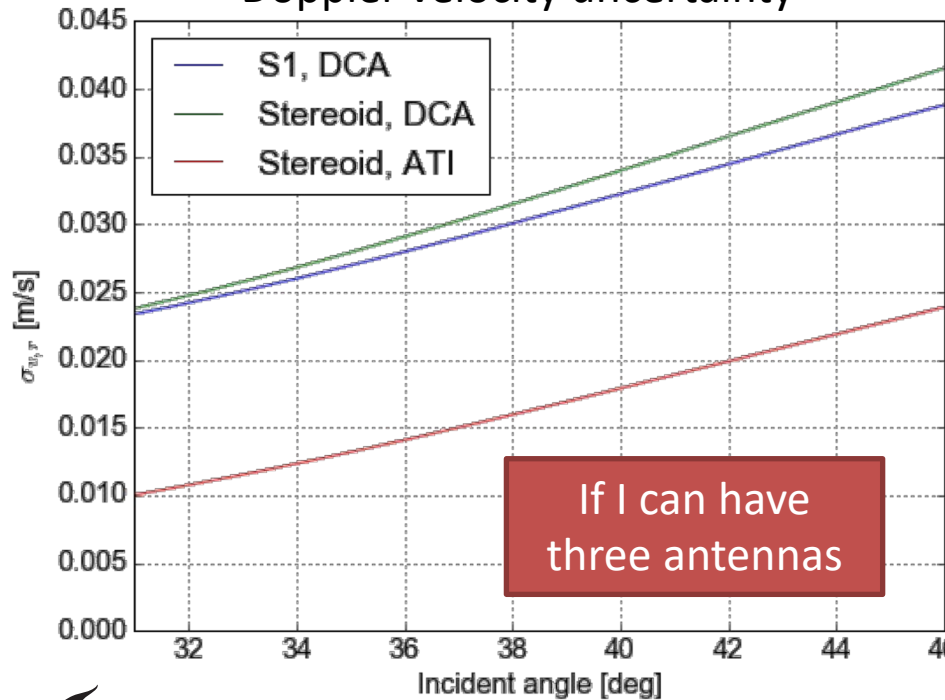
TSCV vector uncertainty



System Sensitivity (ideal retrieval)

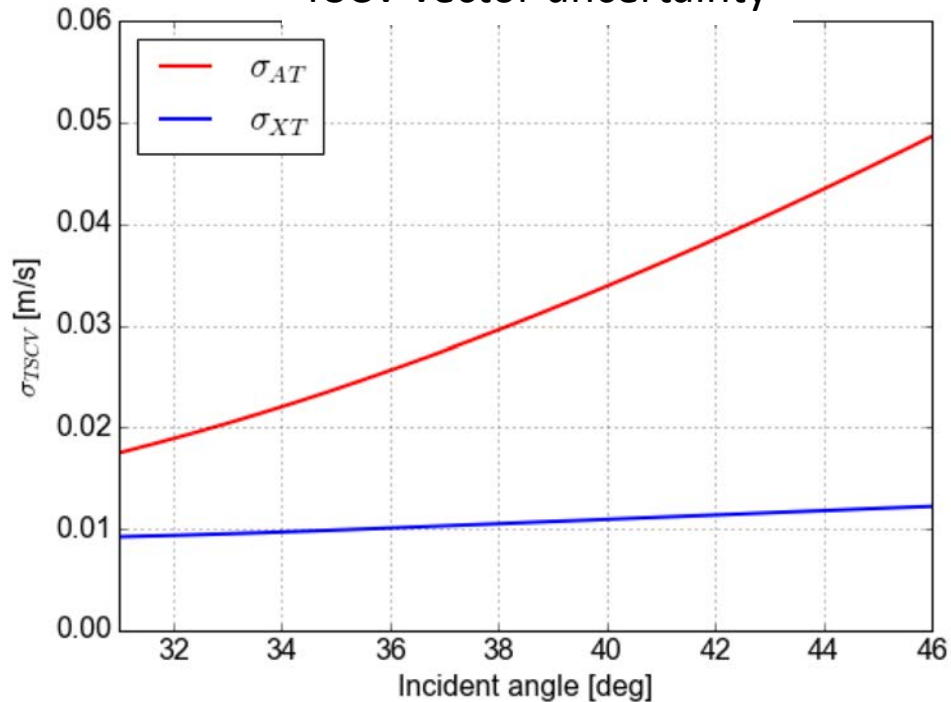
- 3 km resolution
- 6 m/s wind
- **350 km separation**

Doppler velocity uncertainty



If I can have
three antennas

TSCV vector uncertainty



Main sources of systematic (non-geophysical) errors

	DCA	Short-ATI	Long-ATI
Sentinel-1 pointing	Mispointings weighted by $1/\text{beamwidths}$		
Companion mispointing			
Formation knowledge			Leads to ATI phase offset
Oscillator frequency offsets	Is a point of concern, but seems technically solved		

