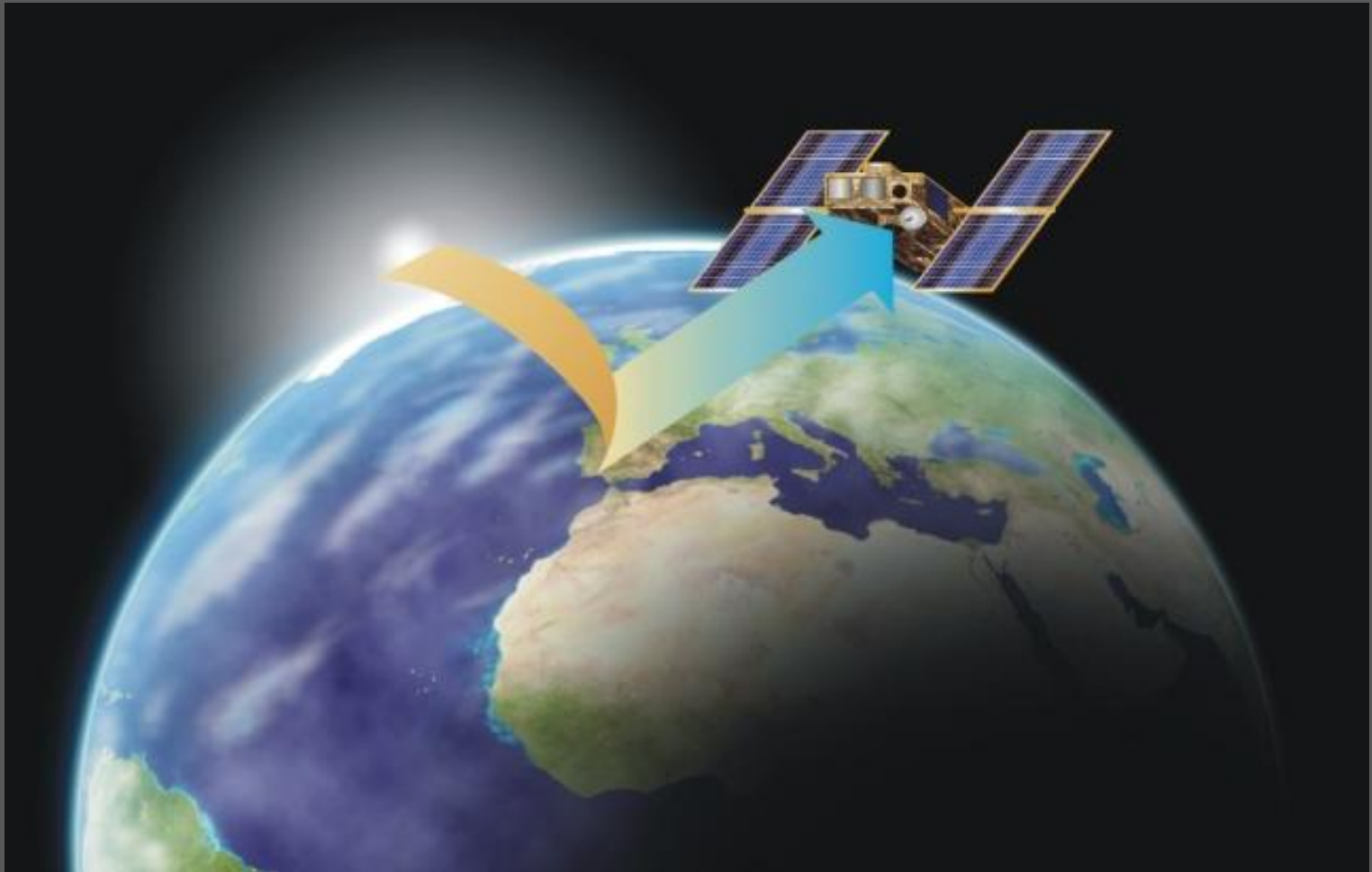


Introduction to UV-Vis Spectrometry Techniques From Space

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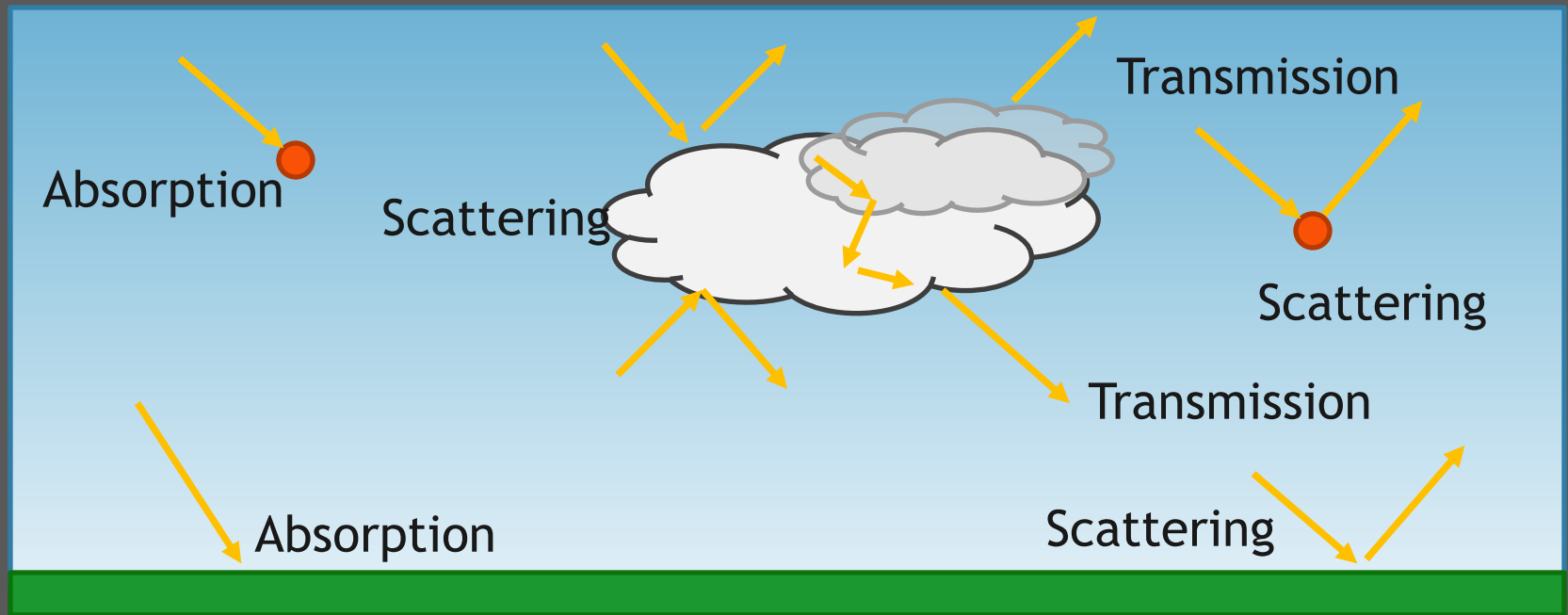
Introduction