But emphasis on gradients



We can mostly live with low-pass systematic errors

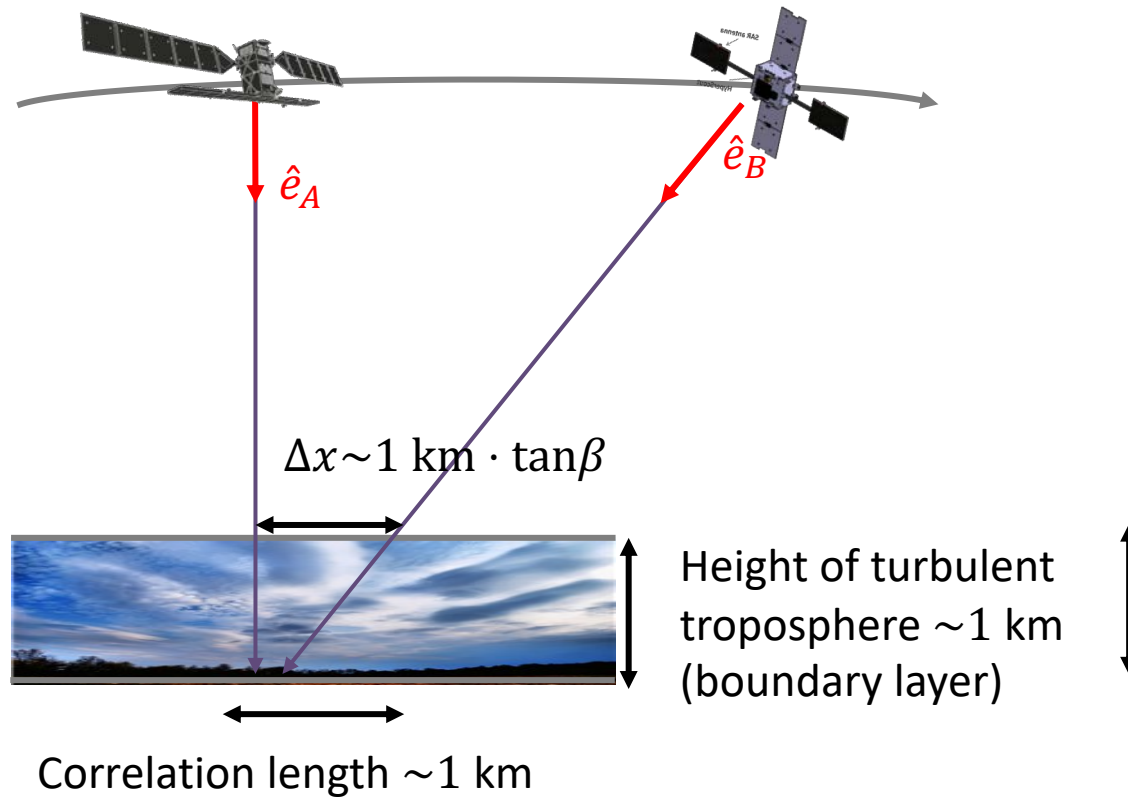
Science objectives: solid Earth

1. Monitor 3D surface

- co-seismic, inter-seismic, and post-seismic deformation;
- tectonic rifting;
- volcanism;
- landslides.

2. Topographic changes (volume!) due to volcanic and landslide activity

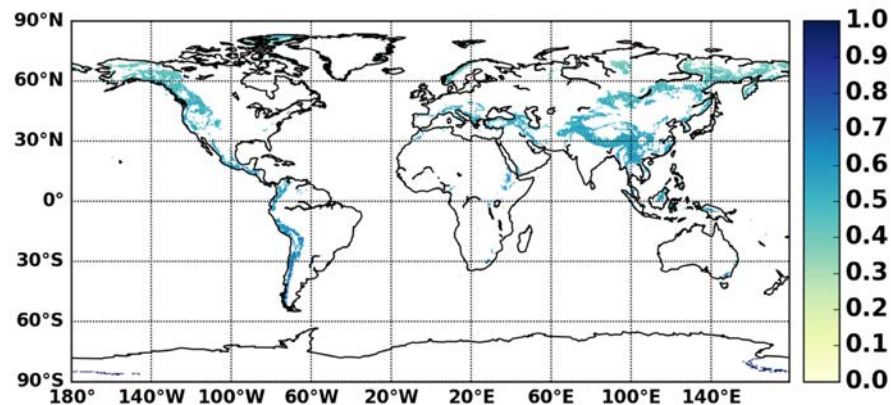
Geometry-diverse DInSAR



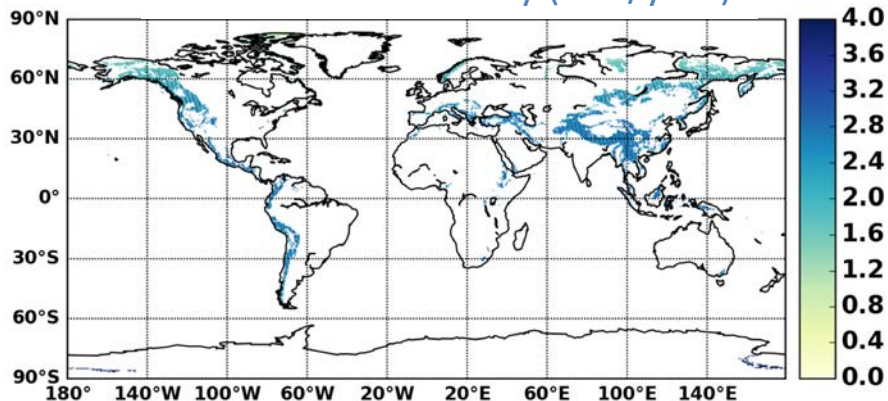
3D deformation

- Along track separation of 250 km
- Performance over 5 years
- Product resolution 100m x 100m
- $\tau = 40\text{days}$, $\gamma_0 = 1$, $\gamma_\infty = 0.15$
- Troposphere power 1 cm²
- Only Sentinel-1 A (12 days repeat)