- First atmospheric chemistry observations made from space over 30 years ago



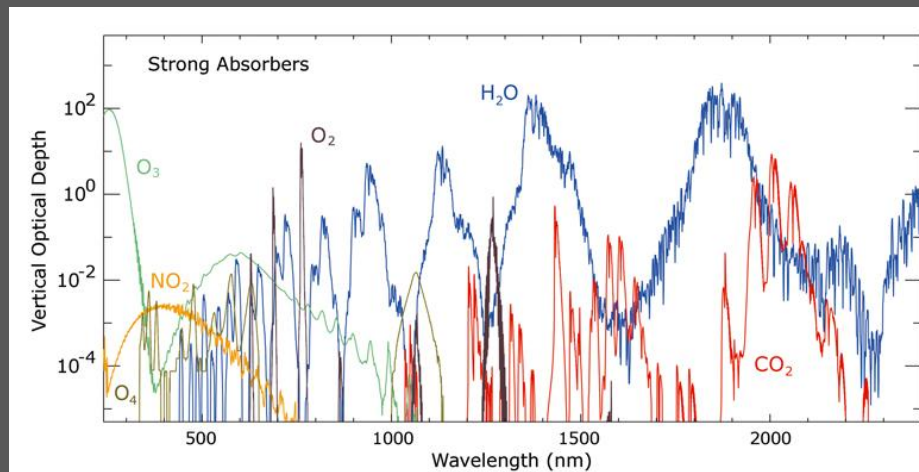
Introduction

- First atmospheric chemistry observations made from space over 30 years ago
- Solar radiation passing through the atmosphere can undergo many different processes
 - Absorption, Scattering, Reflection, Transmission



Measuring Atmospheric Gases from Space

- Absorbing gases can be identified and quantified from spectrally resolved measurements of the radiation
- From space this can be achieved by measuring the radiation that is reflected/scattered from sunlight in the atmosphere in the UV - visible - NIR - SWIR range
- By observing radiation at particular spectral wavelengths can focus on particular gases of interest



Detecting Atmospheric Gases - DOAS

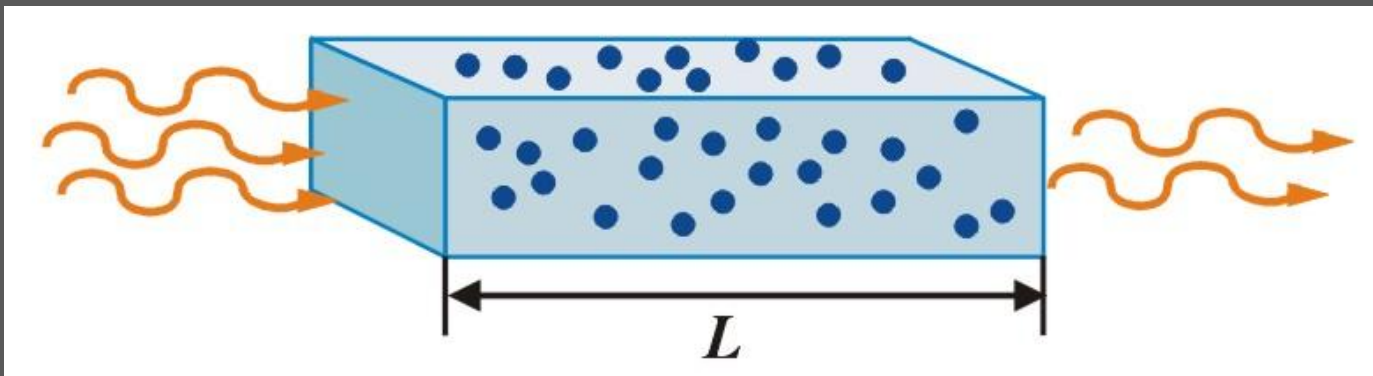
- DOAS - Differential Optical Absorption Spectroscopy
- Based on the measurement of absorption of radiation in the open atmosphere
- Many wavelengths are used simultaneously
- Each gas identified through unique absorption cross section
- Several absorbers can be retrieved at the same, some overlap of species

DOAS - In Detail

- Back to basics - the Beer Lambert law quantitatively describes absorption of radiation

$$I(\lambda) = I_0(\lambda) \exp[-L \sum \sigma_i(\lambda) n_i]$$

- L = path length / cm,
- $\sigma(\lambda)$ = absorption cross section, also dependent upon temperature and pressure / $\text{cm}^2 \text{ molecules}^{-1}$
- n = number density / molecules cm^{-3}



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 - $\sigma(\lambda)$ = absorption cross section, also dependent upon temperature and pressure / $\text{cm}^2 \text{ molecules}^{-1}$
 - n = number density / molecules cm^{-3}
- Optical depth $\tau(\lambda) = L\sigma(\lambda)n$, can be used to provide good estimation of concentration

$$\tau(\lambda) = \ln \left[\frac{I_0(\lambda)}{I(\lambda)} \right]$$

DOAS - Atmospheric Considerations

- Does not account for losses (extinction) processes other than absorption - **Scattering**
 - **Rayleigh scattering**: scattering by particles smaller than the wavelength of radiation
 - **Mie scattering**: scattering by particles similar or larger in size than the wavelength of radiation, e.g. aerosol particles
- Account for scattering by considering as pseudo-absorption process with their own cross sections
- Extinction coefficients

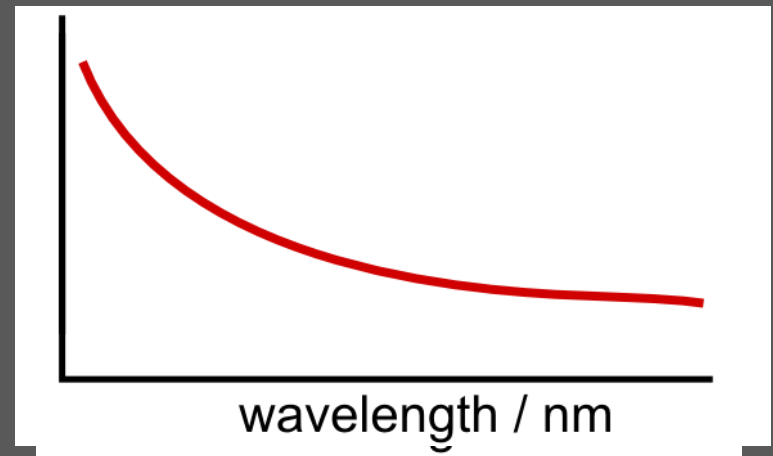
$$\varepsilon_R = \sigma_R(\lambda) n_{air}$$

$$\varepsilon_M = \sigma_M(\lambda) n_{air}$$

Differential Absorption

$$I(\lambda) = I_0(\lambda) \times \exp \left[-L \left(\sum_i (\sigma_i(\lambda) n_i) + \varepsilon_R(\lambda) + \varepsilon_M(\lambda) \right) \right]$$

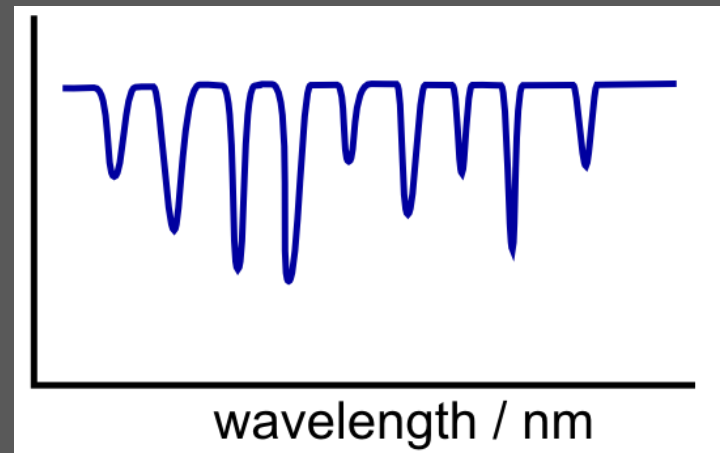
- $I_0(\lambda)$ is difficult to measure in the atmosphere - instead measure the **differential absorption**
- Defined as the part of the total absorption of any molecule that rapidly varies with the wavelength
- Processes such as scattering produce a ‘slow’ variation in the radiance with wavelength



Differential Absorption

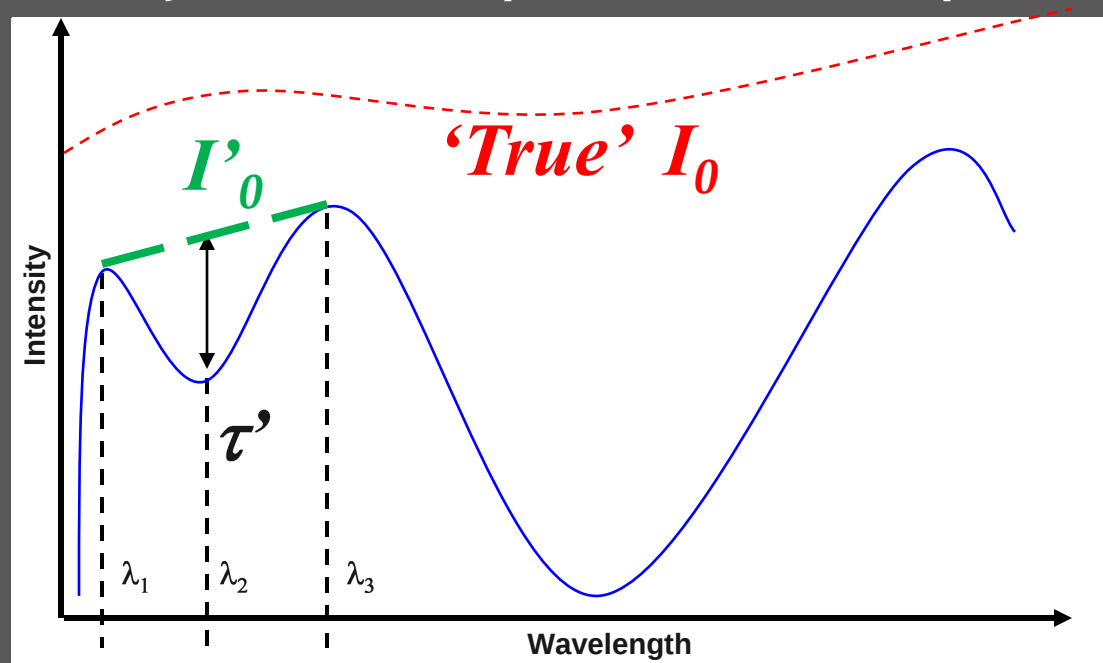
$$I(\lambda) = I_0(\lambda) \times \exp \left[-L \left(\sum_i (\sigma_i(\lambda) n_i) + \varepsilon_R(\lambda) + \varepsilon_M(\lambda) \right) \right]$$

- $I_0(\lambda)$ is difficult to measure in the atmosphere - instead measure the **differential absorption**
- Defined as the part of the total absorption of any molecule that rapidly varies with the wavelength
- Molecular absorption features sharply defined - 'rapid' variation with wavelength



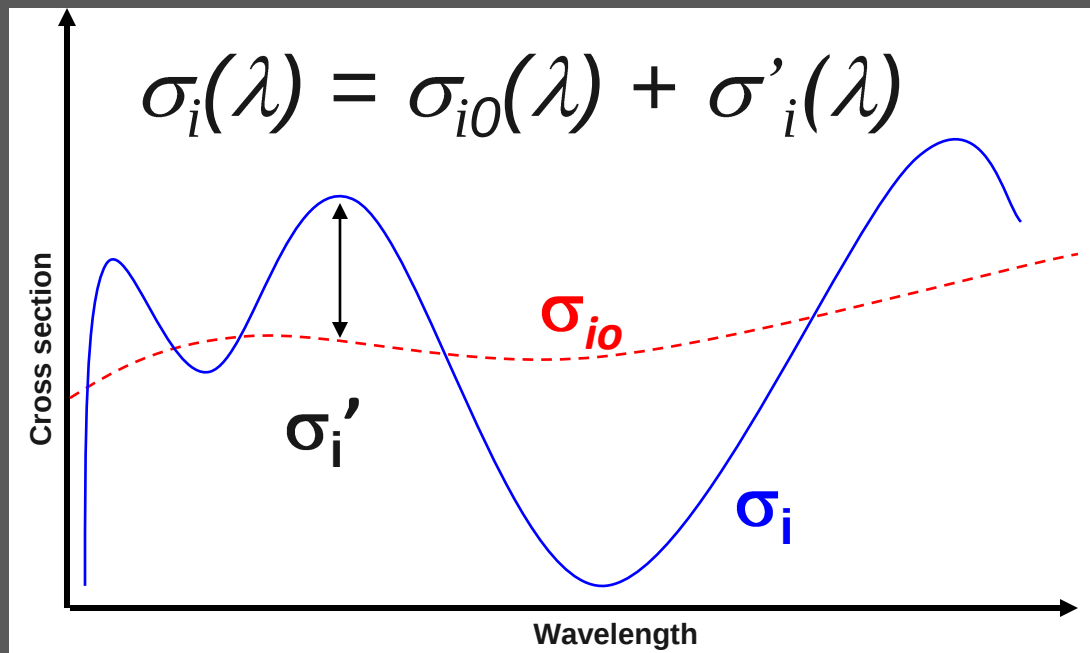
Differential Absorption

- To accurately pick out these features in atmospheric spectra requires a high sampling wavelength interval and high resolution
- Using this system the absorption cross section and intensity will be split into two portions



Differential Absorption

- To accurately pick out these features in atmospheric spectra requires a high sampling wavelength interval and high resolution
- Using this system the absorption cross section and intensity will be split into two portions



Differential Absorption

- Equation becomes:

$$I(\lambda) = I_0(\lambda) \times \exp \left[-L \left(\sum_i \sigma'_i(\lambda) n_i \right) \right] \times \exp \left[-L \left(\sum_i \sigma_{i0}(\lambda) n_i + \varepsilon_R(\lambda) + \varepsilon_M(\lambda) \right) \right] \times A(\lambda)$$