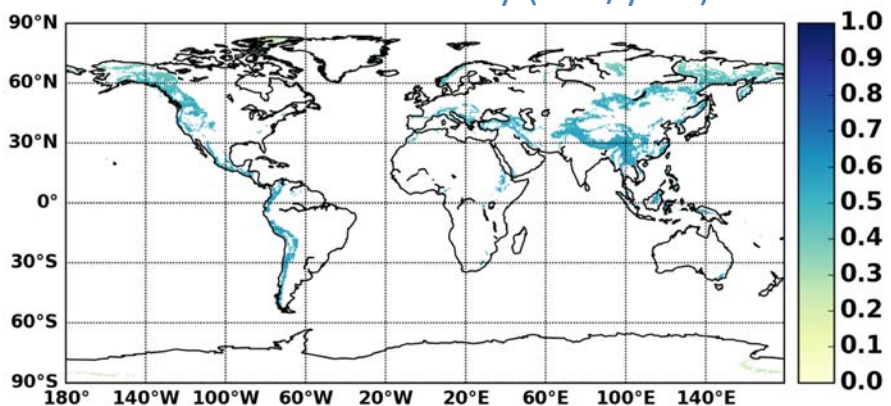
Up-Down accuracy (mm/year)



North-South accuracy (mm/year)

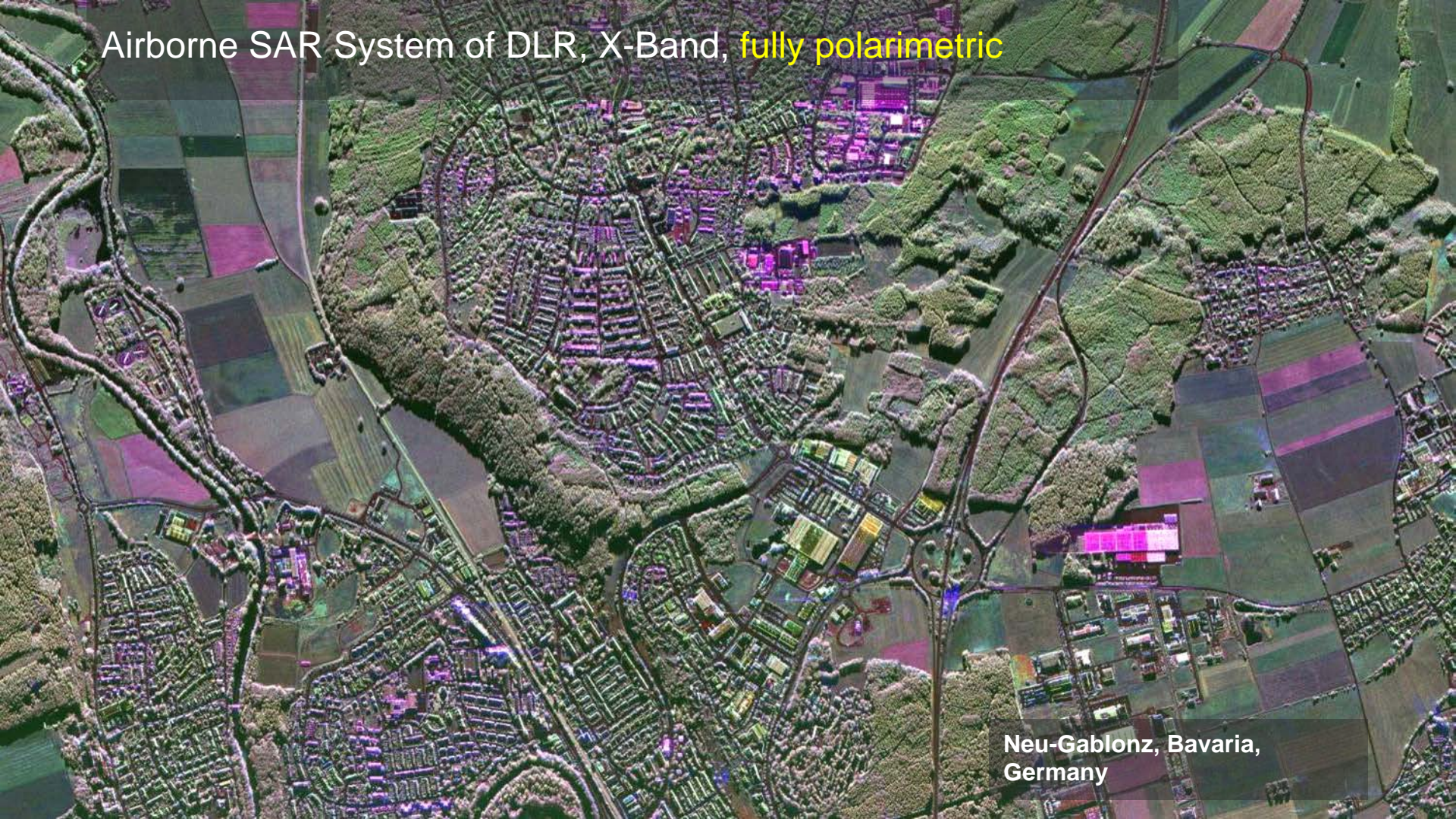


East-West accuracy (mm/year)



From polarimetry to...
multistatimetry?