Rapid

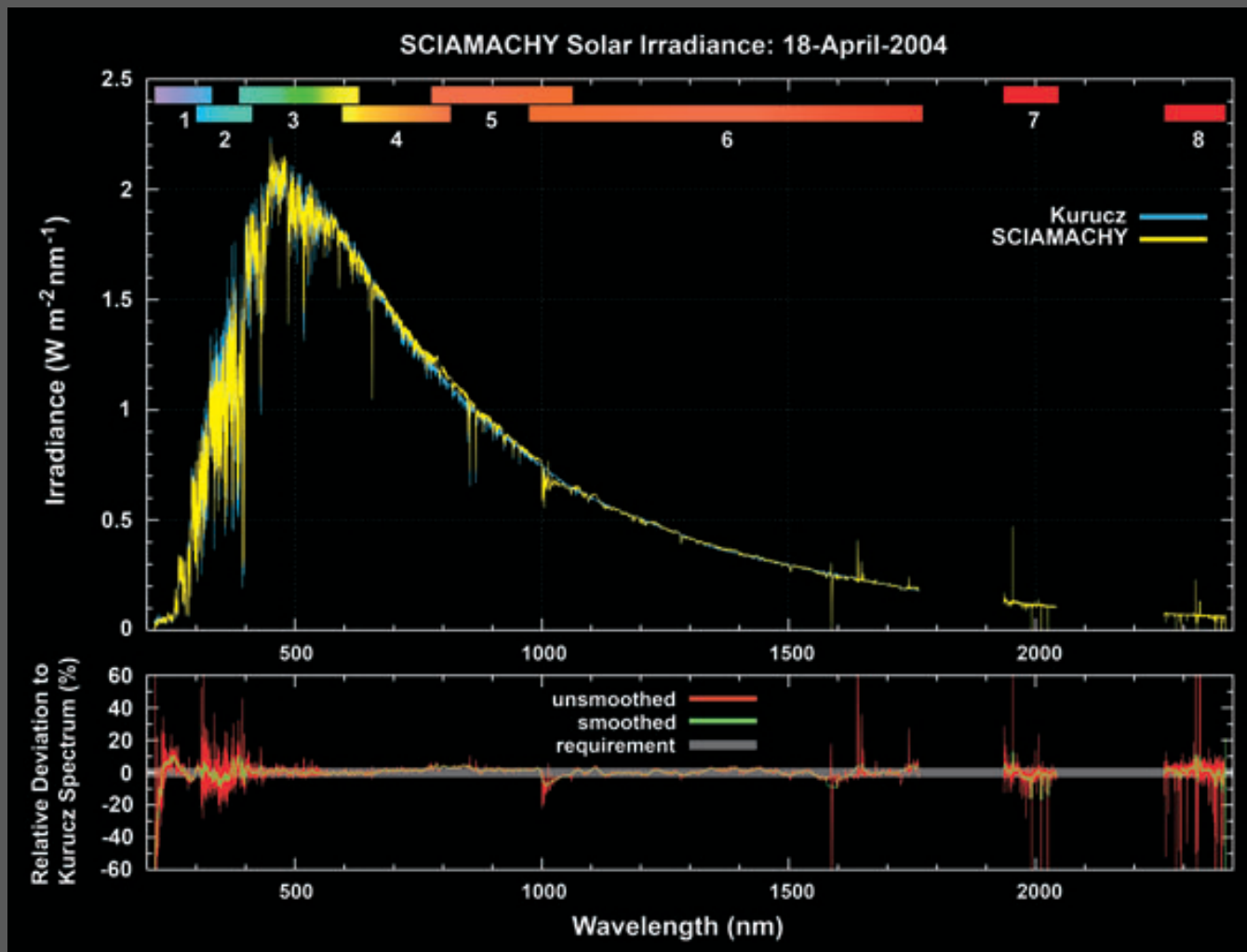
Slow

- where

$$I'_0(\lambda) = I_0(\lambda) \times \exp \left[-L \left(\sum_i \sigma_{i0}(\lambda) n_i + \varepsilon_R(\lambda) + \varepsilon_M(\lambda) \right) \right] \times A(\lambda)$$

- Can determine $I'_0(\lambda)$ from atmospheric spectrum by removing the absorption features, either through using a polynomial fit, digital smoothing or a Fourier transform of the measured intensity $I(\lambda)$

Solar Spectrum



Determining Concentration

- The optical depth can then be determined

$$\tau'(\lambda) = \ln \left(\frac{I'_0(\lambda)}{I(\lambda)} \right) = L \sum_i \sigma'_i(\lambda) n_i$$

- The amount of absorbing gas molecules can be estimated directly from the optical depth

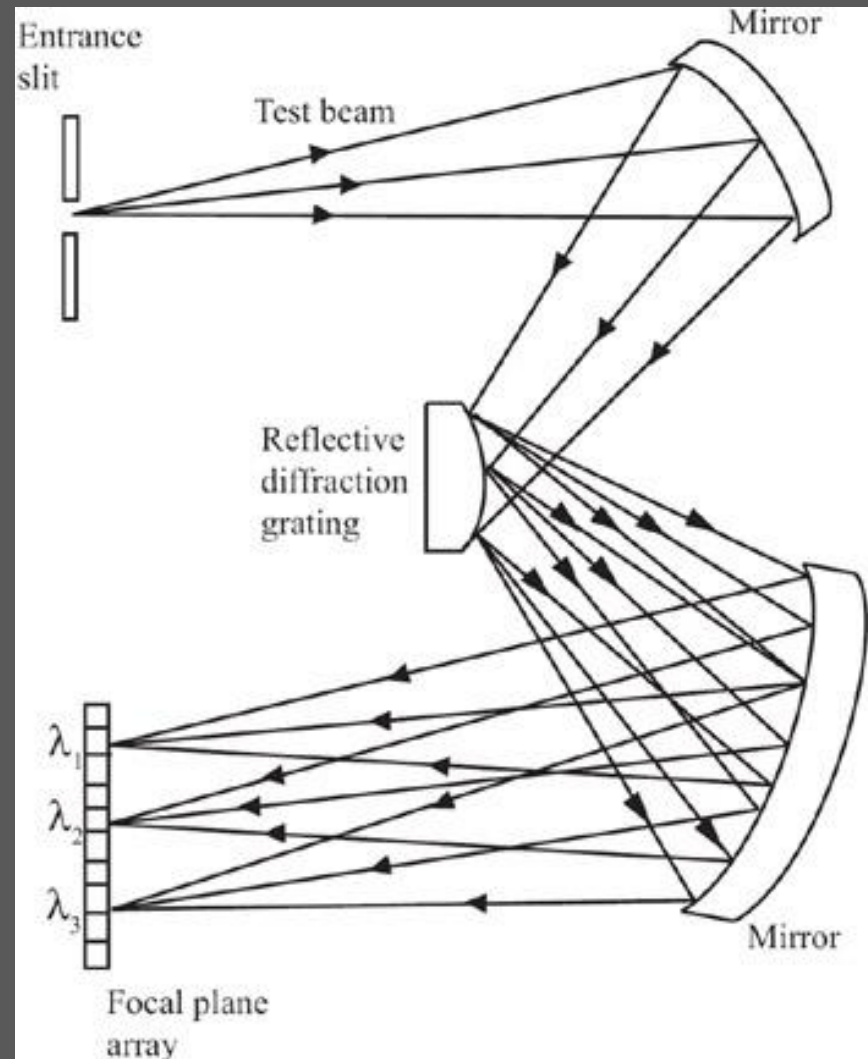
$$\tau'_i(\lambda) \approx \sigma'_i(\bar{T}, \bar{p}) \times SCD$$

- SCD is the slant column density - the number of molecules per area along the path from the Sun to the instrument

$$SCD = \int_0^{\lambda} n(\lambda) d\lambda \quad [\text{molecules cm}^{-2}]$$

DOAS Instrumentation - grating spectrometer

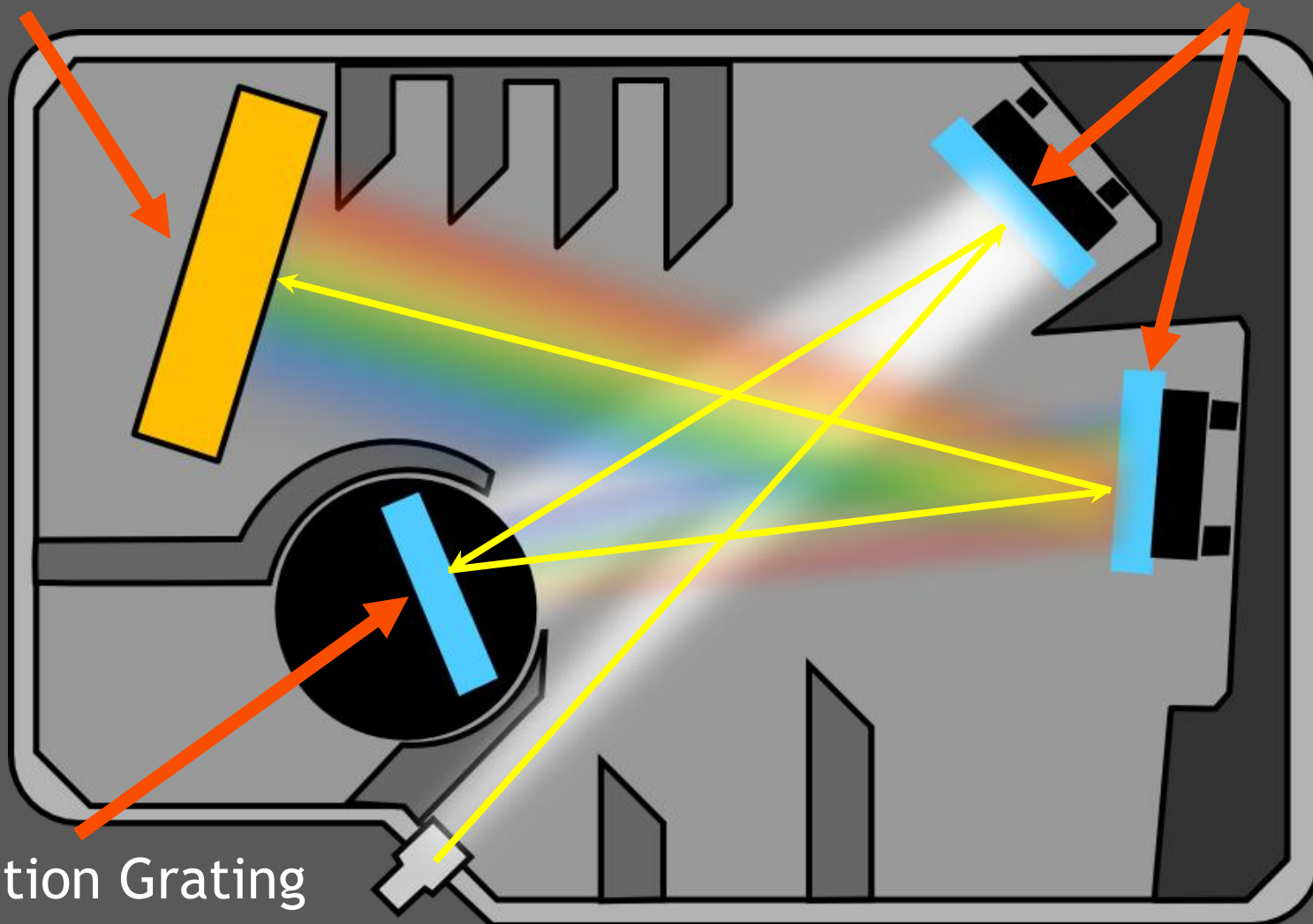
- Grating Spectrometer - standard Offner array



DOAS Instrumentation - grating spectrometer

CCD - detector

Focusing mirror

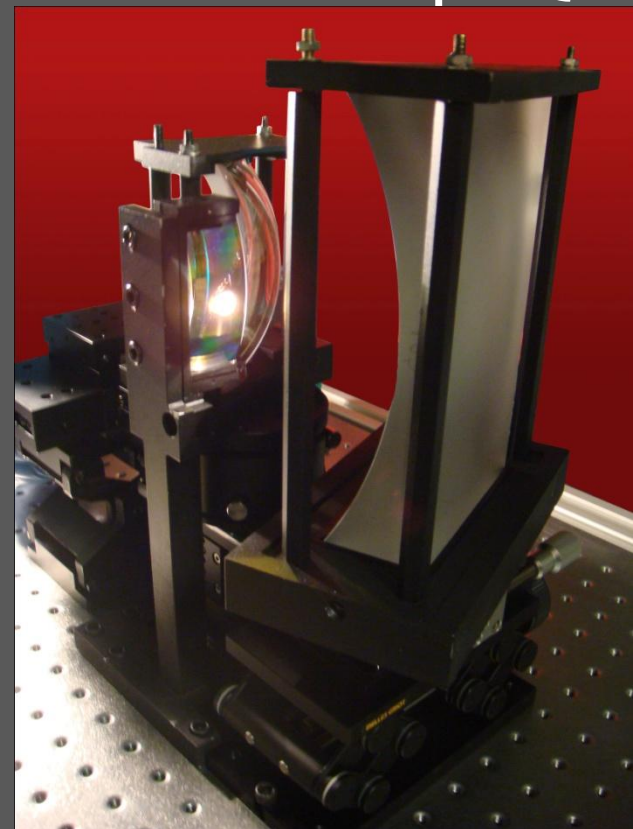
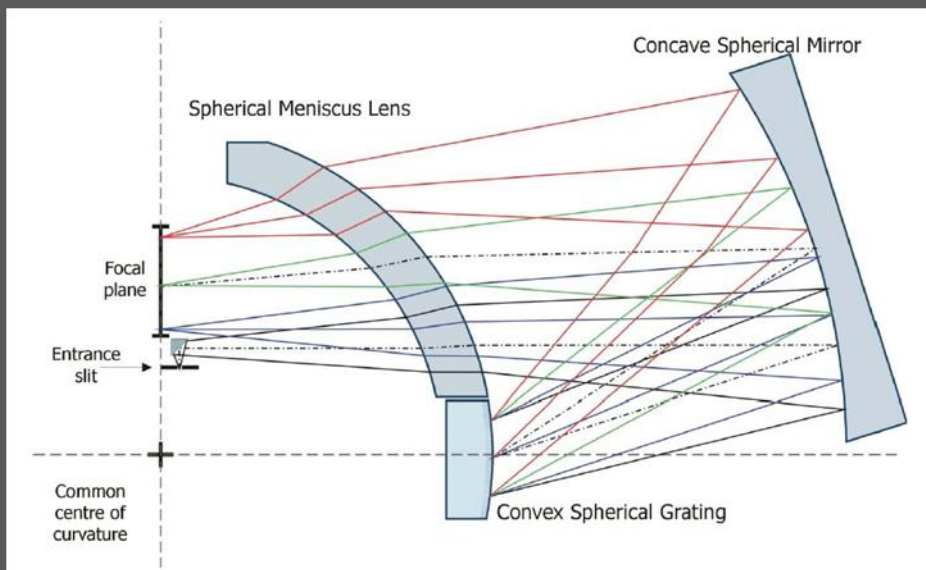