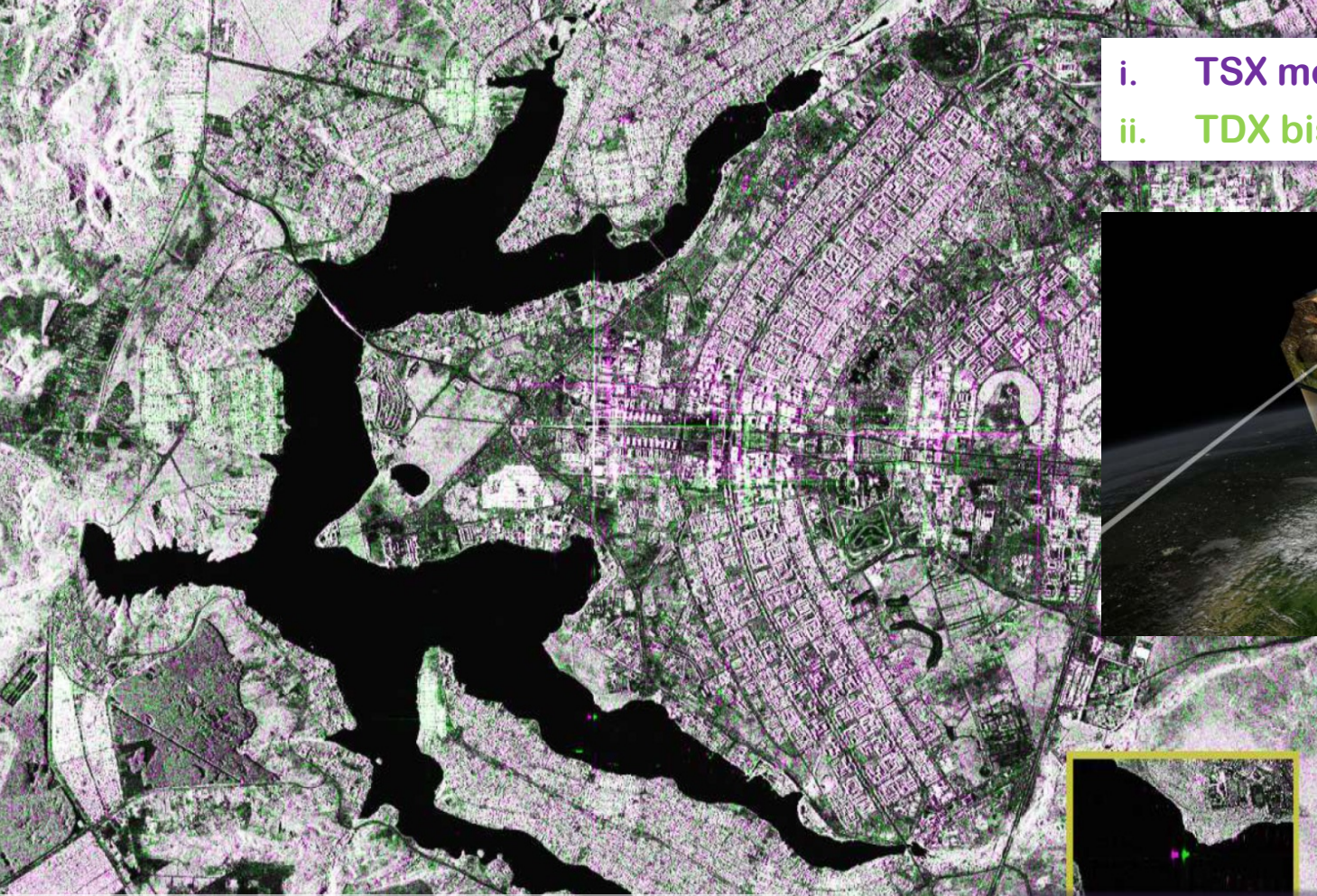
Airborne SAR System of DLR, X-Band, fully polarimetric



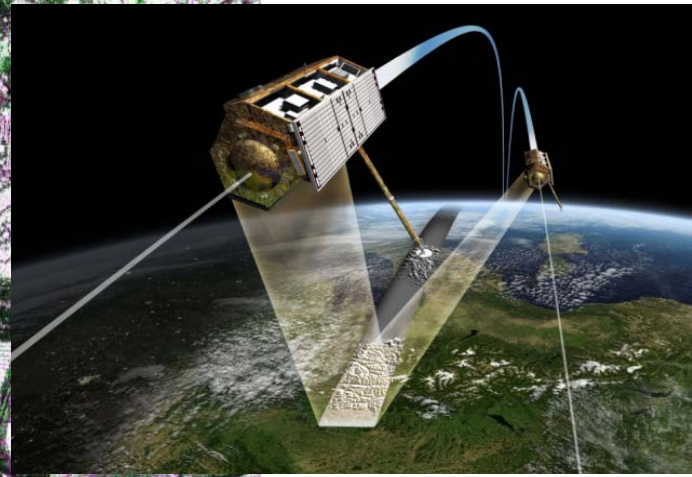
Neu-Gablonz, Bavaria,
Germany

Very limited empirical experience with multistatic data





- i. TSX monostatic image
- ii. TDX bistatic image (20 km)