Diffraction Grating

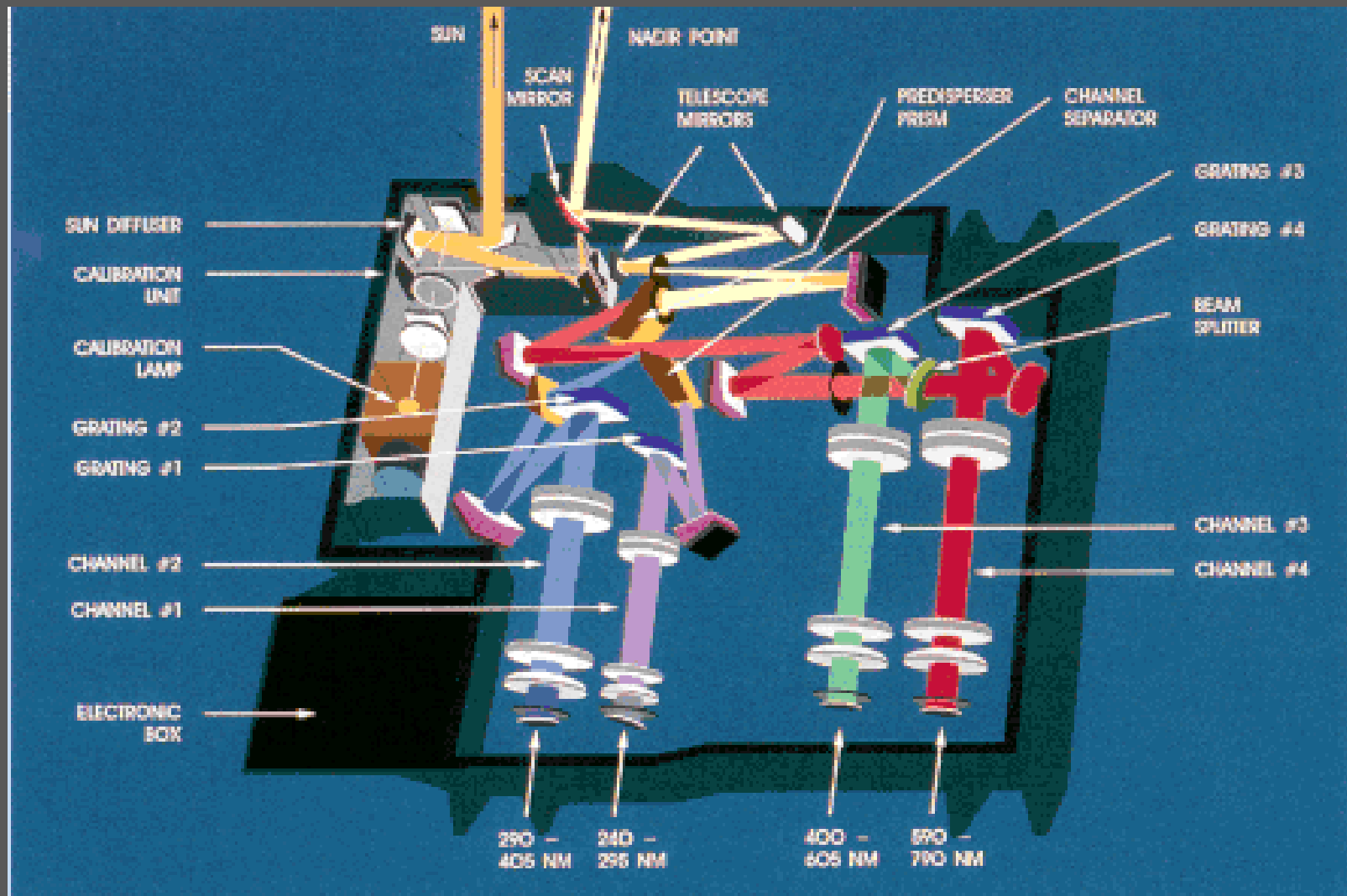
Advanced Instrumentation

- Compact mass spectrometers - lighter, less optical components
- Example: concentric spectrometer: CompAQS
- C Whyte *et al*

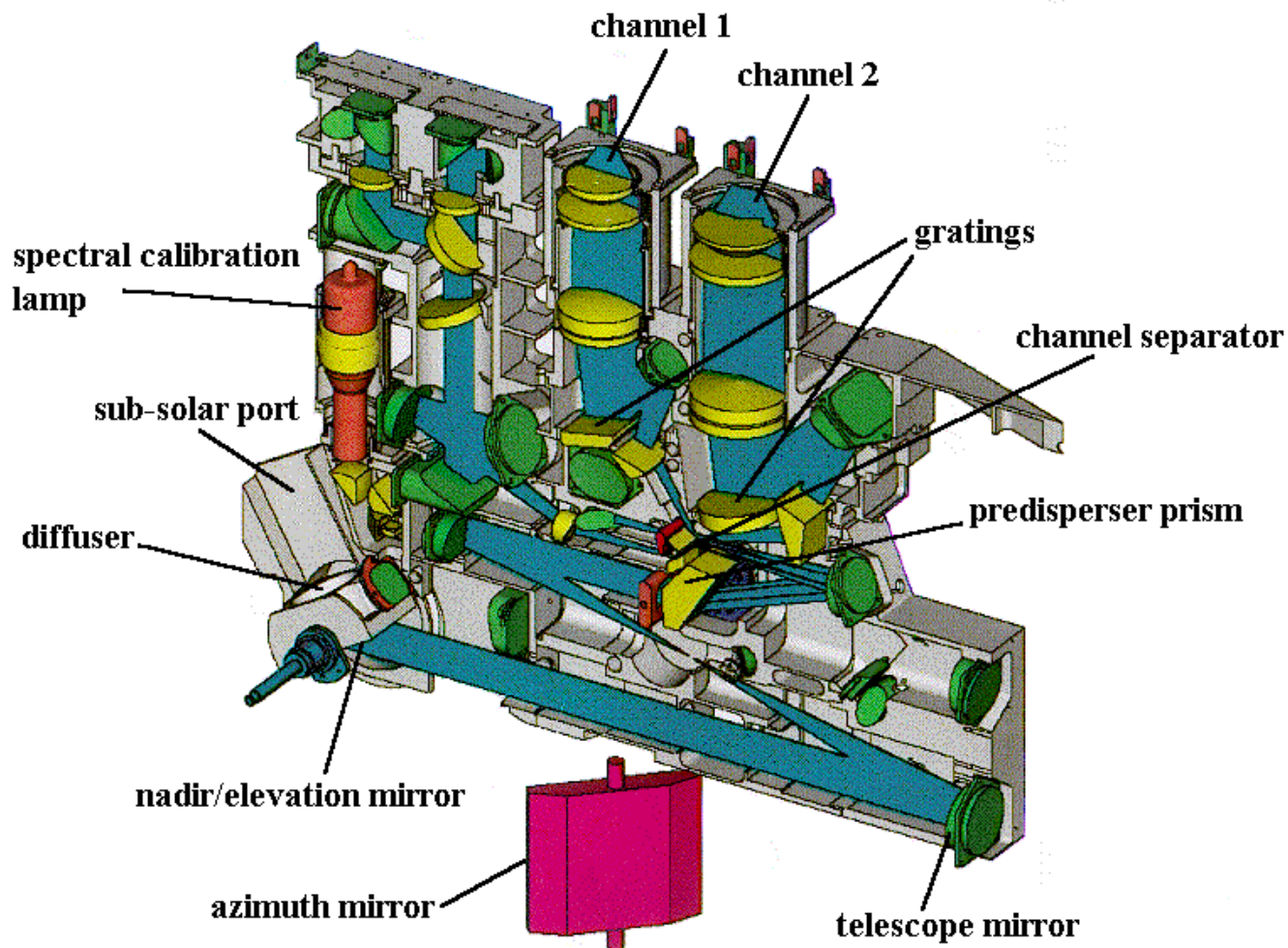
Atmos. Meas. Tech., 2, 789-800, 2009



GOME instrument on ERS-2

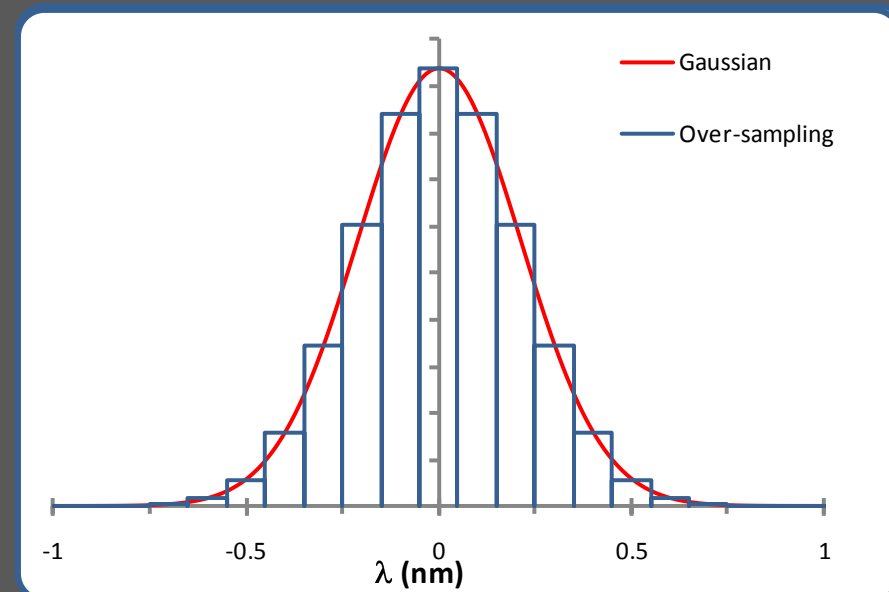


SCIAMACHY on Envisat



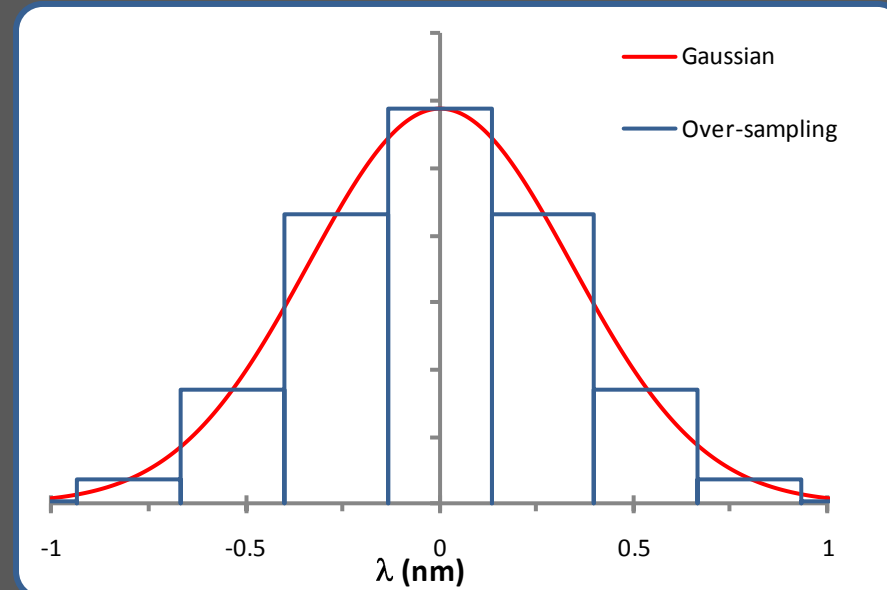
Instrument Line Shape

- Instrument feature - shape of a singly resolved spectral element made up of a varying number of measurements
- FWHM of shape - **resolution**
- Number of measurements defining one resolved element - **oversampling**
- FWHM = 0.5 nm
- 5 pixel oversampling
- Sampling interval = 0.1nm

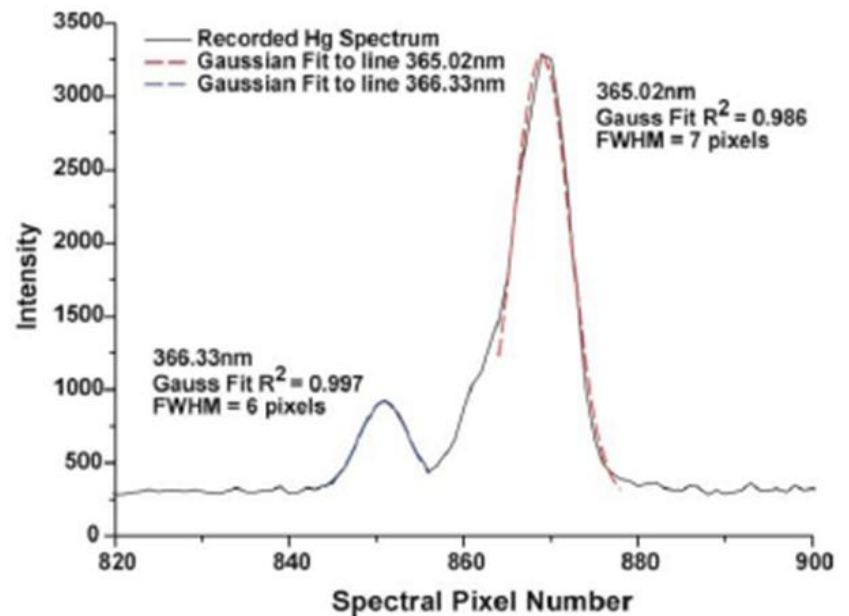
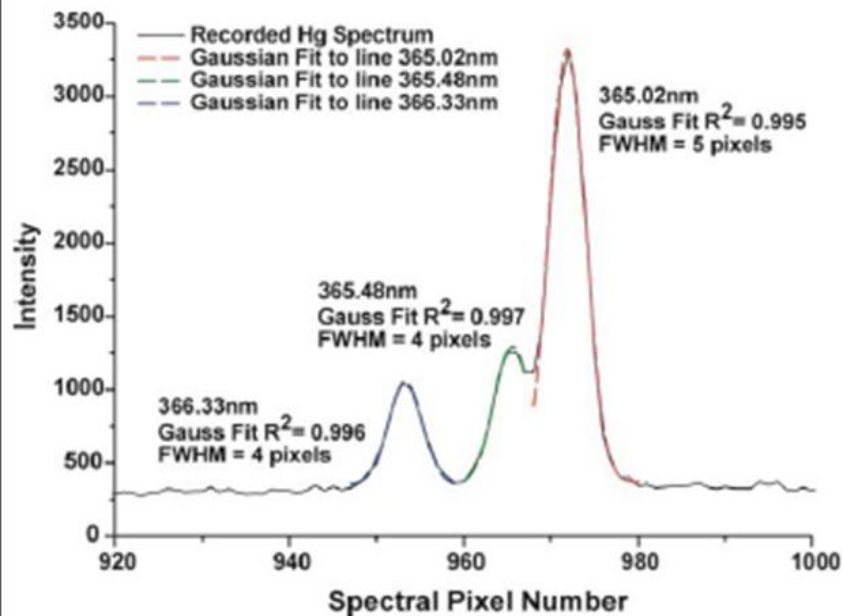


Instrument Line Shape

- Instrument feature - shape of a singly resolved spectral element made up of a varying number of measurements
- FWHM of shape - **resolution**
- Number of measurements defining one resolved element - **oversampling**
- FWHM = 0.8 nm
- 3 pixel oversampling
- Sampling interval = 0.26nm



CompAQS example spectrum



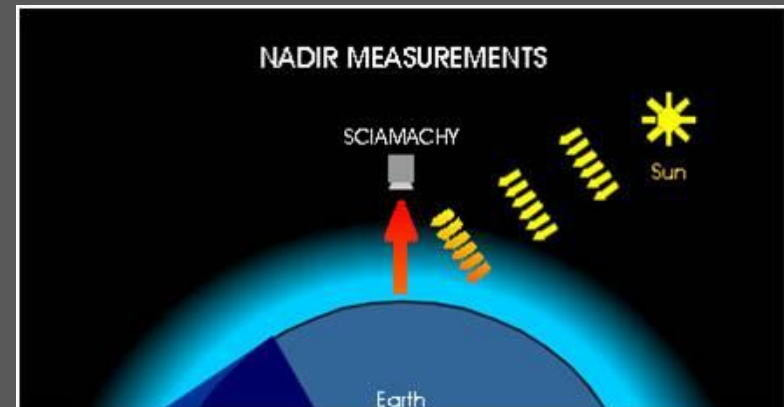
Spatial Resolution

- High spatial resolution desirable
- A smaller spatial resolution will increase the number of cloud free pixels - scenes not obscured by cloud
- Approach urban scale - assessment of local air quality and monitor transport across urban centres
- Practical application to the real world, e.g. AirText service provides local forecasting for AQ in London