TanDEM-X: the 1st bistatic SAR in space

Brasilia city, Brazil, August 8th 2010

Space-segment side challenges

Formation flying

Synchronization

SAR
performance

Formation
safety

Cross-track
baseline
knowledge

Common
Doppler and
along-track
baseline
control

Phase

Echo
window

Sensitivity

Ambiguities

Formation flying: cross-track baseline knowledge

- LOS baseline error translates directly into phase error:

$$\Delta\phi = 2\pi \cdot \frac{\Delta B_{\parallel}}{\lambda_0}$$

$$\epsilon_h = R \cdot \sin \theta_{inc} \cdot \frac{\Delta B_{\parallel}}{B_{\perp}}$$

Baseline rotation

- $R \sim 800$ km
- $B_{\perp} \sim 200$ m to 400 m
- Target $\epsilon_h \sim 10$ cm

Baseline knowledge requirement from $O(0.1$ mm)

Common Doppler loss

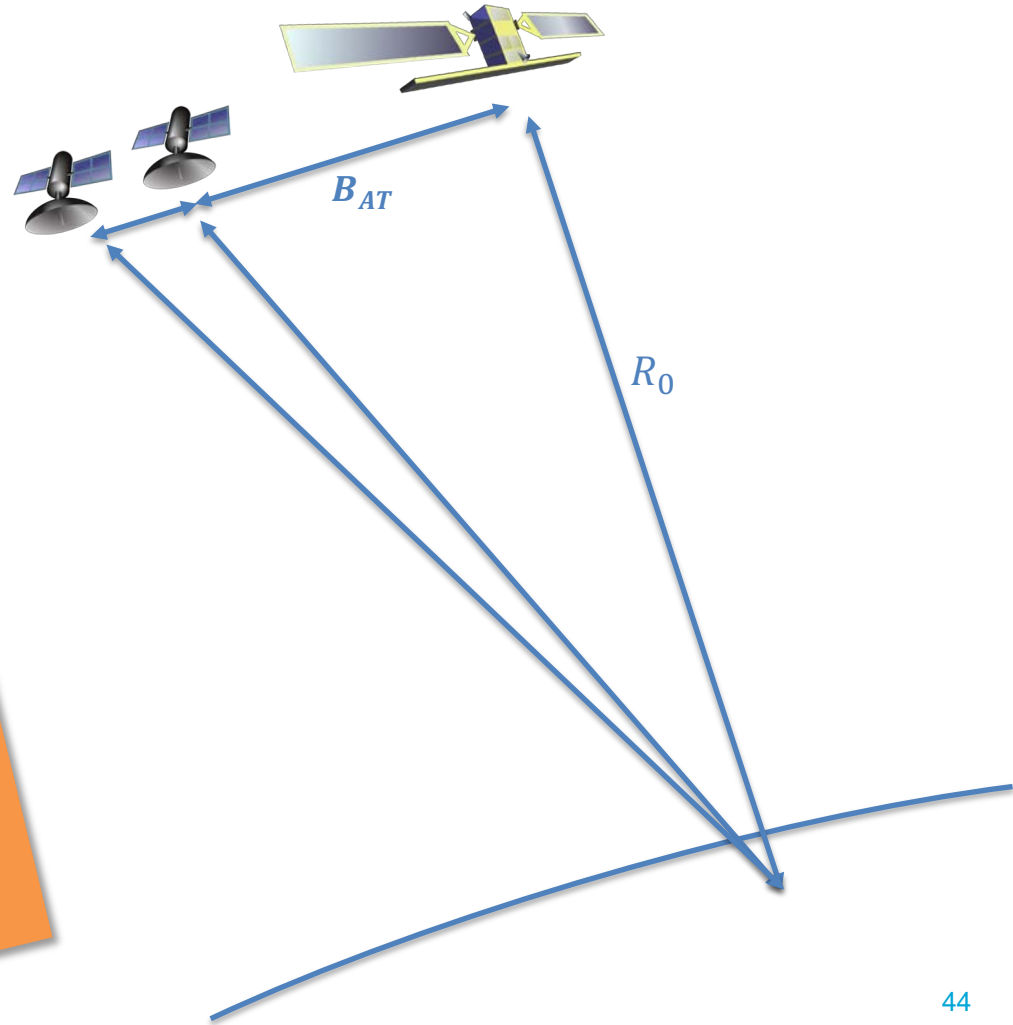
$$\Delta f_{\text{Dop}} = \frac{v_{\text{orb}}}{\lambda} \cdot \frac{B_{AT}}{R_0}$$

Should be small compared to processed Doppler bandwidth.