Satellite Viewing Geometries

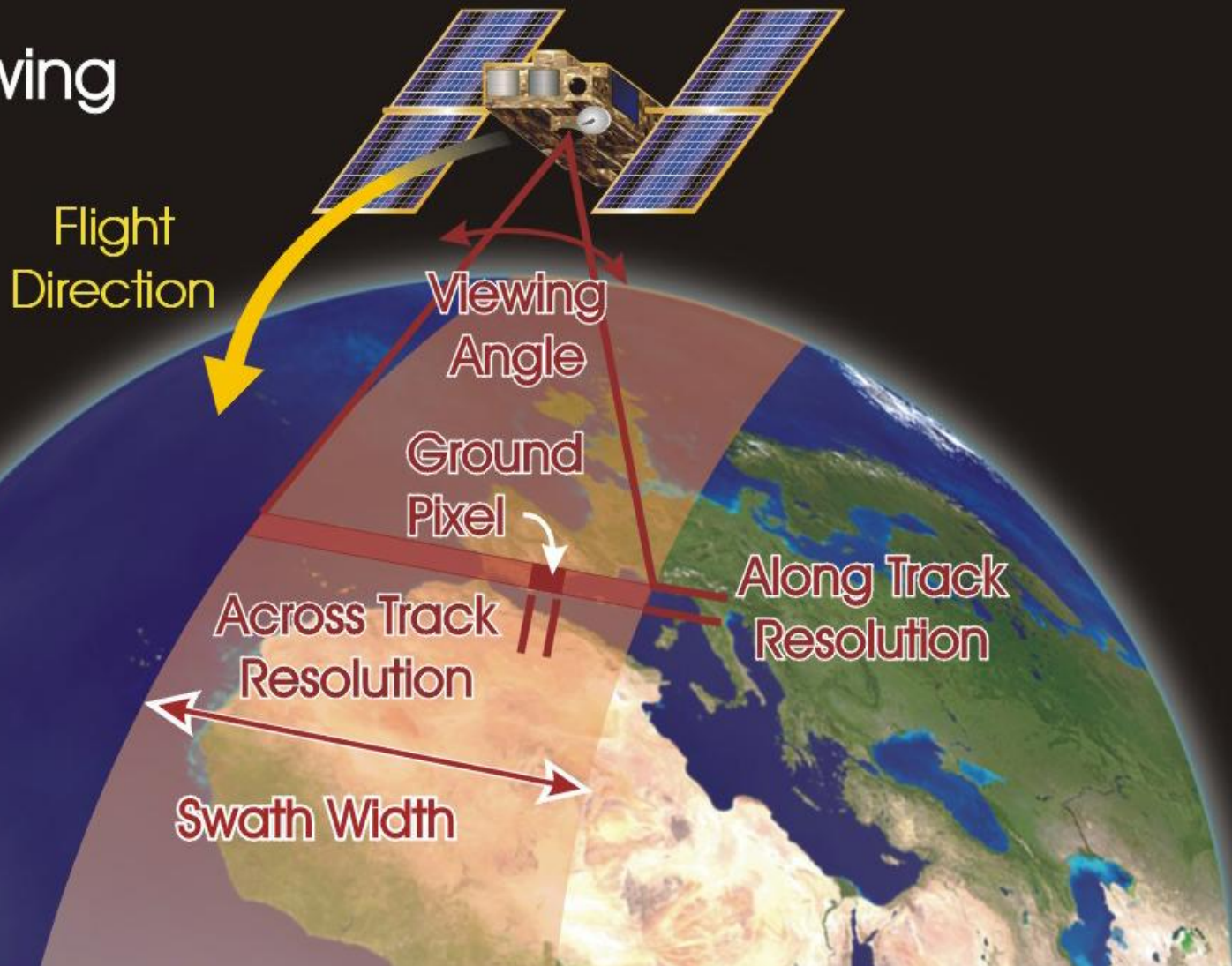
- Low Earth Orbit

- A sun-synchronous orbit crosses the equator at the same local time each day
- Allows consistent scientific observations as the angle of sunlight on the Earth's surface remains relatively constant - seasonal variation
- Nadir observes either emission or backscattered radiation
- Measure total column of absorbing gases
- Global coverage



Satellite Viewing Geometries

Nadir Viewing



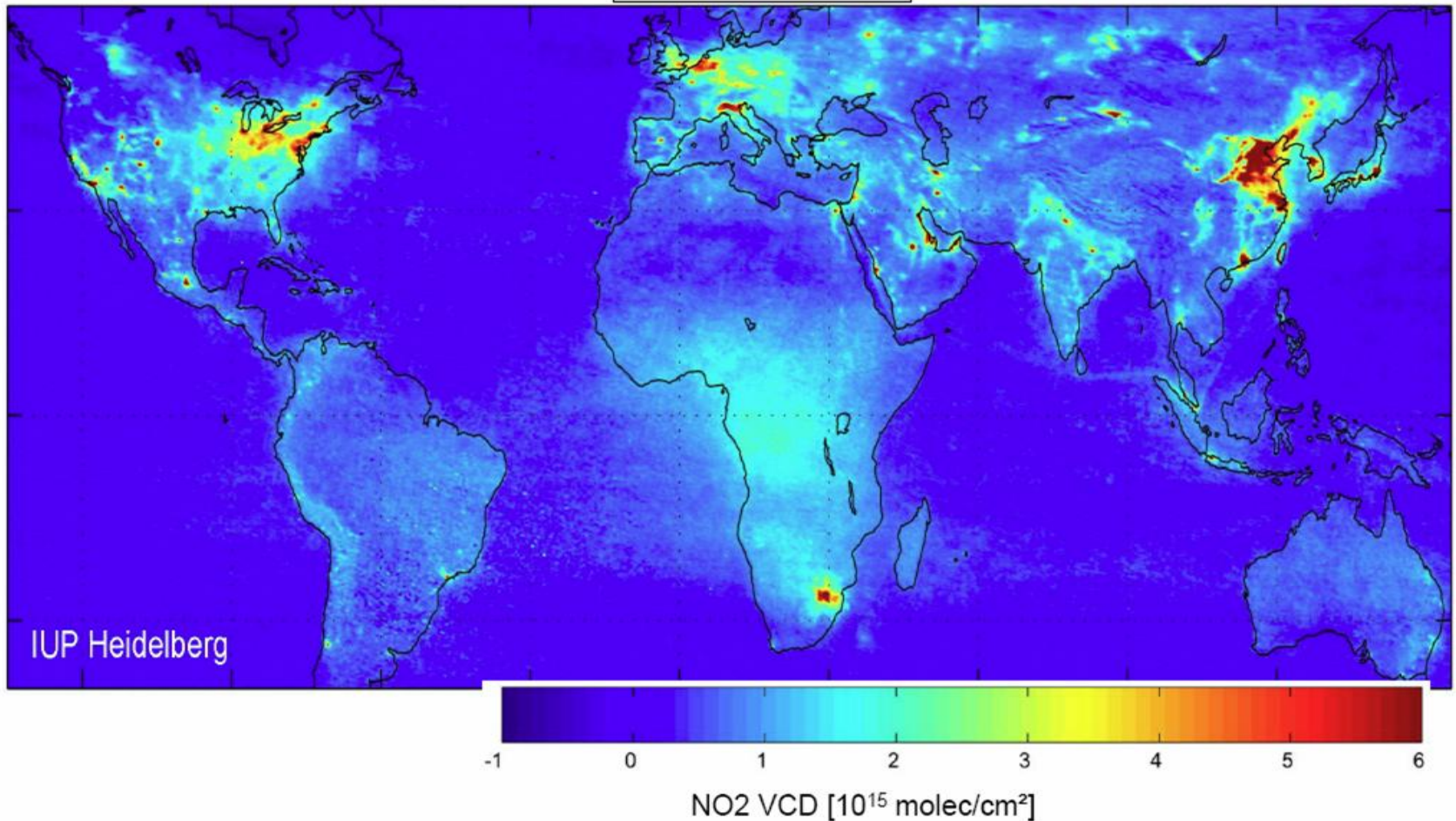
Atmospheric Gases: Nitrogen Dioxide

- Major air pollutant originating mostly from the combustion of fossil fuels (vehicle emissions)
- Source of ground level ozone, plays vital role in formation of tropospheric ozone and aerosols
- Potent lung irritant with known health implications
- Tropospheric (near-surface) amount of NO_2 can be measured by UV/Vis satellite instrument due to their sensitivity to the whole atmospheric column

Global Measurement of NO₂

NO₂ (S. Beirle, IUP Heidelberg)

SCIAMACHY, 2003/04



Nitrogen Dioxide over Europe

OMI NRT tropospheric NO₂ 11 Mar 2010

KNMI/NASA/NIVR