$$B_{AT} \ll \frac{\lambda \cdot R_0}{v_{\text{orb}}} \cdot B_{\text{proc}}$$

For STEREOID
0(300 m)
separation
OK



Space-segment side challenges

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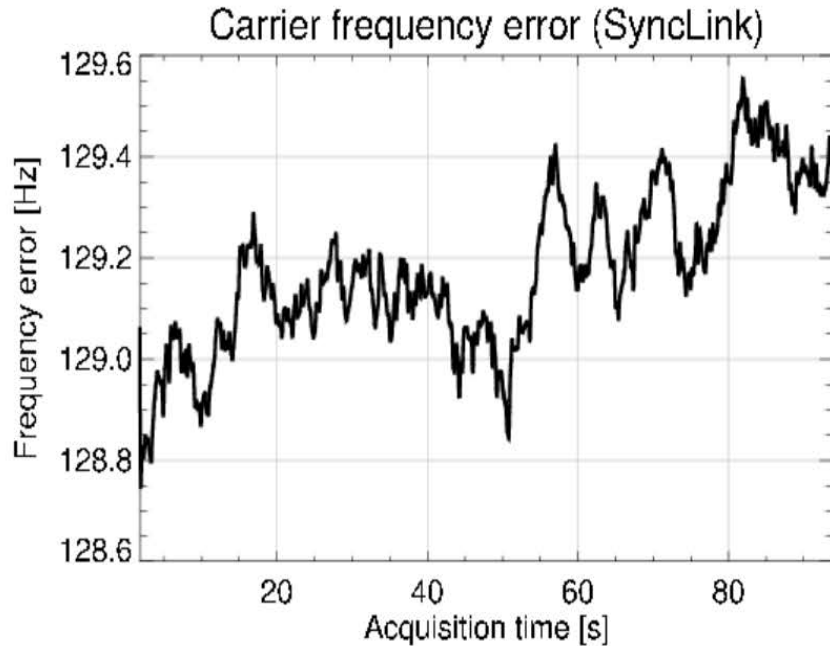
Phase

Echo
window

Sensitivity

Ambiguities

Phase synchronization



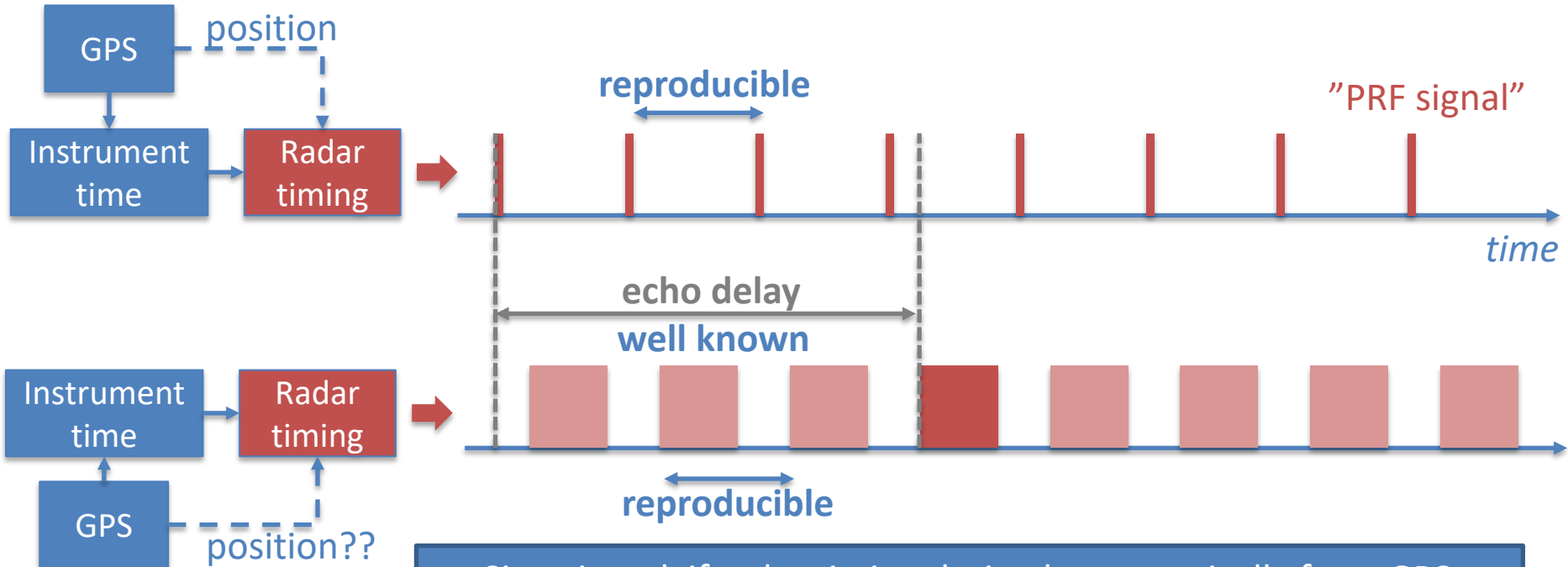
From TanDEM-X SyncLink

Frequency and phase synchronization
always a critical issue.

Carrier frequency and phase synchronization: lessons we think we have learnt

- GPS tagging/disciplining → Frequency offsets
- Data driven (AutoSync, etc) → Relative phase errors
 - Often good enough
 - Issues for sure
- Explicit synchronization link
 - Two way synchronization between receivers needed

Echo window synchronization



- Situation ok if radar timing derived systematically from GPS-referenced instrument time
- Position derived timing potential trouble maker.

Technical challenges

Formation flying

Synchronization

SAR
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Cross-track
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Common
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control

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Sensitivity

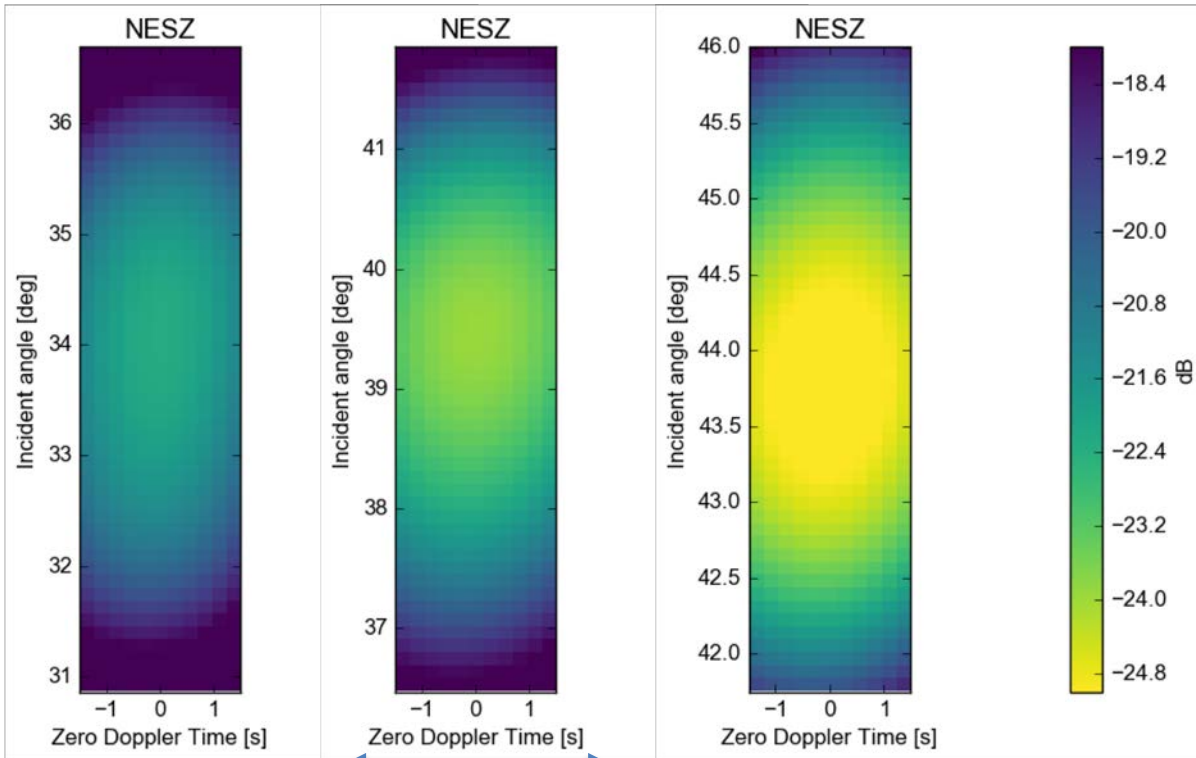
Ambiguities

Imaging performance: NESZ (Imaging mode)

sub-swath 1

sub-swath 2

sub-swath 3



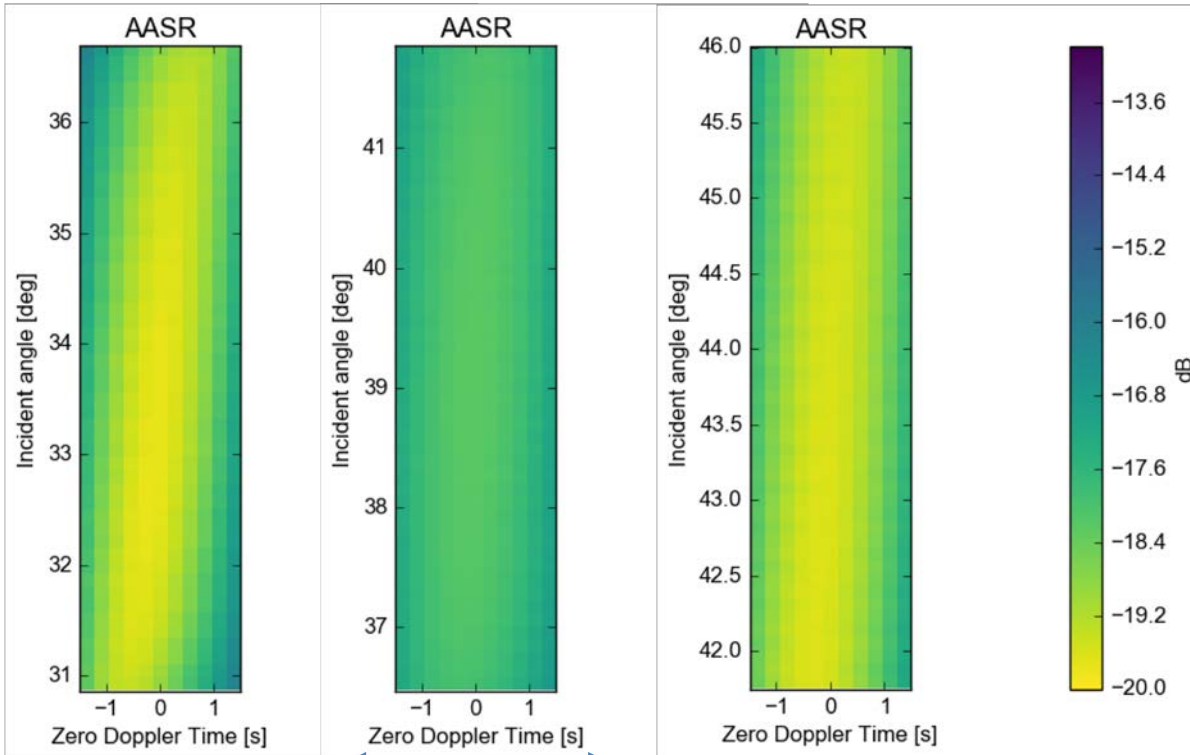
- NESZ generally adequate to good
- Probably a bit less gain in exchange of wider elevation beams would be better.
- Or SCORE
- + 3dB for ATI mode

Imaging performance: AASR (imaging mode)

sub-swath 1

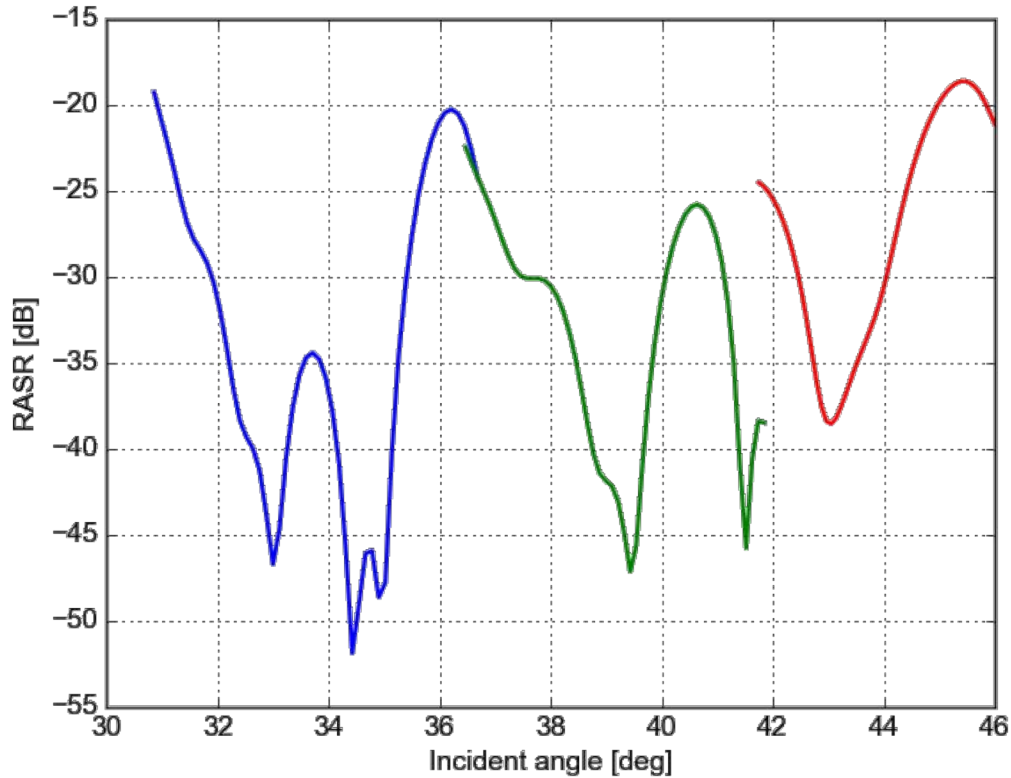
sub-swath 2

sub-swath 3



- AASR in -17 to -18 dB range.
- Quite good given small total antenna area
- Sub-swath variability due to Sentinel-1 PRFs

Imaging performance: RASR



- RASR < -20 dB
- Good, but we need to accommodate large dynamic ranges (varying wind conditions)
- Misalignment between pattern footprint and iso-Doppler lines



Range ambiguities Doppler shifted (need to account for range-azimuth ambiguities)

Outlook (1/2)

- We have a (very strong) MAG
 - “UK” members: Juliet Biggs (U. Bristol & COMET), Andy Hooper (U. Leeds & COMET), Julianne Stroeve (UCL).
 - First meeting in January
- ESA in process of preparing Phase-0/A ITTs
 - Preliminary MRD and SoW being discussed just now at ESA.

Outlook (2/2)

- No technical show-stoppers up to know
- We are sure that we would produce good science
- Some level of scientific prioritization needs to happen
 - Multi-purpose nature both a strength and a weakness
 - For selection we need compelling but also simple story line.
- We need to formulate & justify mission requirements

